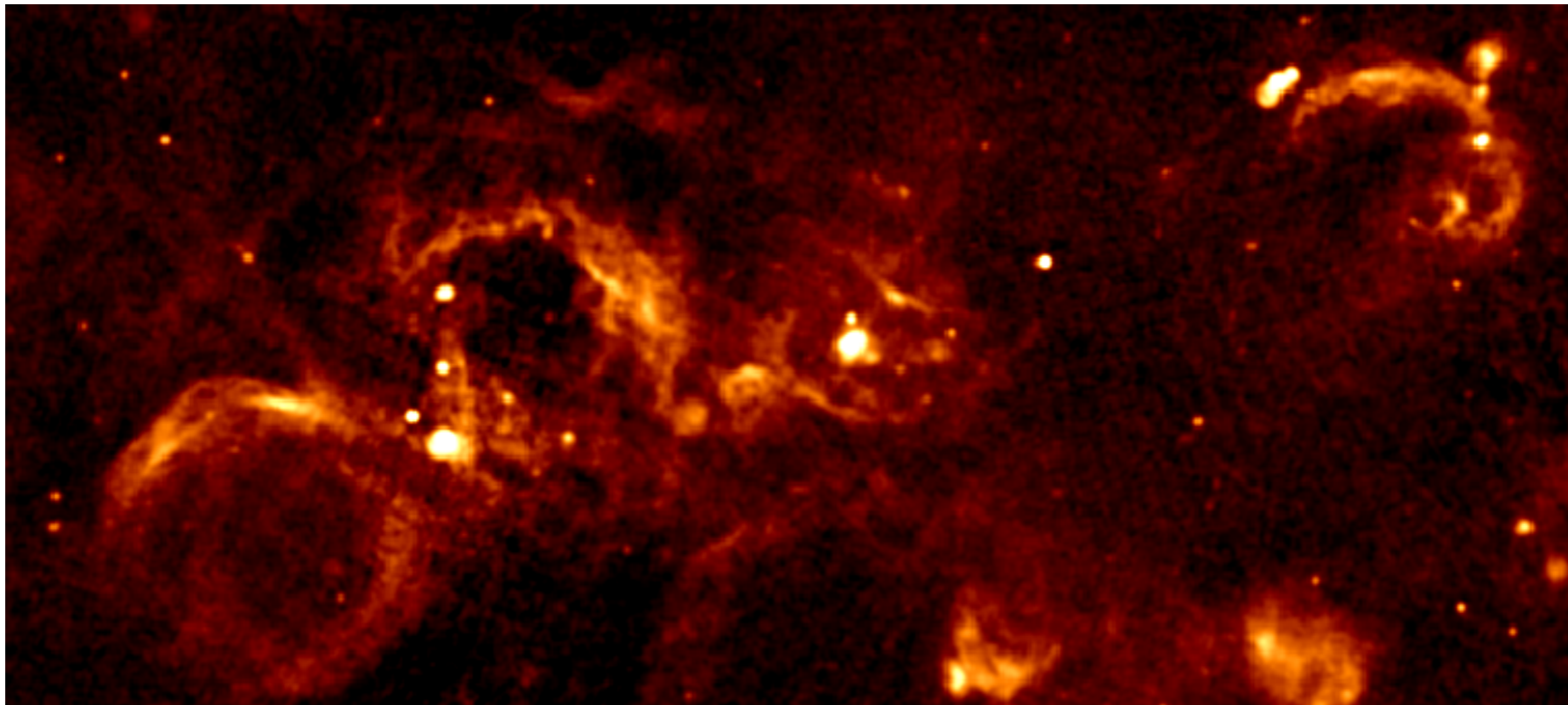


# G53.41+0.03 – A NEWLY DISCOVERED GALACTIC SUPERNOVA REMNANT

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**JACCO VINK**  
**JASON W.T. HESSELS**  
**MARIA ARIAS**  
**JOSEPH D. GELFAND**

