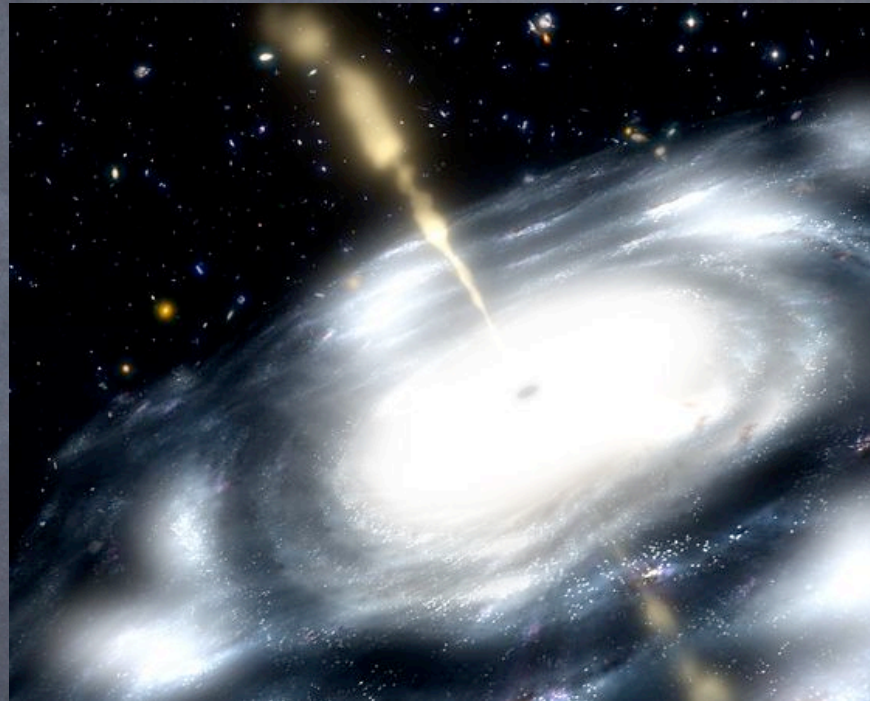


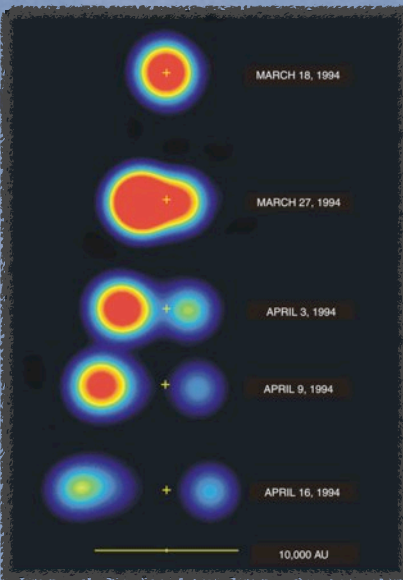
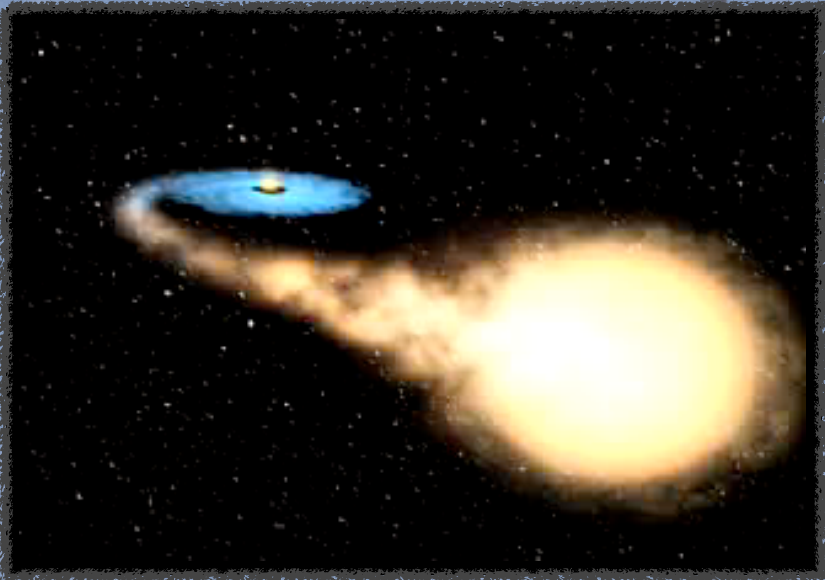
# Expanding the Efficacy of the Fundamental Plane of Black Hole Activity



