



High-frequency QPOs from Black Hole Binaries: the RossiXTE Legacy

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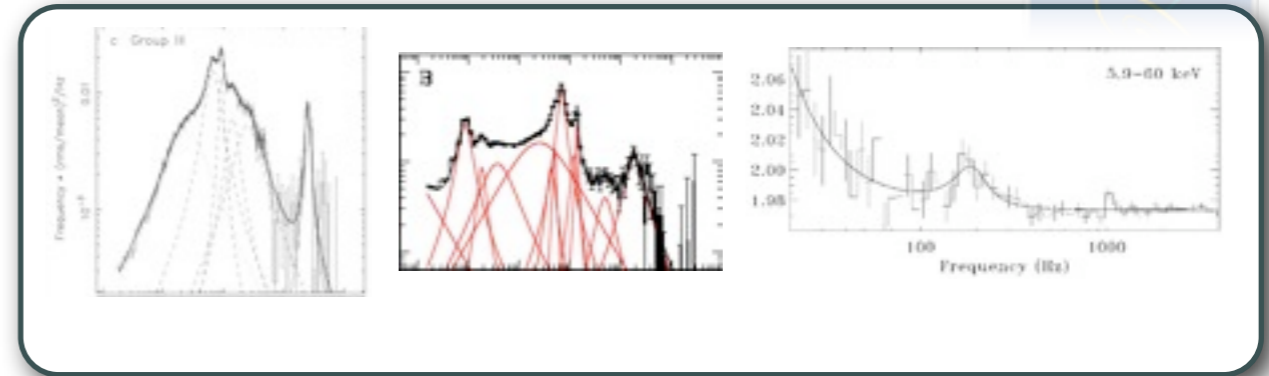
Leverhulme Visiting Professor, University of Southampton

with Andrea Sanna & Mariano Méndez (Univ. Groningen)