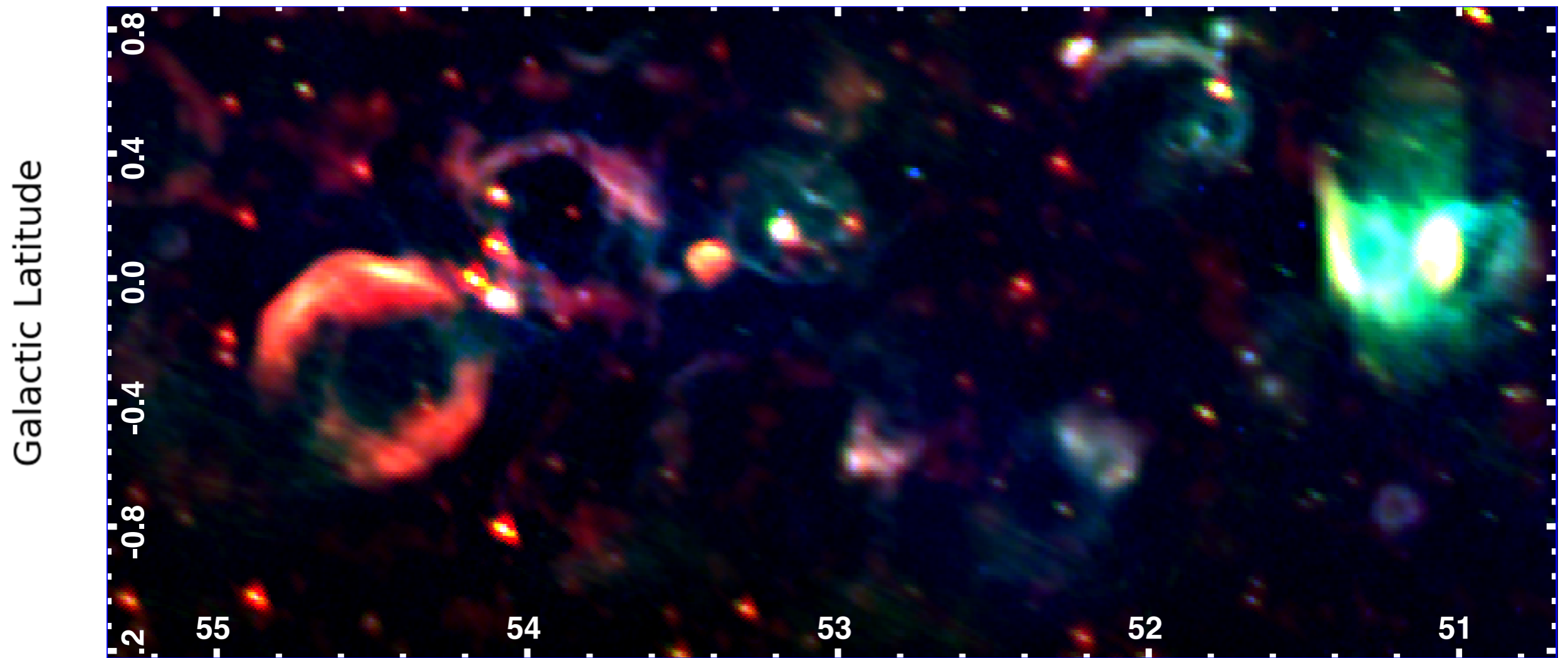
- ▶ Only ~300 known SNR (Green catalogue),
- ▶ 5-10x more expected (Li+1991)
- ▶ SNR and HII regions hard to distinguish
  - ▶ X-ray data
  - ▶ Radio data of 2 frequencies