Rich Plotkin

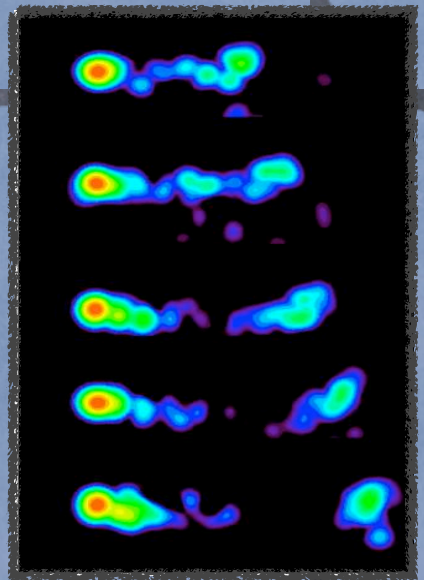
University of Amsterdam

with Sera Markoff (Amsterdam), Brandon Kelly (CfA),  
Elmar Körding (Nijmegen) and Scott Anderson (U. Washington)



Stellar  
Mass  
Black  
Hole  
( $\sim 10 M_{\text{Sun}}$ )

GRS 1915+105



Super-  
massive  
Black Hole  
( $10^6 - 10^9 M_{\text{Sun}}$ )