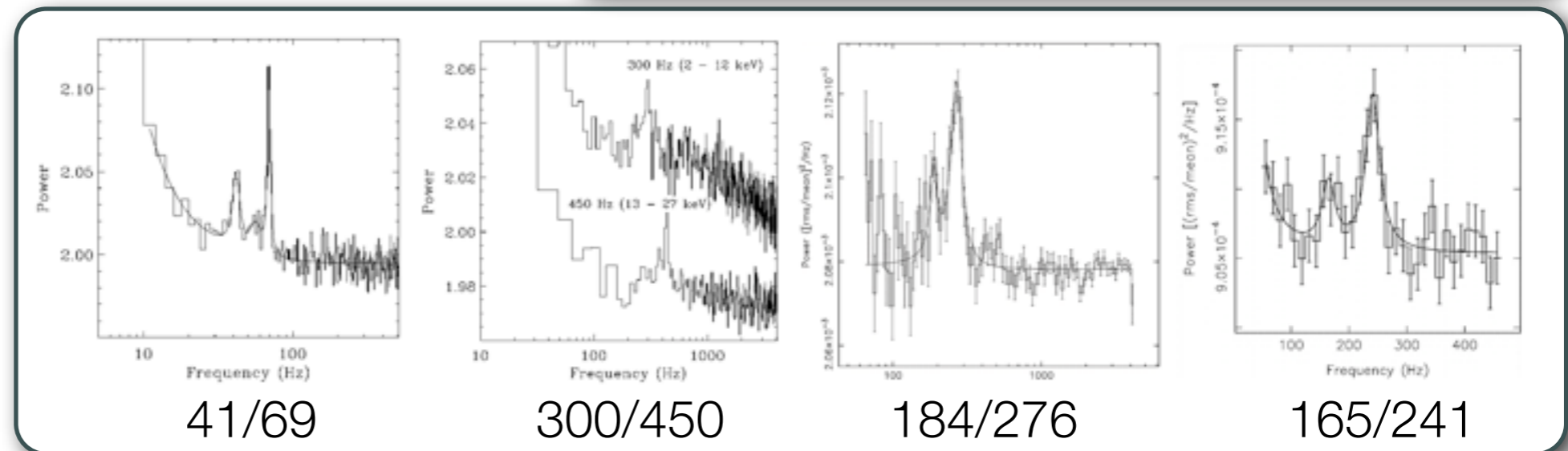


High-Frequency oscillations in BHT

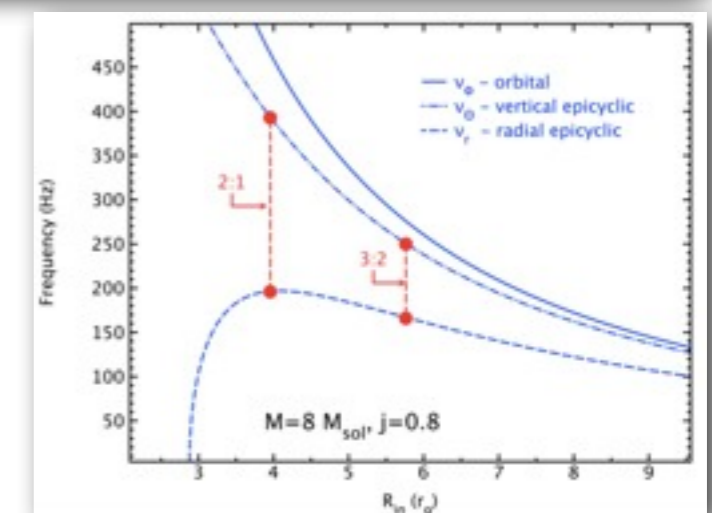
- 30-450 Hz
- Very few detections



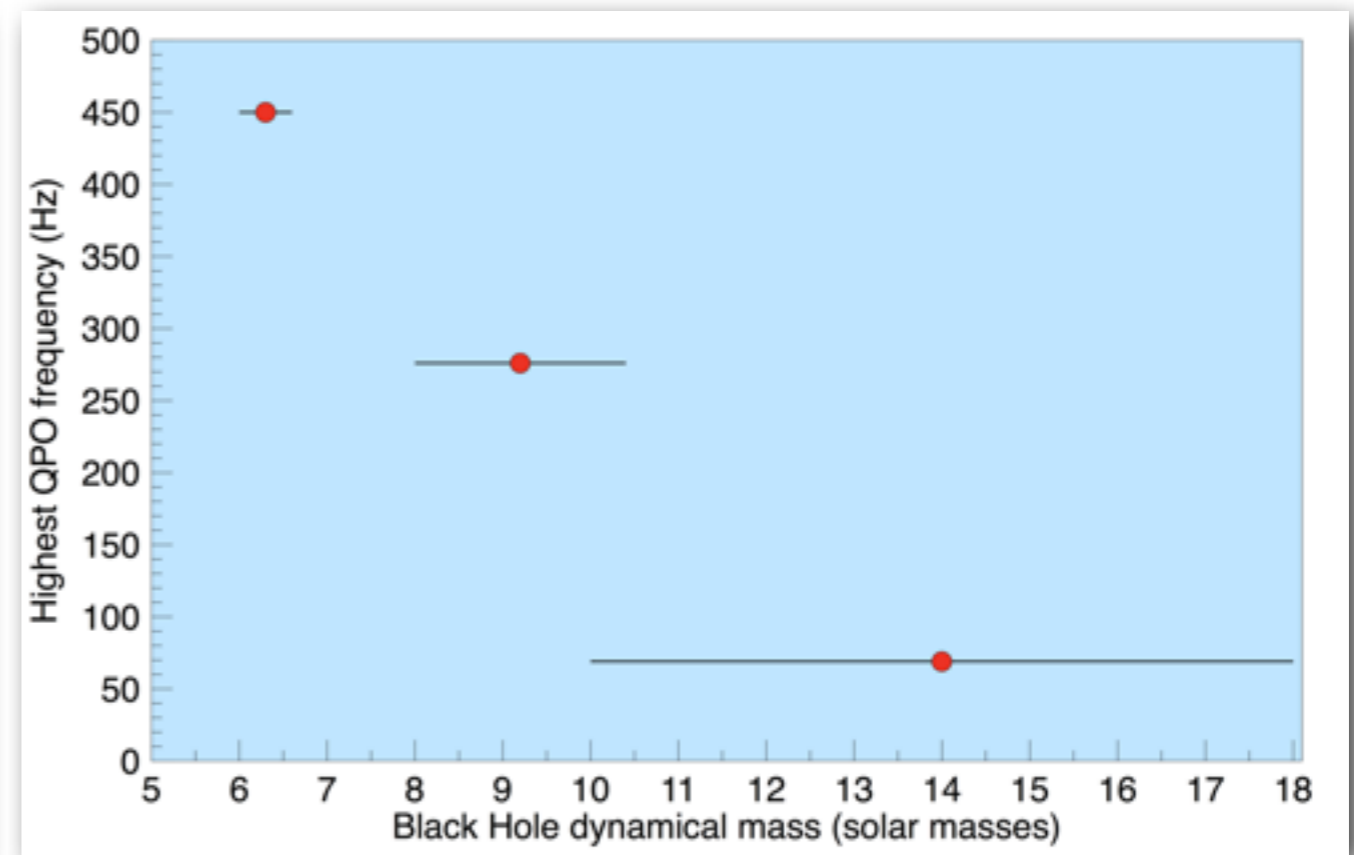