VLA 1.4 GHz

# LOFAR OBSERVATION

VLA 1.4 GHz    WSRT 327 MHz    LOFAR 150 MHz

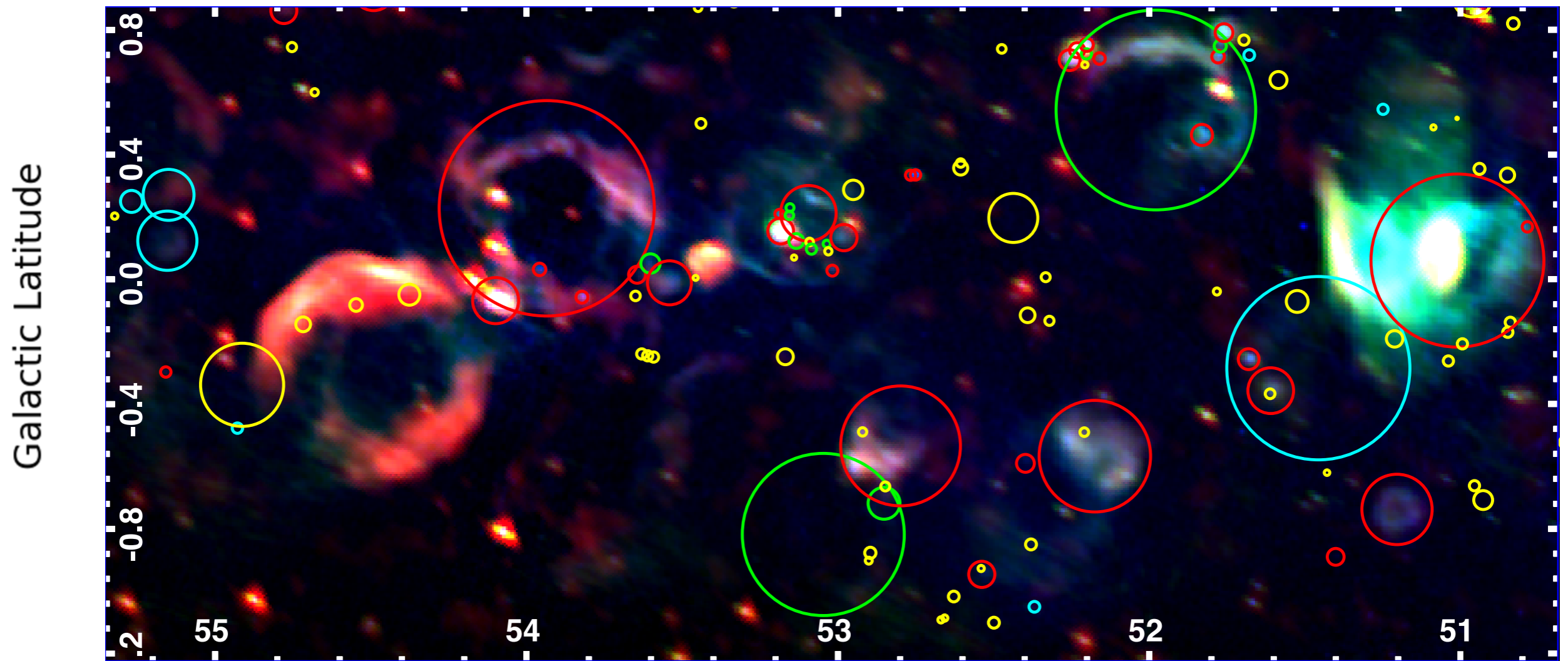


Galactic Longitude