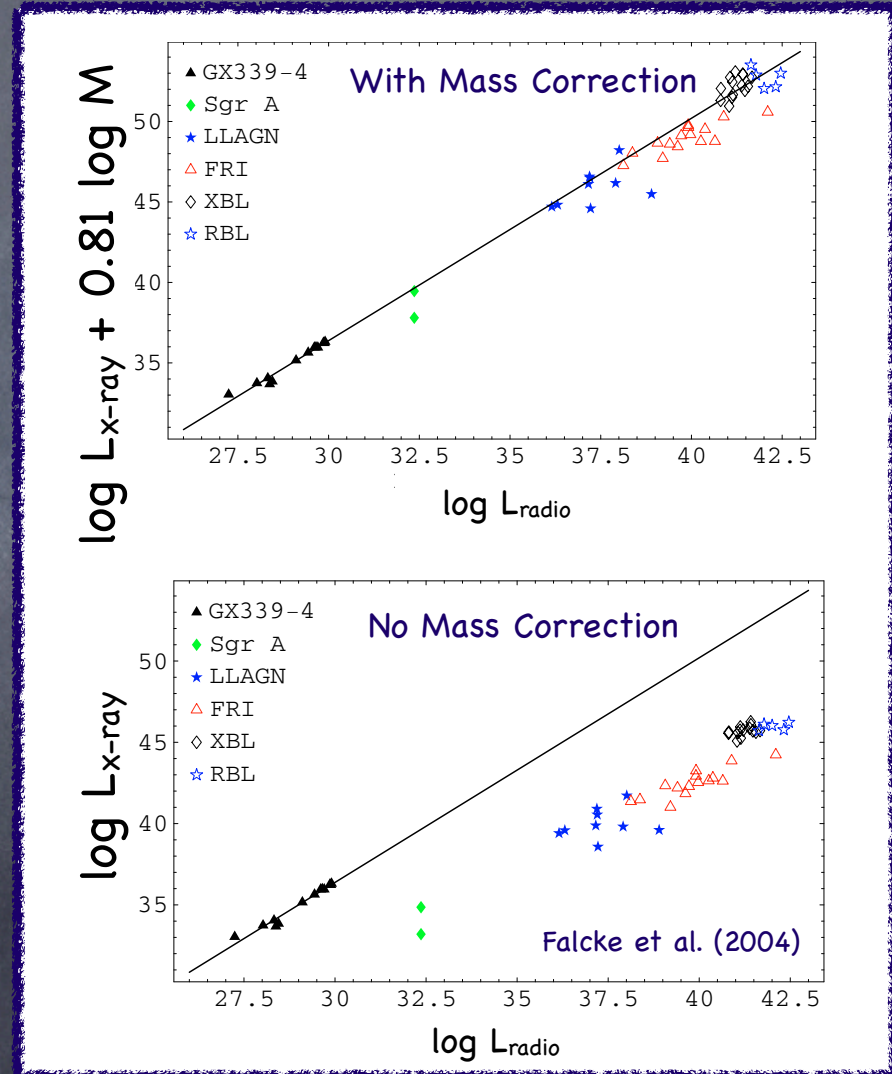
3C 279

# Fundamental Plane of Black Hole Activity

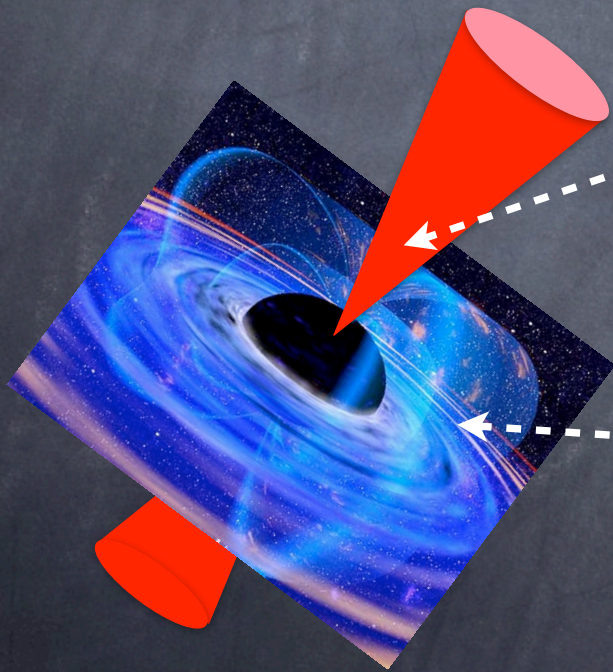
(Merloni et al. 2003, Falcke et al. 2004)

For “hard state” BHs:  
the conversion of  
the accretion flow  
into radiative output  
is universal across  
the black hole mass  
scale

Also a “variability” plane  
(e.g., McHardy et al. 2006,  
Körding et al. 2006)



# Fundamental Plane as a Tool to Diagnose X-ray Radiative Processes/Geometries



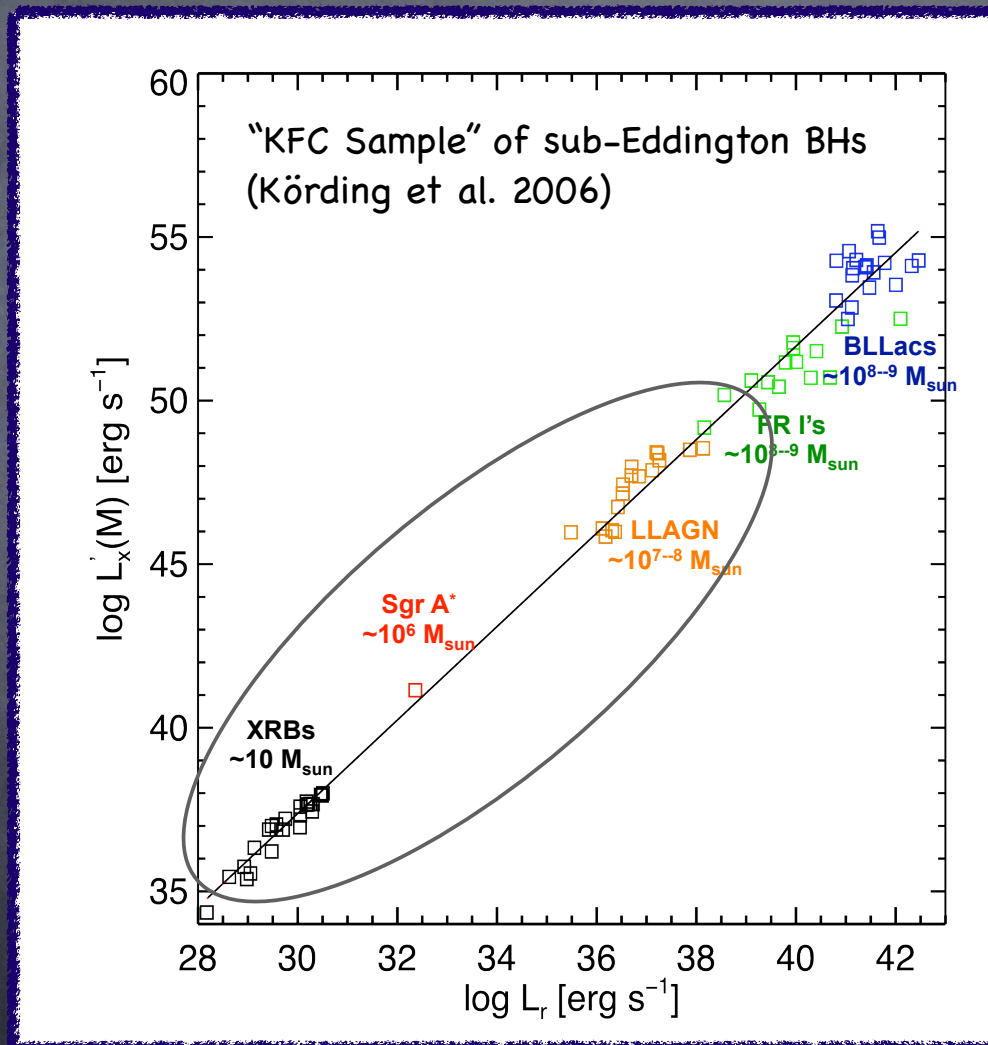
Optically Thin Jet Synchrotron X-rays:

$$L_{\text{x-ray}} \sim (L_{\text{radio}})^{1.38} M^{-0.81}$$

X-rays from Radiatively Inefficient Accretion Flow:

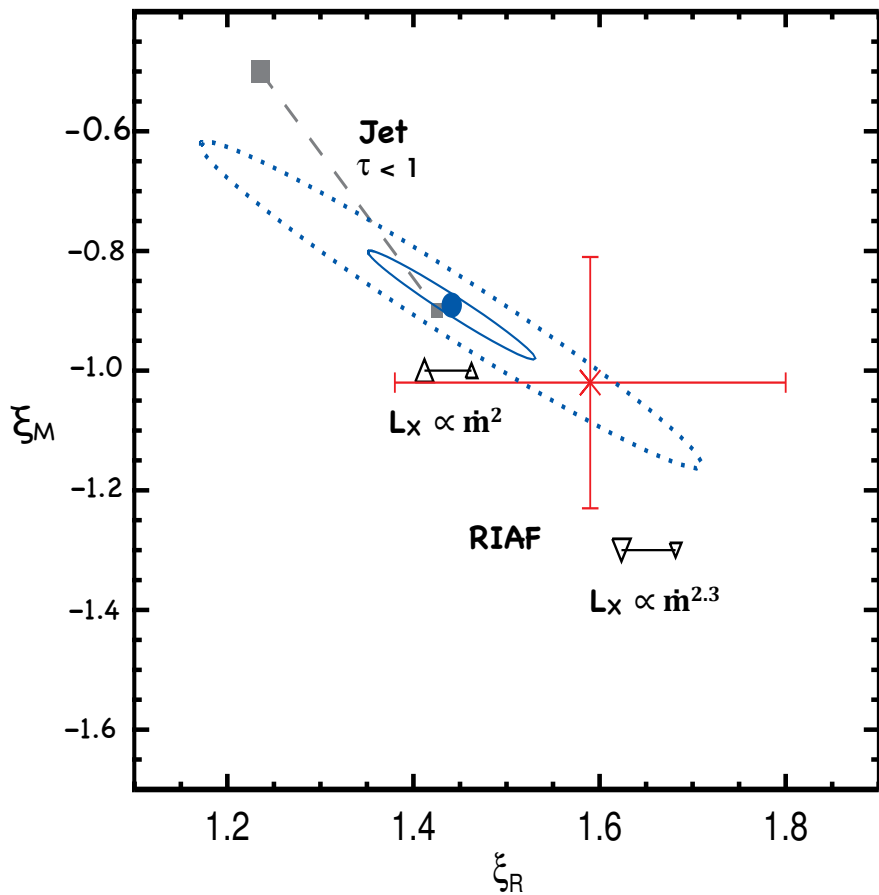
$$L_{\text{x-ray}} \sim (L_{\text{radio}})^{1.76} M^{-1.56}$$