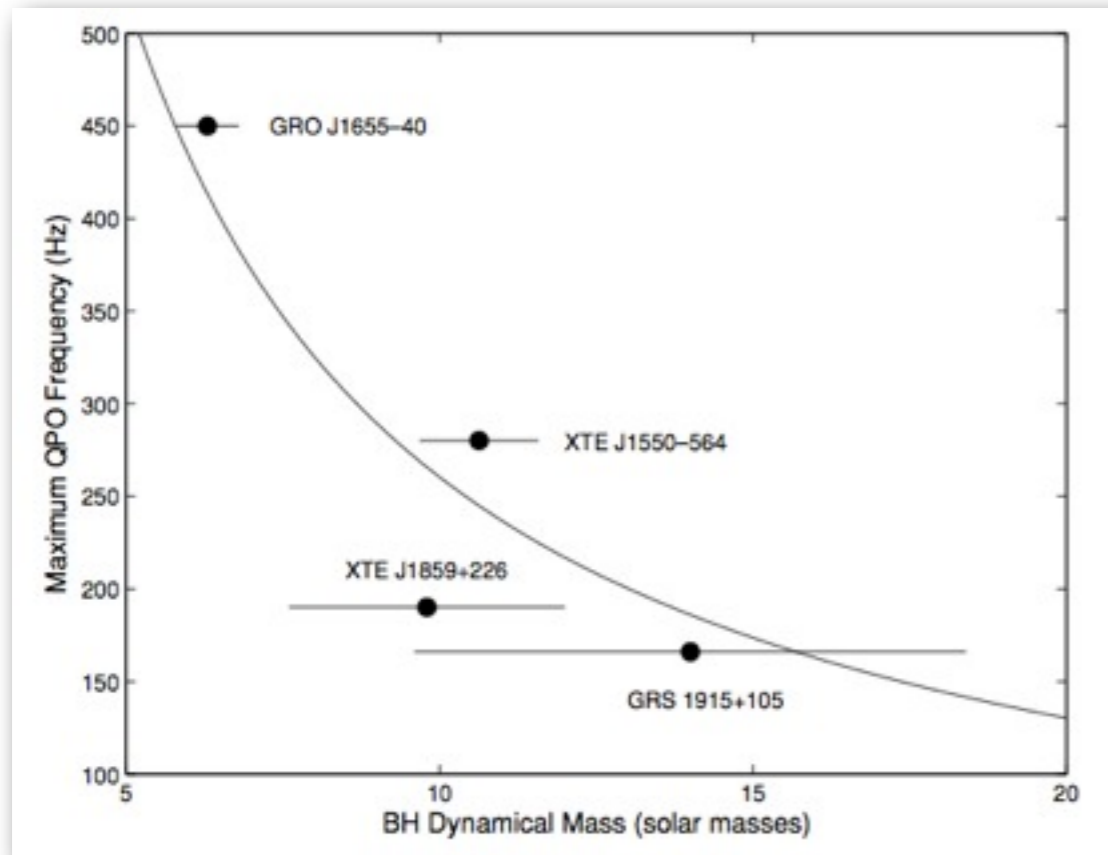
- Ratios?
- BH mass?



- Resonance model
- Too few points... some of them not even real

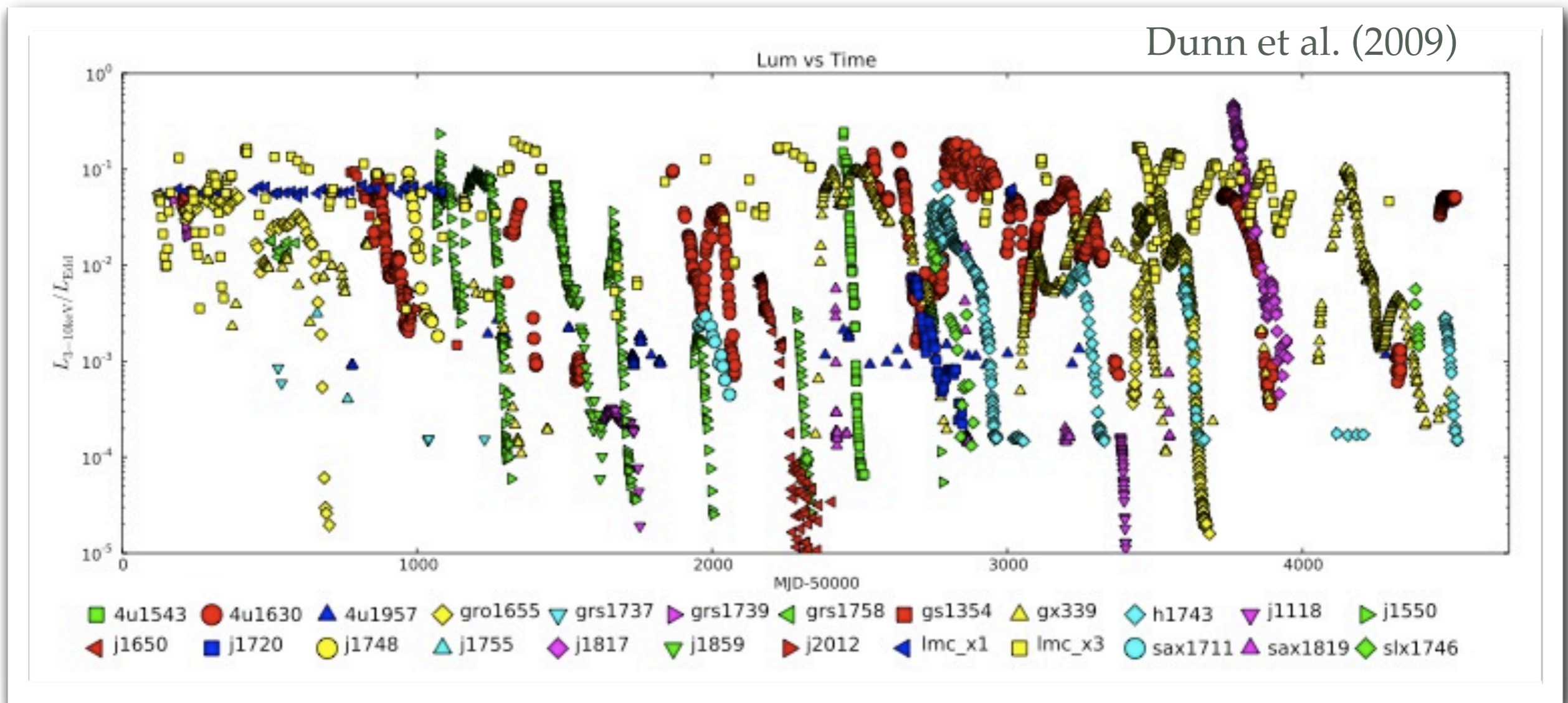
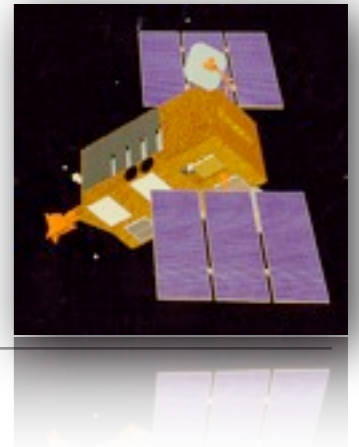