Driessen, Domček et. al. 2017

# LOFAR OBSERVATION - HII REGIONS

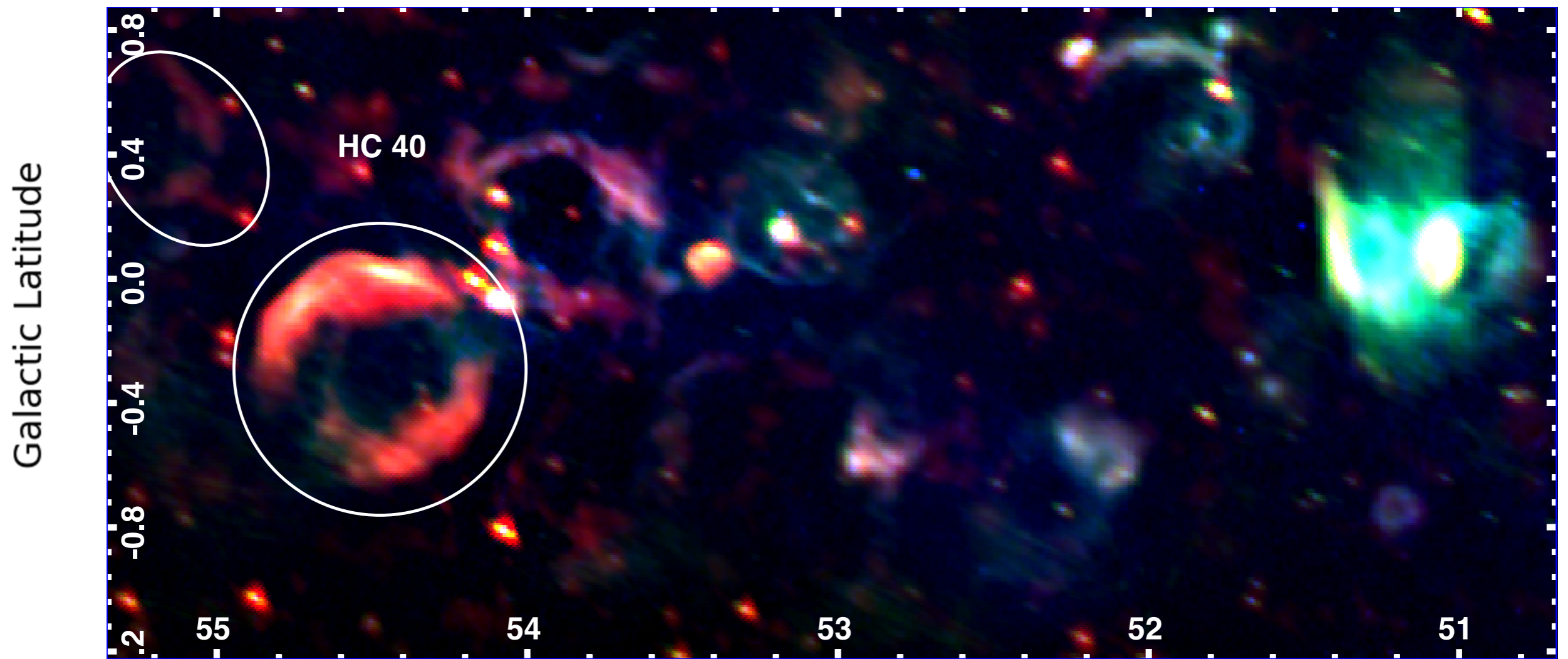
VLA 1.4 GHz    WSRT 327 MHz    LOFAR 150 MHz