# Finding the best-fit slope is a non-trivial statistical problem



# X-rays from the "average" low $L/L_{\text{Edd}}$ BHs are Dominated by Optically Thin Jet Synchrotron

43 Galactic BH, SgrA\*, LLAGN  
 $10 M_{\text{Sun}} - 10^7 M_{\text{Sun}}$



$$L_{\text{x-ray}} \sim (L_{\text{radio}})^{\xi_R} M^{\xi_M}$$



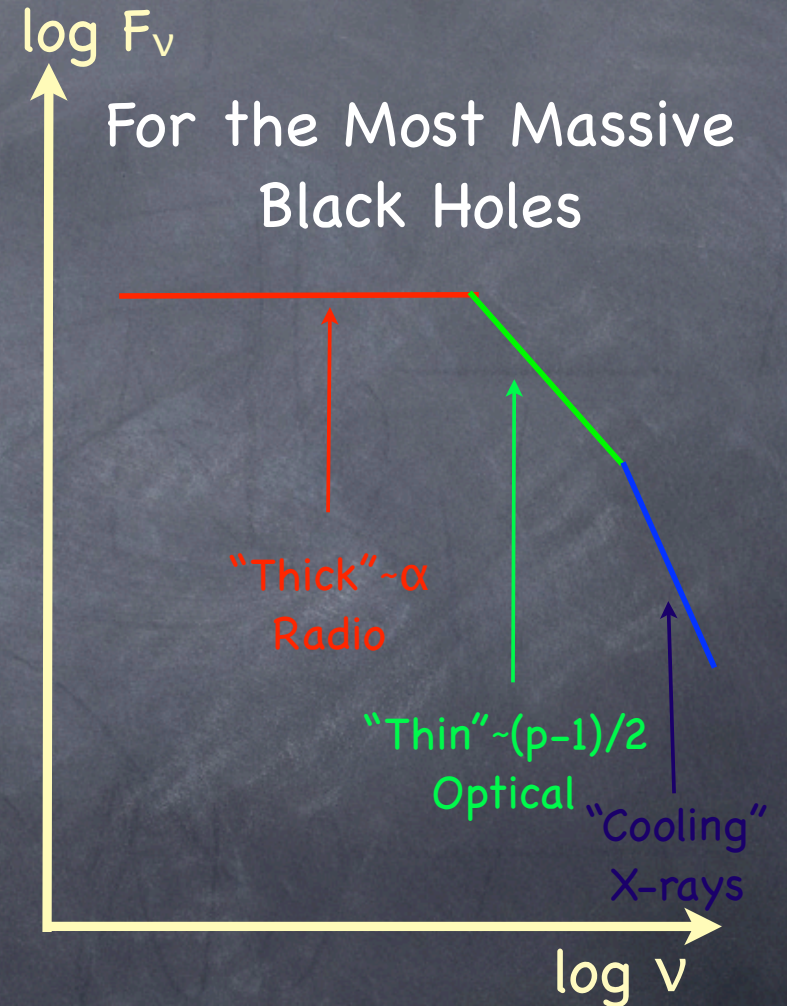
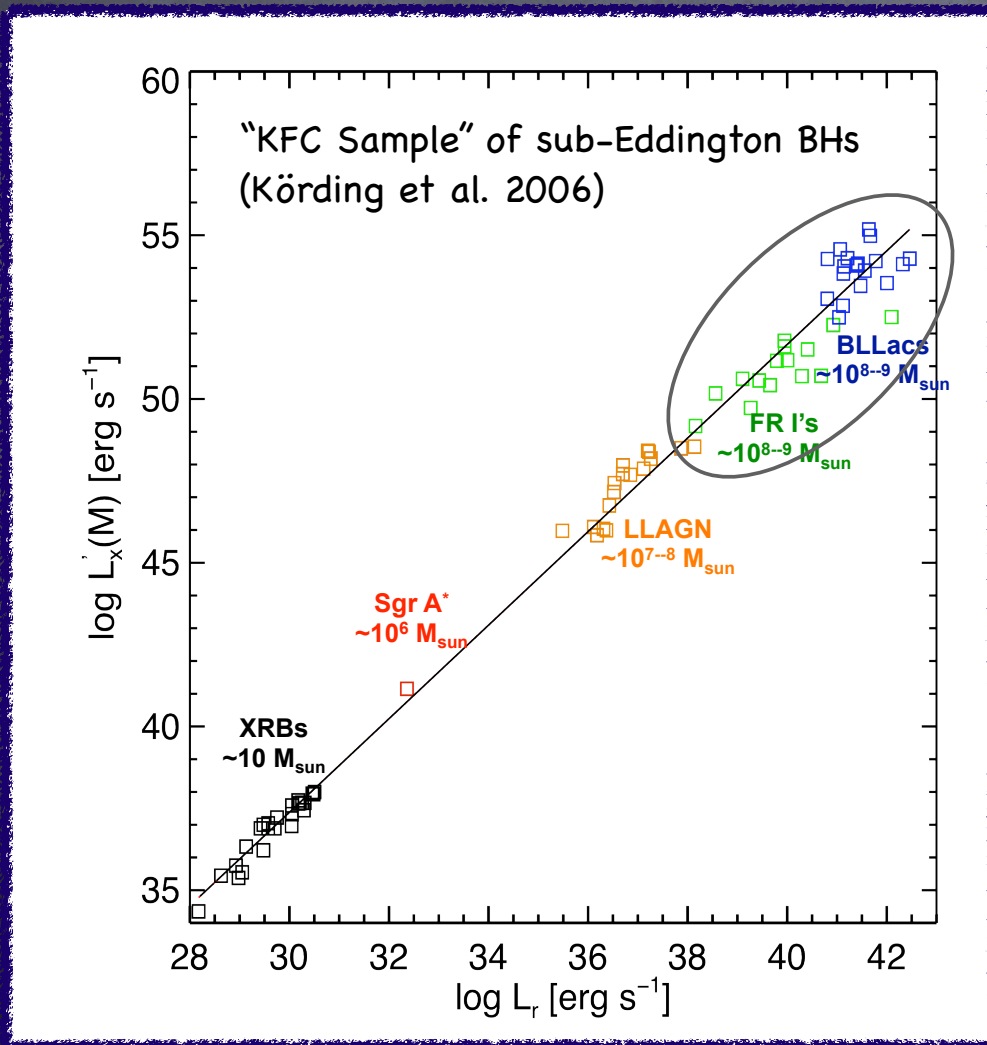
Bayesian  
Regression