Mass correlation?



Last RXTE one from IGR J17091-3624 (Altamirano's talk)

RXTE observed many BHT many times



Need for a systematic analysis



Source name	N. obs.	Valid T.	Valid H.	Exp. T (ks)	Exp. H (ks)	$> 3\sigma$ T	$> 3\sigma$ H	Good T	Good H	Final
GX 339-4	1353	893	517	2190	1657	31	34	4	11	4
4U 1630-47	1002	846	701	1934	1672	31	27	11	7	4
GRO J1655-40	596	516	491	2295	2214	18	24	7	8	11
H 1743-322	504	386	300	1153	941	19	20	6	8	4
XTE J1550-564	409	319	280	860	790	46	31	15	14	16
Swift J1753.5-0127	278	240	172	671	527	14	8	3	3	0
XTE J1752-223	208	138	79	343	250	6	6	0	4	2
XTE J1650-500	182	101	69	219	155	17	5	4	2	0
Swift J1539.2-6227	156	61	42	153	122	1	3	1	1	0
XTE J1817-330	155	112	79	372	256	3	2	1	1	0
XTE J1859+226	131	115	102	314	281	6	4	2	2	1
4U 1543-47	104	53	38	161	125	6	3	0	1	0
XTE J1720-318	100	75	19	233	101	1	1	0	1	0
4U 1957+115	100	97	46	529	334	1	0	0	0	0
XTE J1118+480	94	62	50	174	154	0	3	0	1	0
MAXI J1659-152	66	63	54	148	135	1	1	0	1	0
SLX J1746-331	65	47	23	125	72	2	2	0	0	0
XTE J1652-453	57	30	1	67	2	1	0	0	0	0
XTE J1748-288	24	23	21	103	99	0	1	0	0	0
SAX J1711.6-3808	17	17	12	43	30	0	0	0	0	0
GS 1354-644	8	7	7	52	52	0	1	0	0	0
GRS 1737-31	5	4	3	40	28	0	0	0	0	0
Total	7108	4205	3106	12177	9996	204	176	54	65	42

Belloni, Sanna & Méndez (2012)

Strict procedure



- Exclusion of hopeless observations
 - Automatic detection procedure
 - Look for one significant peak ($> 3\sigma$ 1 trial)
 - Check all by hand
 - Re-fit manually
-
- Found 42 detections ($> 3\sigma$ 1 trial)
 - 7 sources (no GRS 1915+105 & XTE J17091)



Number of trials

- Very important issue
- For each source:
 - Number of soft+hard observations
 - Number of independent frequencies

- Select only detections with $P < 1\%$

- Left with 11 detections
- No double peaks



The final detections

- Very important issue
- For each source:
 - Number of soft+hard observations
 - Number of independent frequencies

- Select only detections with $P < 1\%$

- Left with 11 detections
- Two sources: XTE J1550-564 & GRO J1655-40

Belloni, Sanna & Méndez (2012)