Galactic Longitude

Driessen, Domček et. al. 2017

VLA 1.4 GHz    WSRT 327 MHz    LOFAR 150 MHz

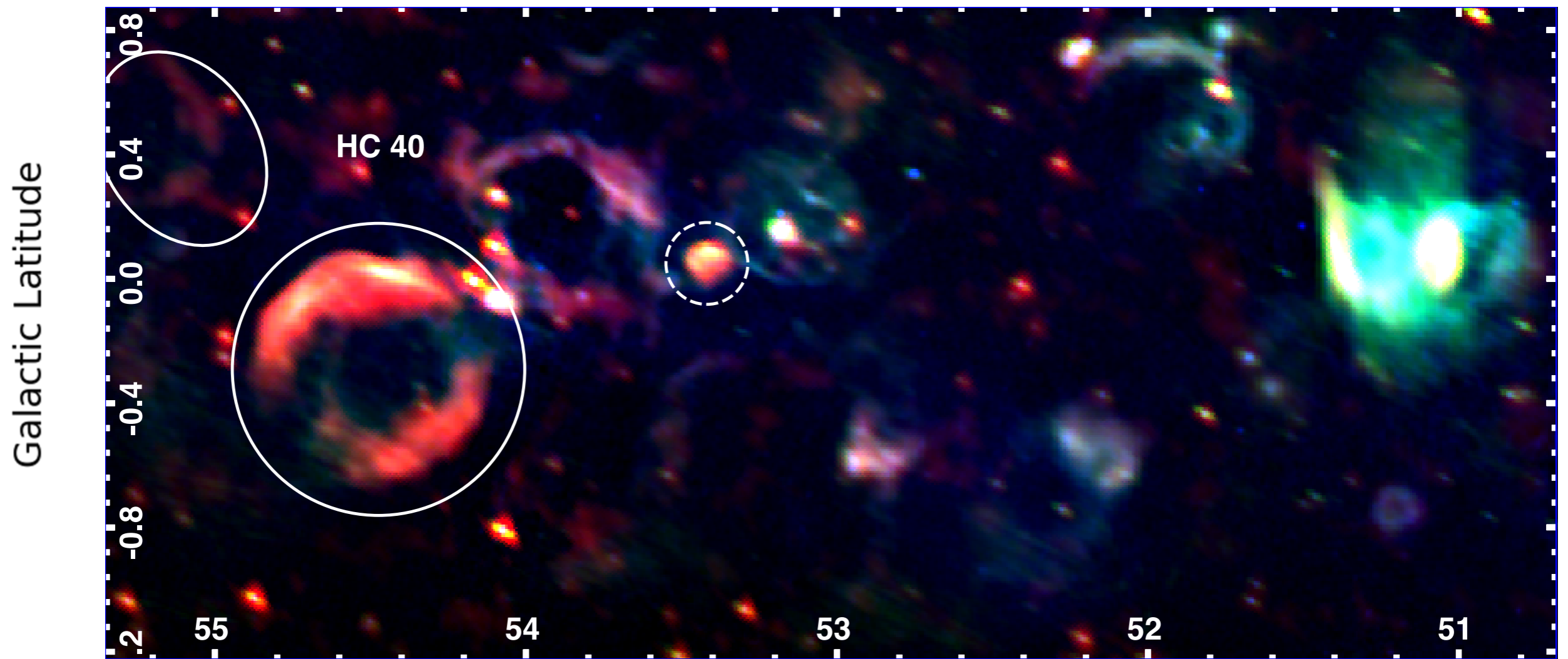


Galactic Longitude

Driessen, Domček et. al. 2017

# LOFAR OBSERVATION - NEW SNR CANDIDATE?