Merit  
Function

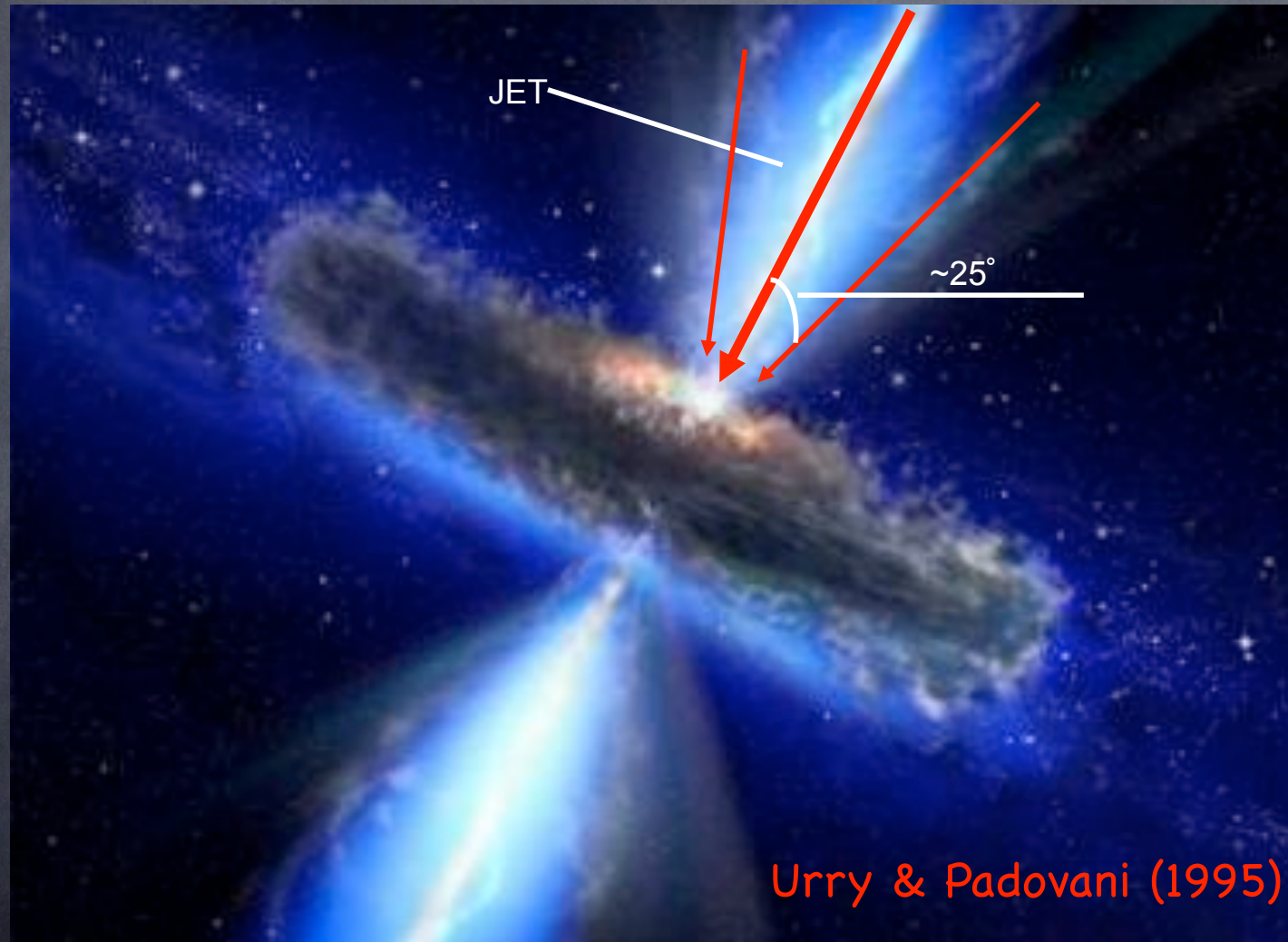
- Merit Function is a  $\chi^2$  estimator (from Körding et al. 2006)
- Bayesian Linear Regression (application of Kelly 2007 technique)