The final detections

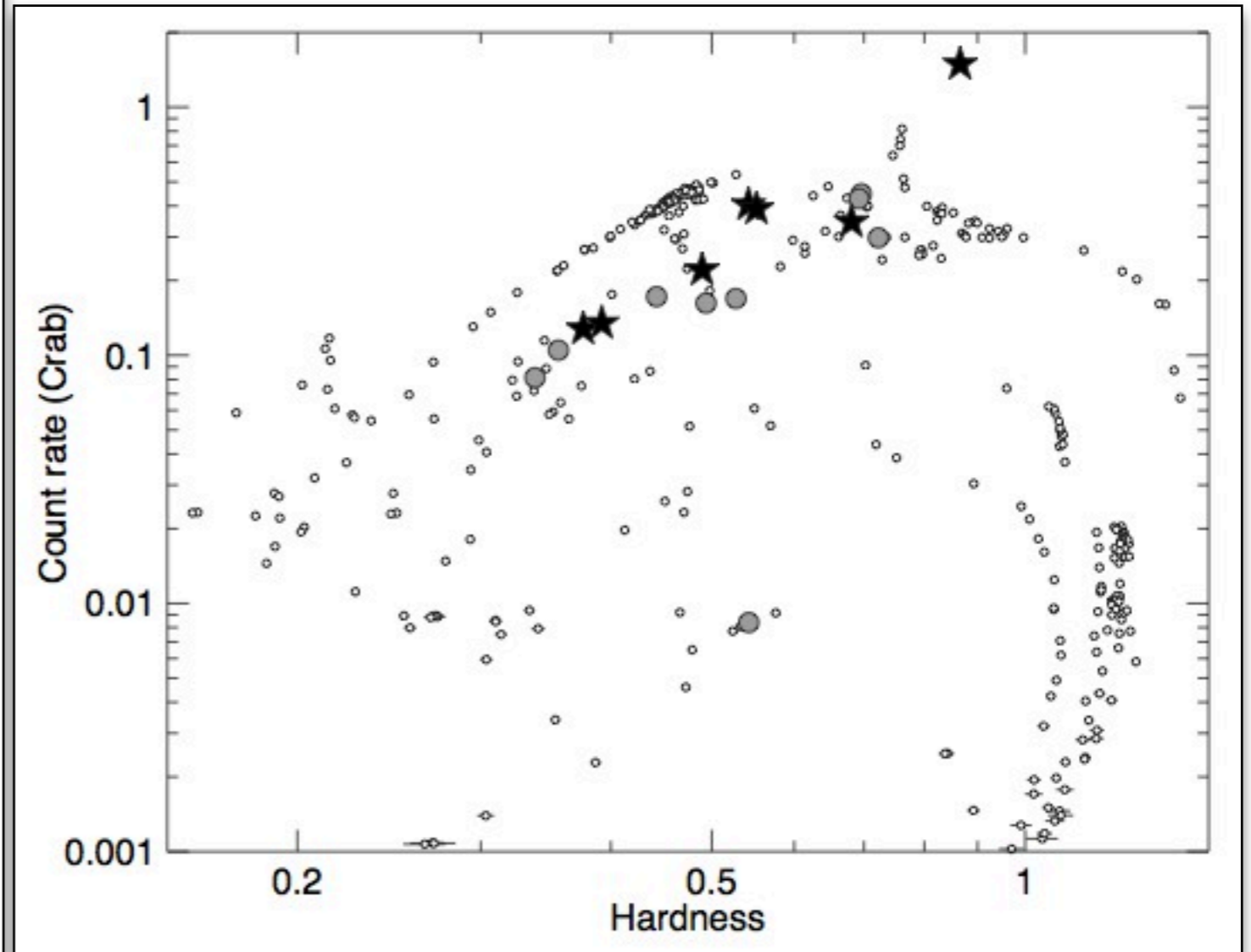
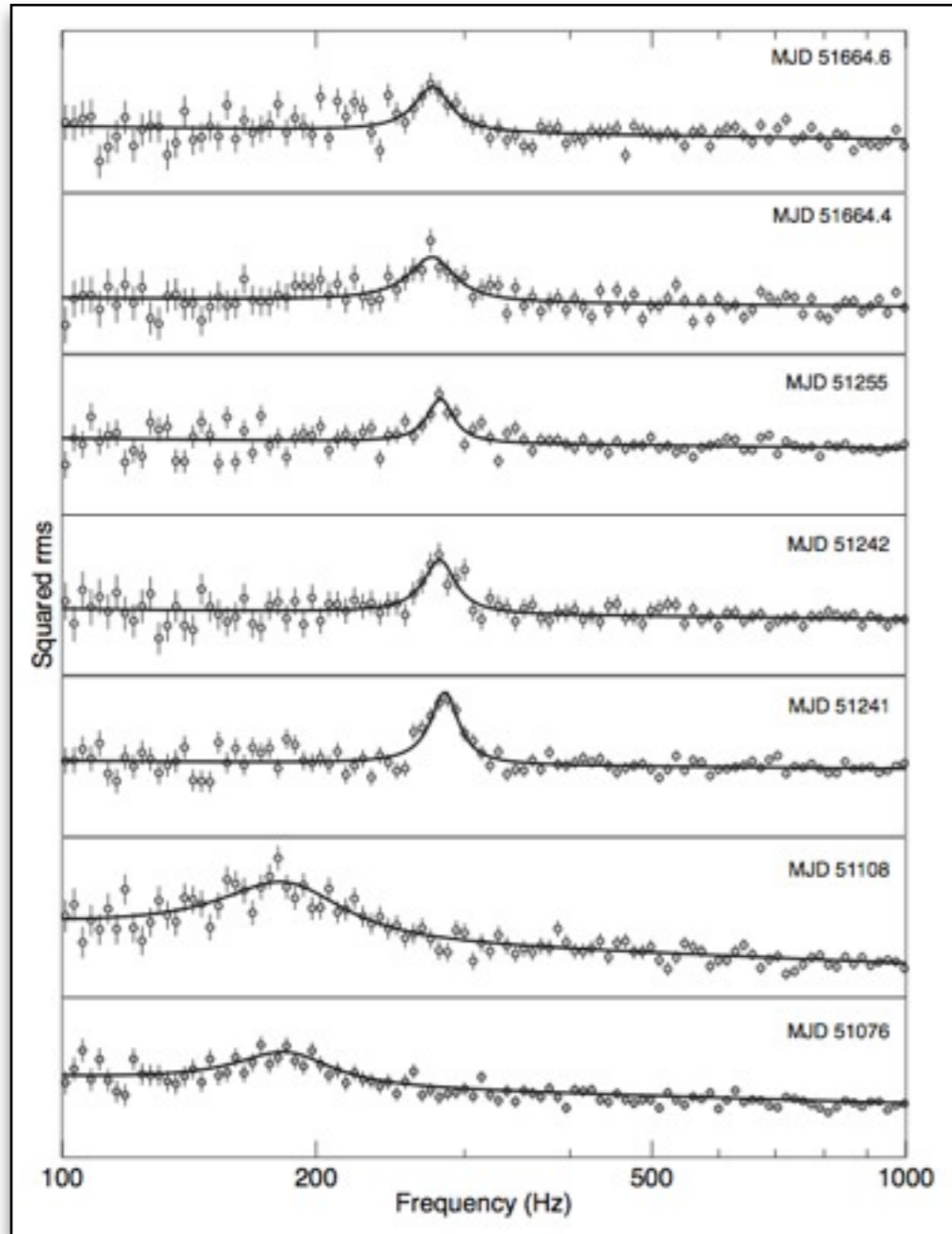
Belloni, Sanna & Méndez (2012)