VLA 1.4 GHz    WSRT 327 MHz    LOFAR 150 MHz

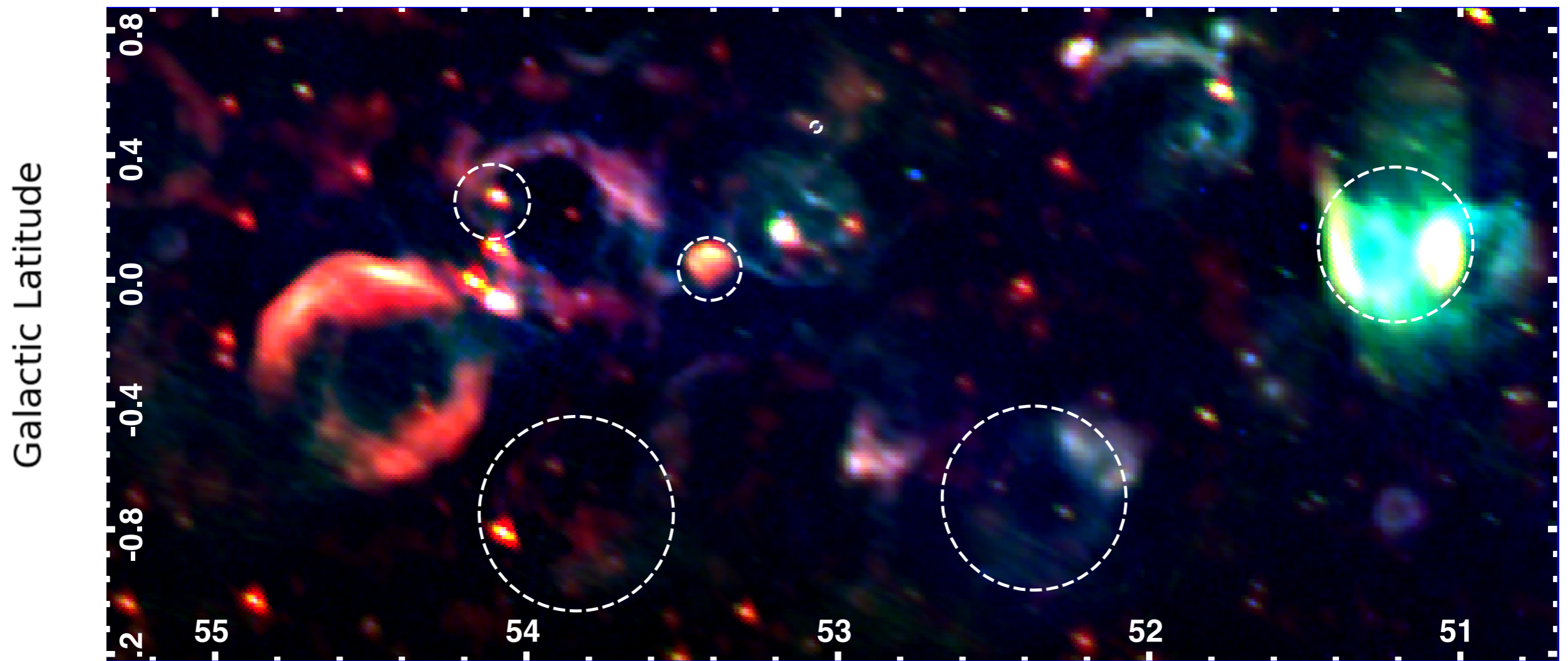


Galactic Longitude

Driessen, Domček et. al. 2017

# OTHER CANDIDATES IN LOFAR FOV

VLA 1.4 GHz    WSRT 327 MHz    LOFAR 150 MHz