# The Massive End Poses a Challenge: Synchrotron Cooling (Heinz 2004\*)



\*Also see Merloni et al. 2003;  
Falcke et al. 2004; Körding et al. 2006

# Explore the Massive End with BL Lacs: Uncontroversially Jet Dominated AGN

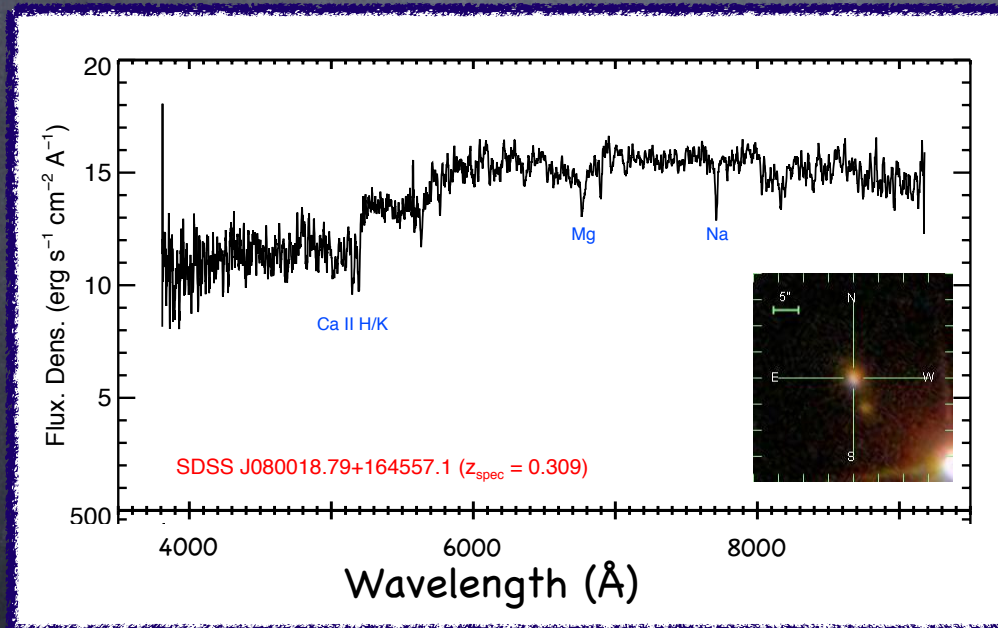




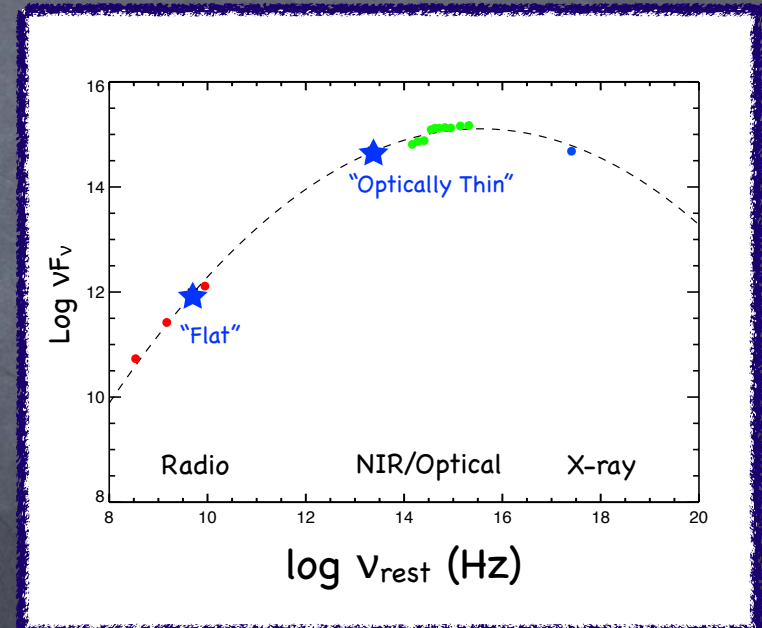
# 55 BL Lacs from the SDSS with Black Hole Mass Measurements

(Plotkin et al. 2010; 2011)