MJD	Exp. (s)	B	ν_0 (Hz)	FWHM (Hz)	%rms	P_0	P_F	Rate	HR
XTE J1550-564									
51076.000	2960	T	180.76 ^{+3.92} _{-4.37}	95.12 ^{+16.12} _{-22.00}	1.15 ^{+0.14} _{-0.12}	2.87E-07	1.62E-03	15700	0.68
51101.607	1584	T	140.70 ^{+1.71} _{-1.92}	16.51 ^{+19.85} _{-4.84}	1.12 ^{+0.59} _{-0.17}	6.64E-04	1.00E+00	3135	0.57
51108.076	9872	T	182.27 ^{+3.25} _{-3.20}	77.88 ^{+11.65} _{-9.77}	1.74 ^{+0.20} _{-0.16}	1.90E-08	1.31E-04	3636	0.53
51115.281	2048	H	272.43 ^{+3.42} _{-4.09}	30.59 ^{+16.30} _{-8.95}	4.02 ^{+0.78} _{-0.47}	7.80E-06	1.28E-01	1785	0.41
51241.802	3104	H	284.07 ^{+1.35} _{-1.36}	26.86 ^{+3.88} _{-3.12}	3.66 ^{+0.21} _{-0.17}	8.77E-29	0.00E+00	4258	0.42
51242.507	1504	H	280.00 ^{+2.87} _{-2.49}	30.00 ^{+9.80} _{-6.85}	3.78 ^{+0.46} _{-0.32}	2.05E-09	3.69E-05	4107	0.43
51245.354	2752	T	180.75 ^{+2.02} _{-2.54}	16.76 ^{+8.71} _{-6.21}	1.08 ^{+0.25} _{-0.13}	8.54E-06	2.40E-01	4718	0.54
51247.979	2816	T	185.20 ^{+6.73} _{-4.93}	46.04 ^{+15.69} _{-12.13}	1.32 ^{+0.30} _{-0.17}	3.30E-05	3.21E-01	4509	0.54
51255.158	4080	H	280.84 ^{+2.00} _{-1.92}	24.37 ^{+8.53} _{-6.28}	4.35 ^{+0.50} _{-0.41}	7.20E-08	1.59E-03	2329	0.38
51258.497	1104	H	276.02 ^{+6.64} _{-6.20}	36.77 ^{+23.00} _{-12.36}	5.22 ^{+0.01} _{-0.76}	2.91E-04	9.86E-01	1816	0.35
51259.253	880	H	274.87 ^{+4.30} _{-3.61}	26.00 ^{+14.57} _{-8.68}	5.09 ^{+1.00} _{-0.67}	7.53E-05	7.90E-01	1710	0.39
51291.184	1536	T	134.75 ^{+1.76} _{-1.45}	7.81 ^{+5.16} _{-3.25}	6.84 ^{+1.28} _{-1.09}	8.45E-04	1.00E+00	88	0.42
51664.409	2240	H	274.40 ^{+3.43} _{-3.47}	41.96 ^{+19.88} _{-10.66}	6.66 ^{+1.23} _{-0.64}	1.11E-07	1.42E-03	1347	0.29
51664.637	2864	H	274.97 ^{+2.90} _{-3.21}	33.51 ^{+13.97} _{-8.16}	5.42 ^{+0.87} _{-0.51}	4.65E-08	7.47E-04	1419	0.31
51665.406	2144	H	264.66 ^{+7.09} _{-3.19}	42.87 ^{+22.01} _{-15.43}	6.50 ^{+1.35} _{-1.02}	6.87E-04	1.00E+00	1105	0.28
51668.829	2976	H	263.29 ^{+9.41} _{-9.96}	53.90 ^{+30.55} _{-17.58}	5.50 ^{+1.56} _{-0.75}	1.36E-04	7.44E-01	856	0.26
GRO J1655-40									
50296.311	9056	T	273.72 ^{+3.97} _{-3.96}	70.33 ^{+19.40} _{-10.92}	1.15 ^{+0.15} _{-0.08}	1.15E-13	1.483E-9	7443	0.67
50301.665	6464	T	307.16 ^{+8.17} _{-6.77}	68.46 ^{+0.00} _{-38.07}	0.88 ^{+0.21} _{-0.11}	3.91E-05	4.04E-01	6414	0.61
50311.391	1856	T	285.56 ^{+8.91} _{-11.19}	63.78 ^{+42.48} _{-17.38}	1.14 ^{+0.36} _{-0.16}	1.93E-04	9.35E-01	5450	0.54
50317.441	6176	H	443.16 ^{+4.12} _{-3.79}	41.44 ^{+20.27} _{-13.66}	5.39 ^{+0.81} _{-0.64}	1.28E-05	2.44E-01	5210	0.51
50324.380	4896	T	289.67 ^{+6.26} _{-6.03}	98.78 ^{+40.67} _{-17.67}	1.41 ^{+0.30} _{-0.11}	1.59E-10	1.46E-06	7267	0.65
50330.254	6144	H	453.72 ^{+6.29} _{-5.32}	37.69 ^{+16.46} _{-12.16}	5.69 ^{+0.89} _{-0.71}	2.91E-05	5.03E-01	4964	0.49
50335.913	8160	H	446.35 ^{+4.10} _{-4.15}	38.74 ^{+9.80} _{-8.00}	6.03 ^{+0.59} _{-0.54}	9.01E-09	2.11E-04	4983	0.50
50383.565	6192	H	442.34 ^{+4.71} _{-4.28}	31.50 ^{+15.62} _{-10.61}	4.55 ^{+0.70} _{-0.62}	1.31E-04	9.77E-01	6092	0.54
50394.884	5792	T	313.45 ^{+9.38} _{-9.81}	98.11 ^{+38.42} _{-19.10}	1.17 ^{+0.25} _{-0.11}	3.72E-08	3.44E-04	5497	0.59
53498.415	13968	T	125.81 ^{+1.10} _{-1.20}	7.49 ^{+3.86} _{-2.91}	0.65 ^{+0.11} _{-0.09}	2.16E-04	1.00E+00	3190	0.29
53508.507	7504	T	266.63 ^{+7.96} _{-7.80}	61.76 ^{+30.61} _{-13.60}	0.76 ^{+0.19} _{-0.09}	2.67E-05	3.24E-01	10164	0.57

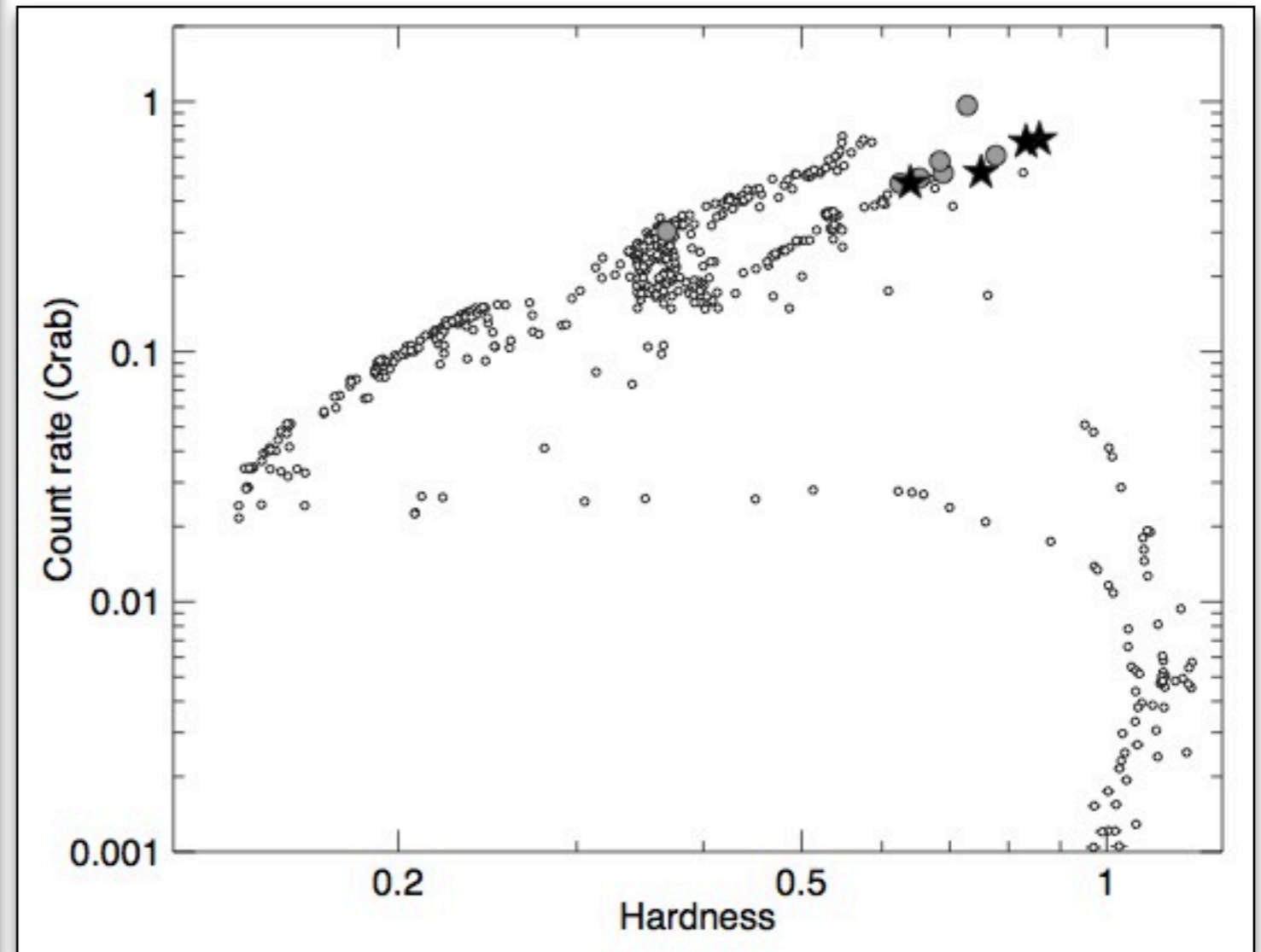
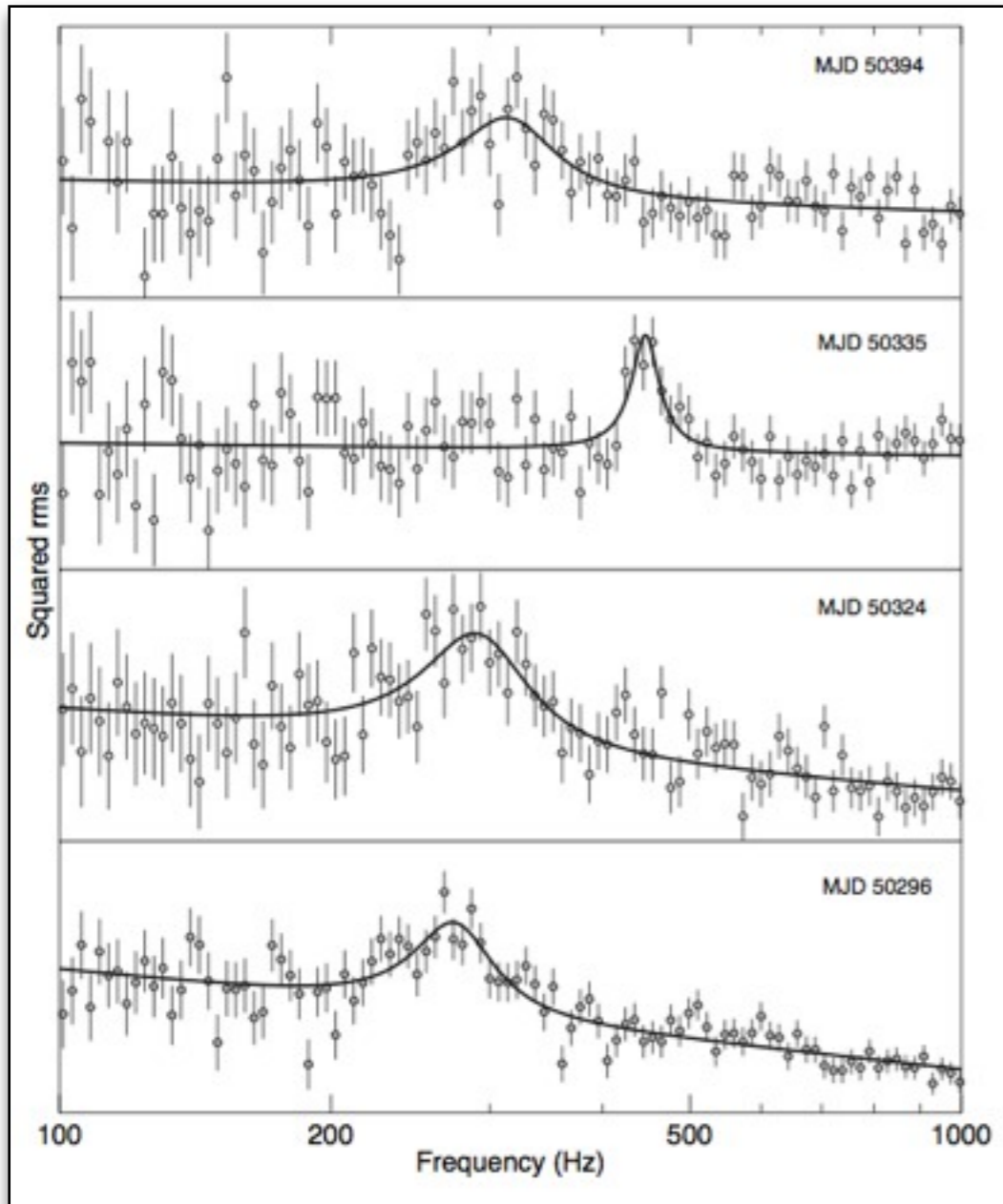
XTE J1550+564