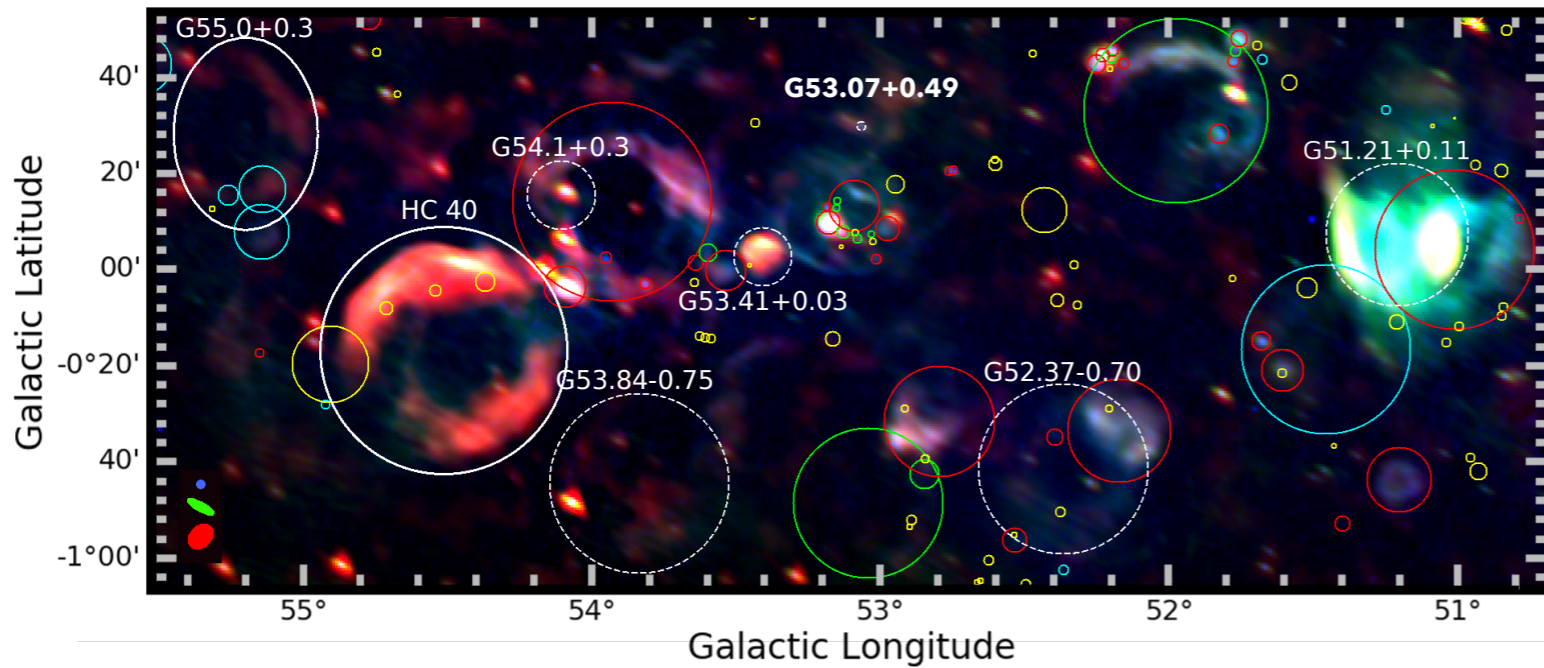


Galactic Longitude

Driessen, Domček et. al. 2017

# OBJECTS IN LOFAR FIELD OF VIEW

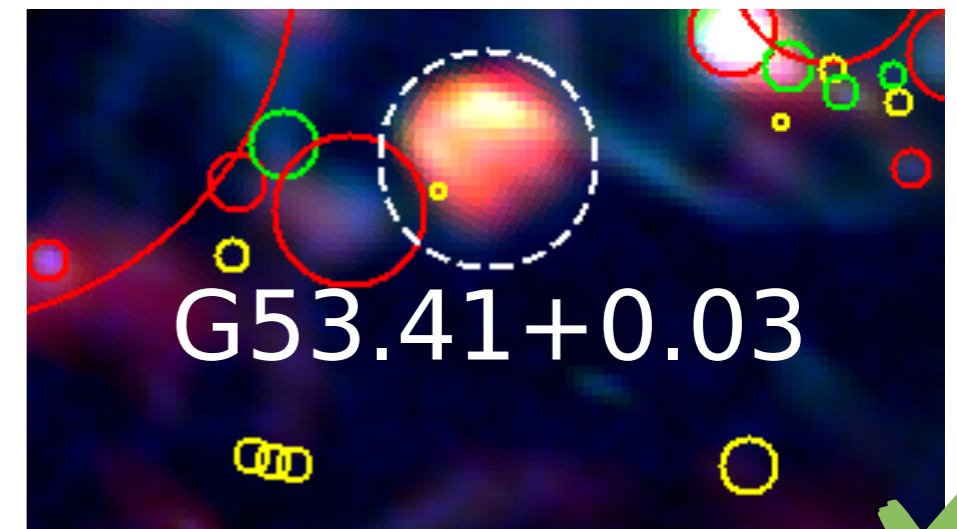
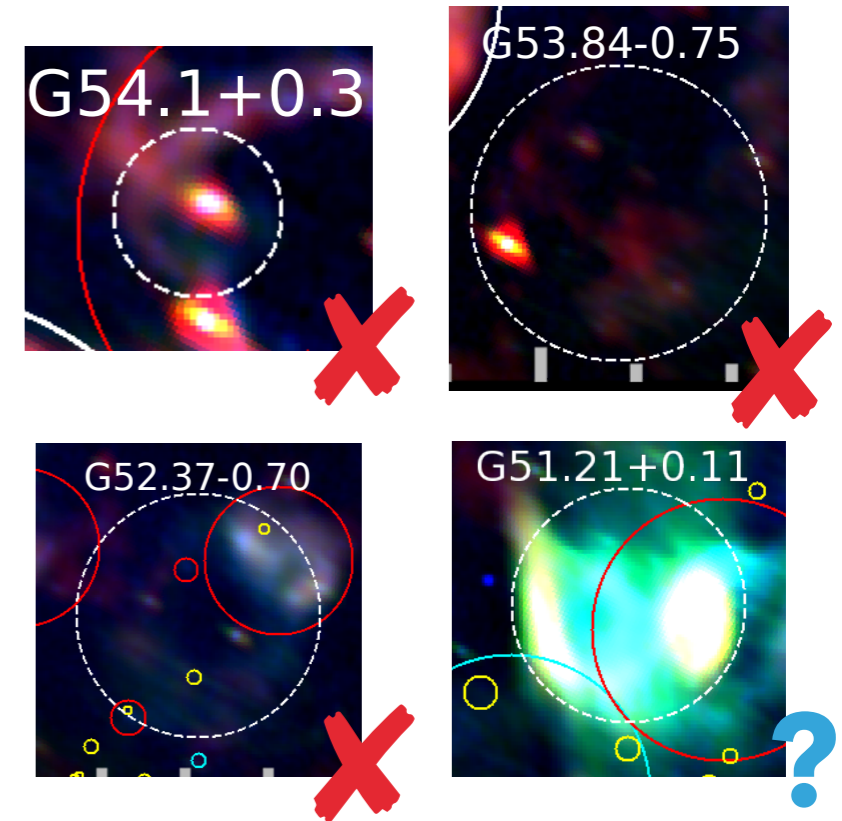
VLA 1.4 GHz    WSRT 327 MHz    LOFAR 150 MHz