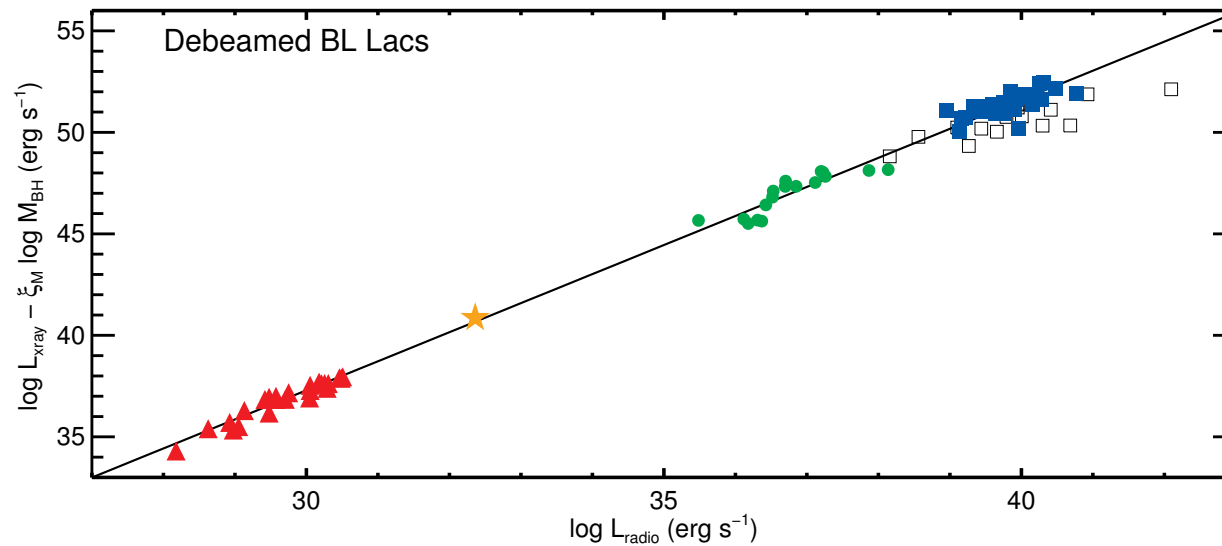
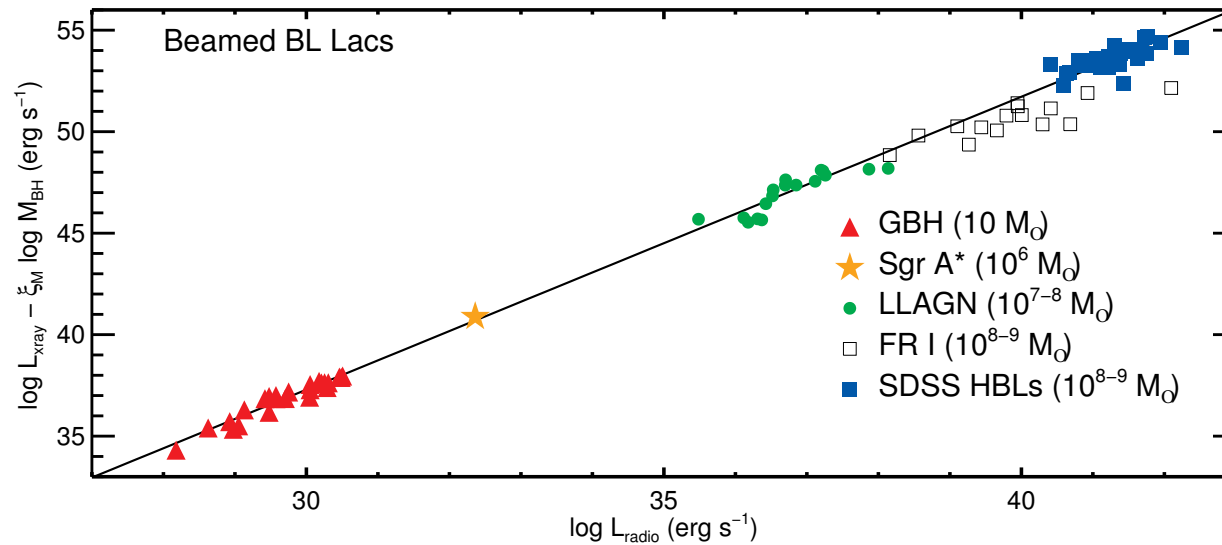
## SDSS Spectrum



## SED



$$\log L_X = (1.44 \pm 0.04) \log L_R - (0.88 \pm 0.06) \log M_{\text{BH}} - 6.03 \pm 1.16$$



# Summary

1. Bayesian Regression is Very Promising for Refining the Fundamental Plane
2. X-rays from the "average" low-accretion rate BH are dominated by optically thin jet synchrotron
3. There are Systematic \*Observational\* Challenges to Including the Most Massive Black Holes
  - Need to account for how SED "shifts" with black hole mass