Belloni, Sanna & Méndez (2012)



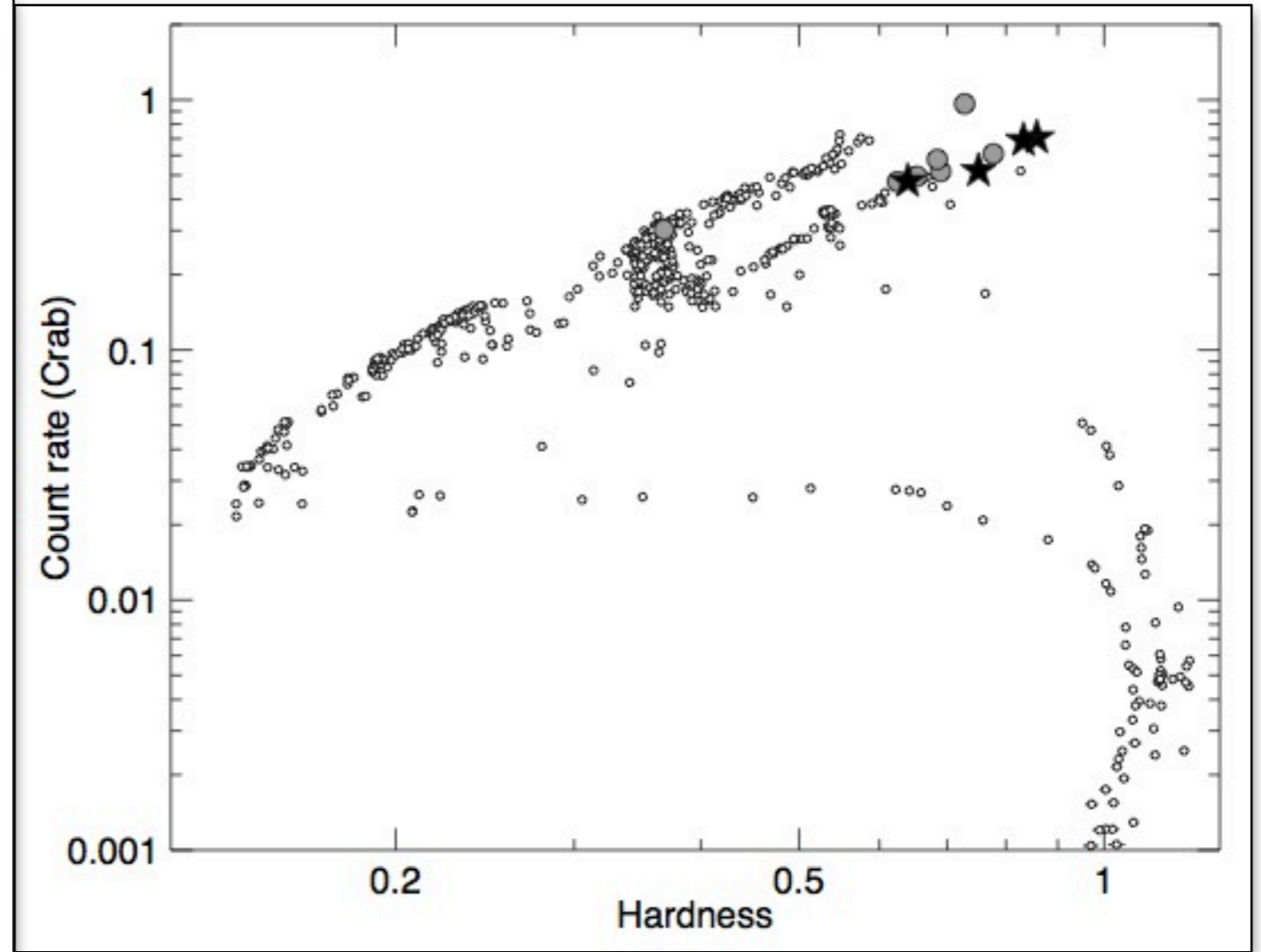
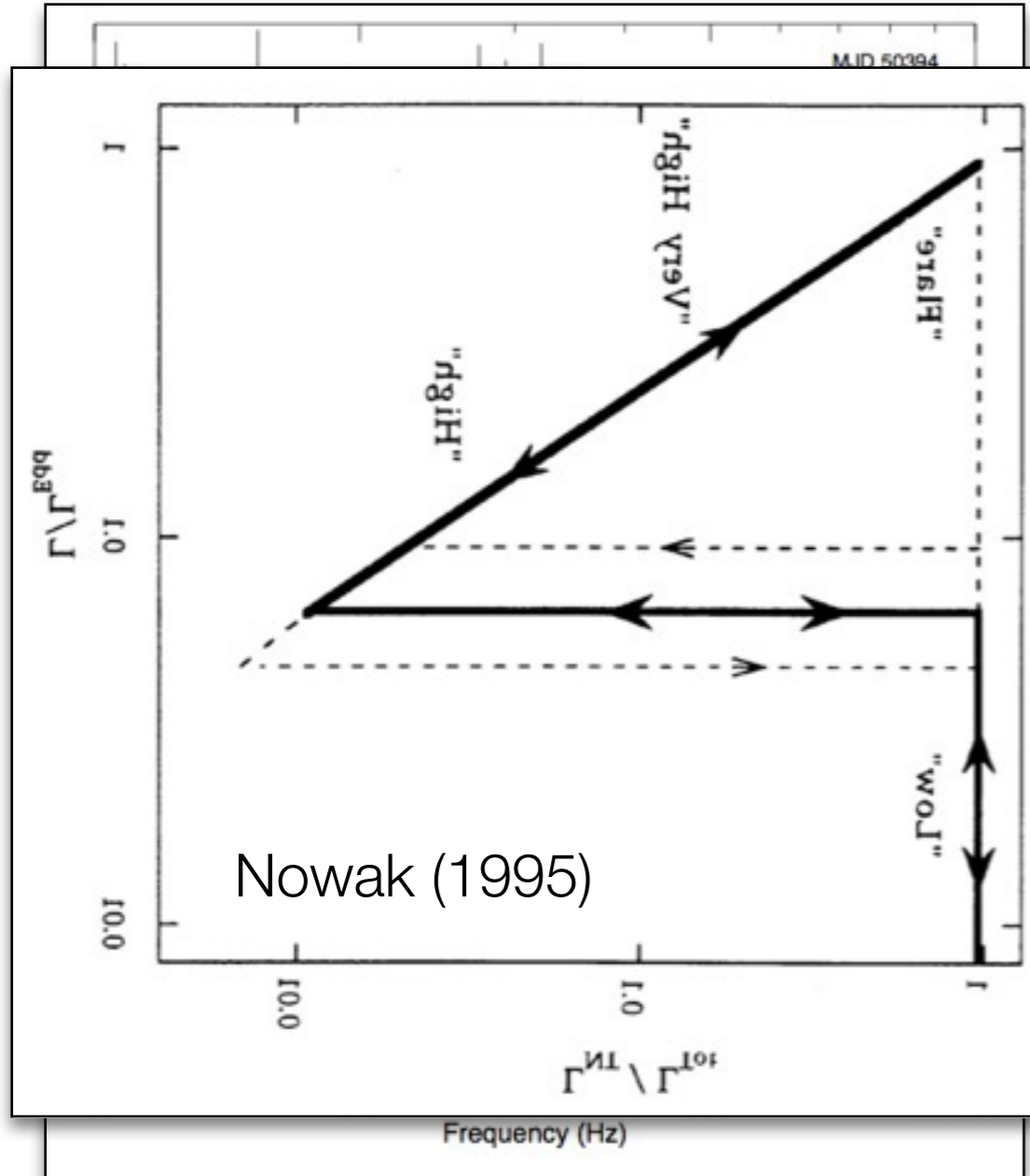
GRO J1655-40

Belloni, Sanna & Méndez (2012)



GRO J1655-40

Belloni, Sanna & Méndez (2012)



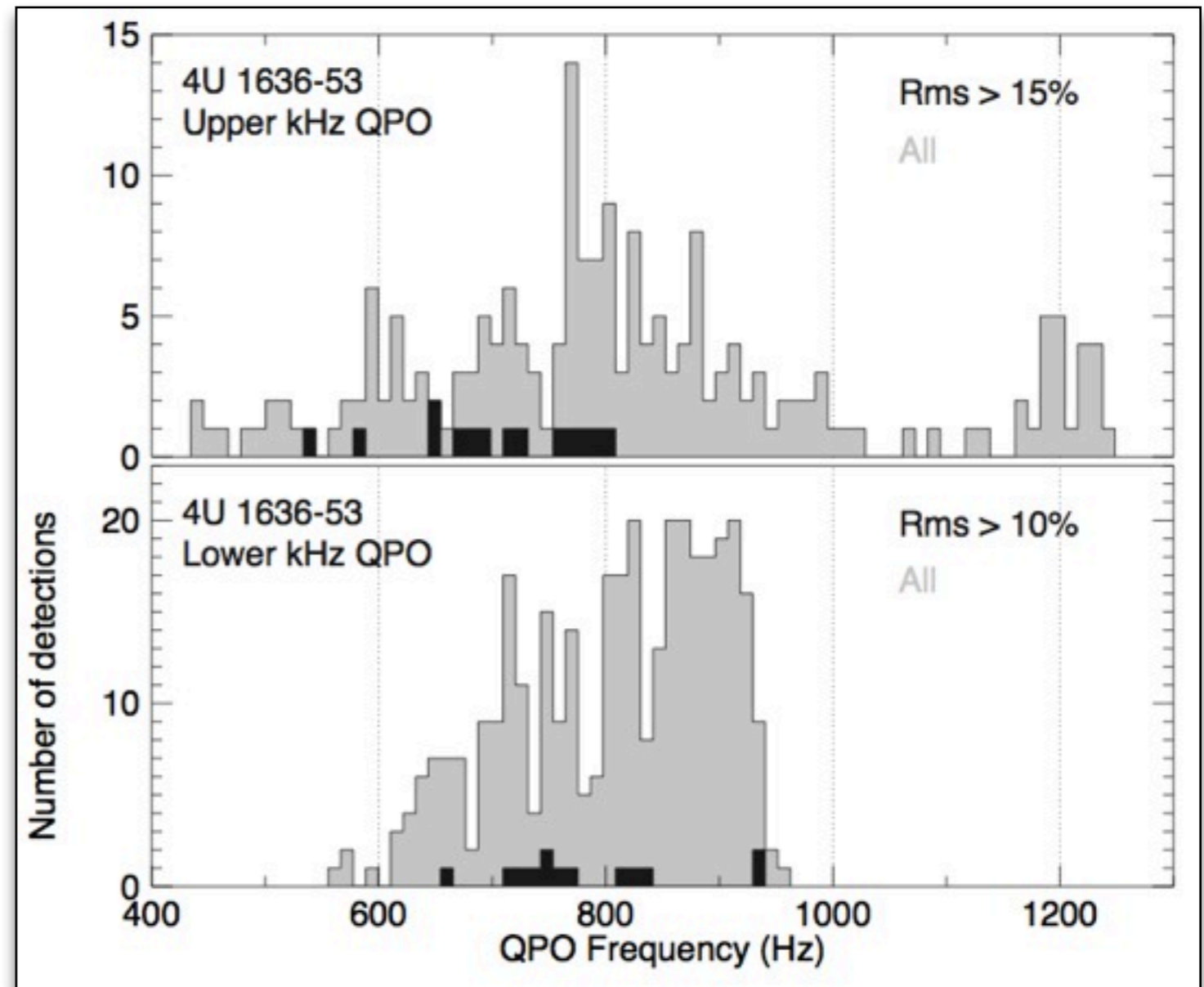
What we have

Belloni, Sanna & Méndez (2012)

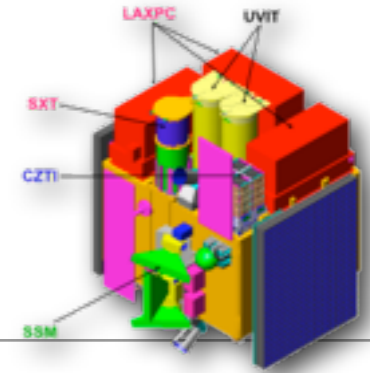


- Few detections (others)
- Special ratios?
- NO** type-C QPO
- Only intermediate states

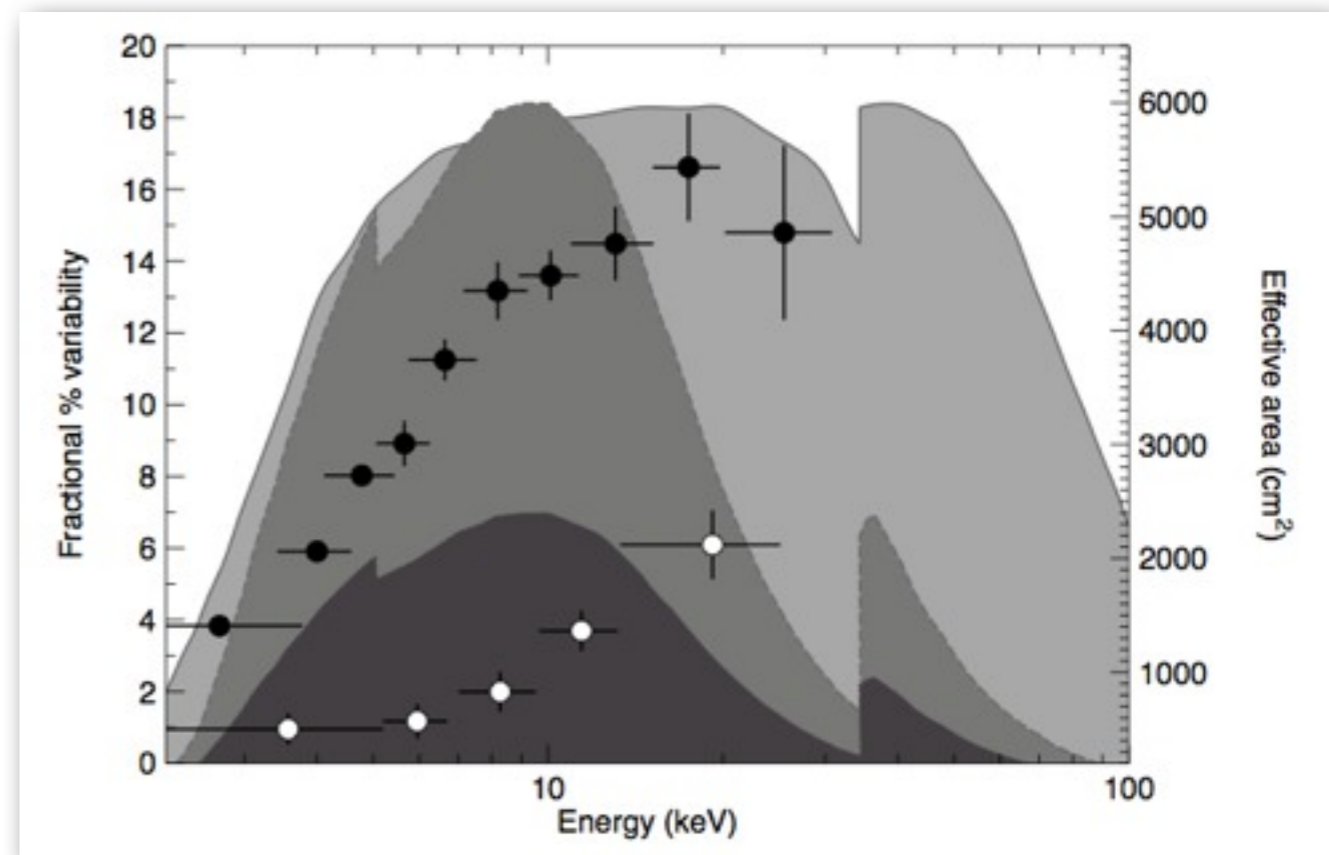
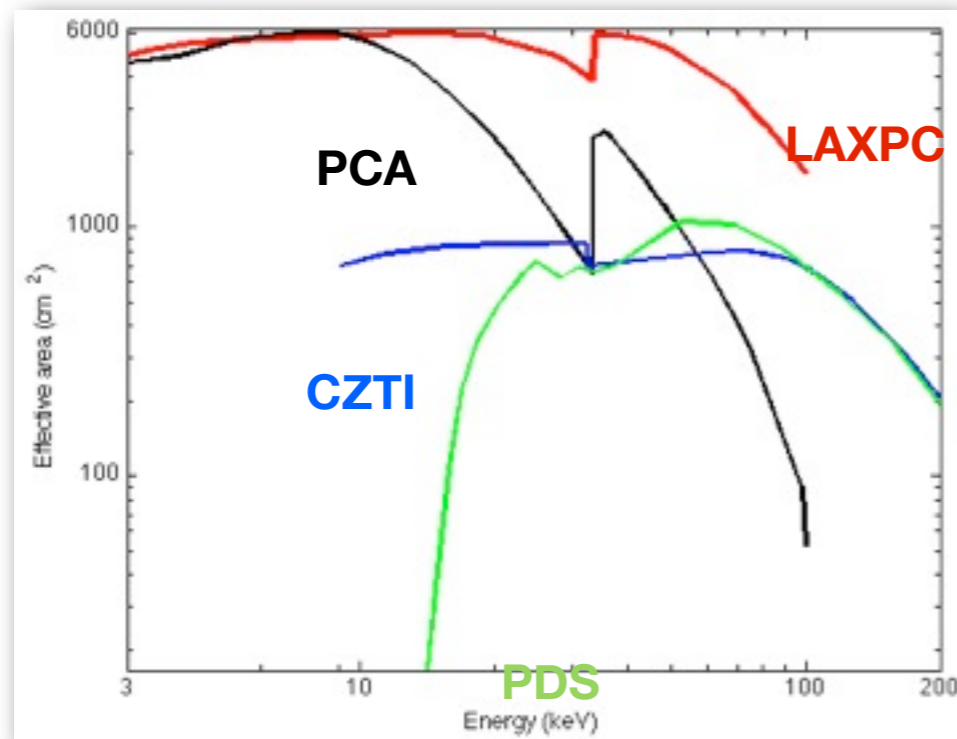
- Higher ν :
 - Narrower peak
 - Harder spectrum
- Same as kHz QPOs in NS?
- Boundary layer effects?



RXTE and THE FUTURE



- The near future: **ASTROSAT** (2013?)
- Better than RXTE above > 20 keV
- Coverage?



LOFT



- Selected for assessment by ESA
- Launched in 2020+ ?
- 12 m² effective area