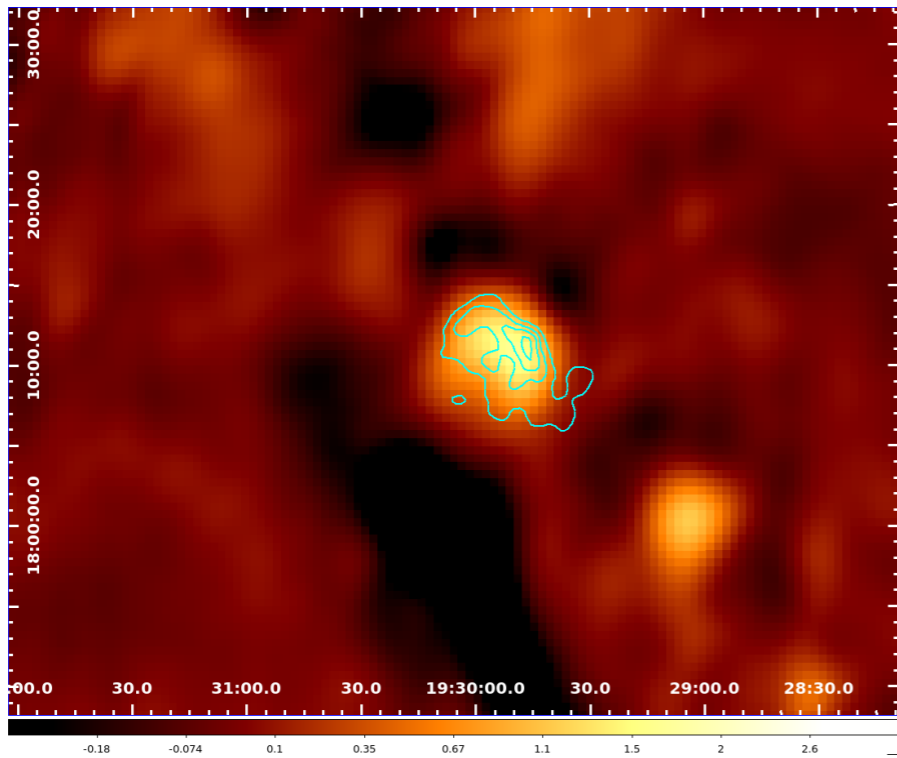
DRIESSEN ET AL.

SNR		flux density (Jy)			$\alpha$
		1.4 GHz	327 MHz	150 MHz	
?	G51.21+0.11	$24.35 \pm 2.1$	$66.1 \pm 0.1$		$-0.7 \pm 0.21$
X	G52.37-0.70	$5.24 \pm 1.75$	$3.2 \pm 0.03$		$0.3 \pm 0.3$
✓	G53.41+0.03	$1.21 \pm 0.21$	$2.2 \pm 0.03$	$3.11 \pm 0.2$	$-0.6 \pm 0.2$
X	G53.84-0.75	$1.31 \pm 3.43$	$0.06 \pm 0.02$	$1.2 \pm 0.07$	$0.05 \pm 3.9$
X	G54.1+0.3	$1.46 \pm 0.28$	$1.21 \pm 0.05$	$0.4 \pm 0.8$	$0.3 \pm 4.3$

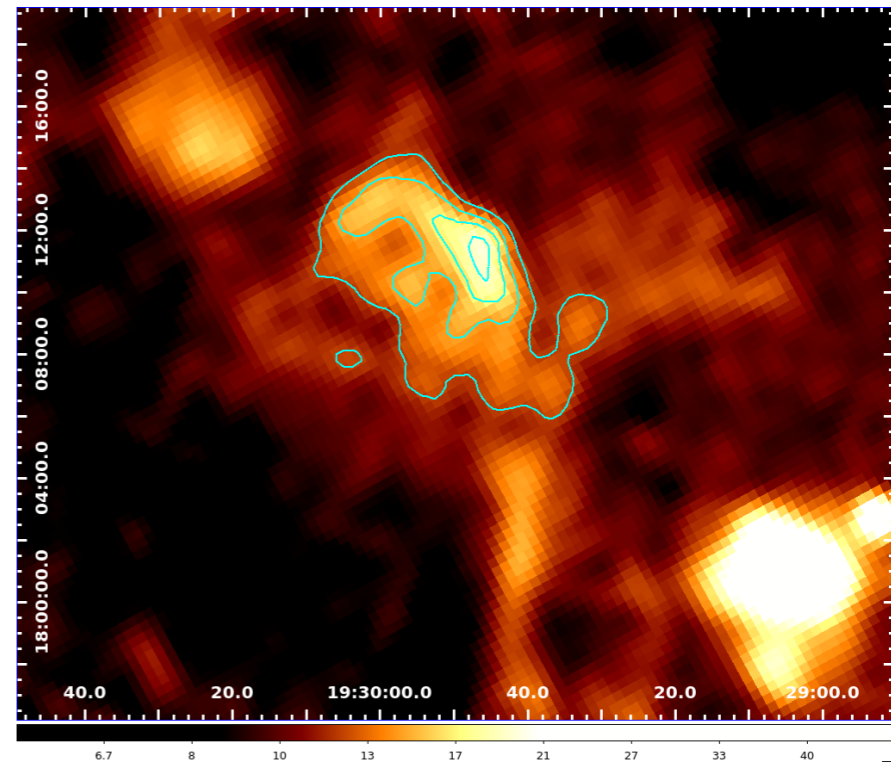
SNR Candidates from Anderson et. al. 2017



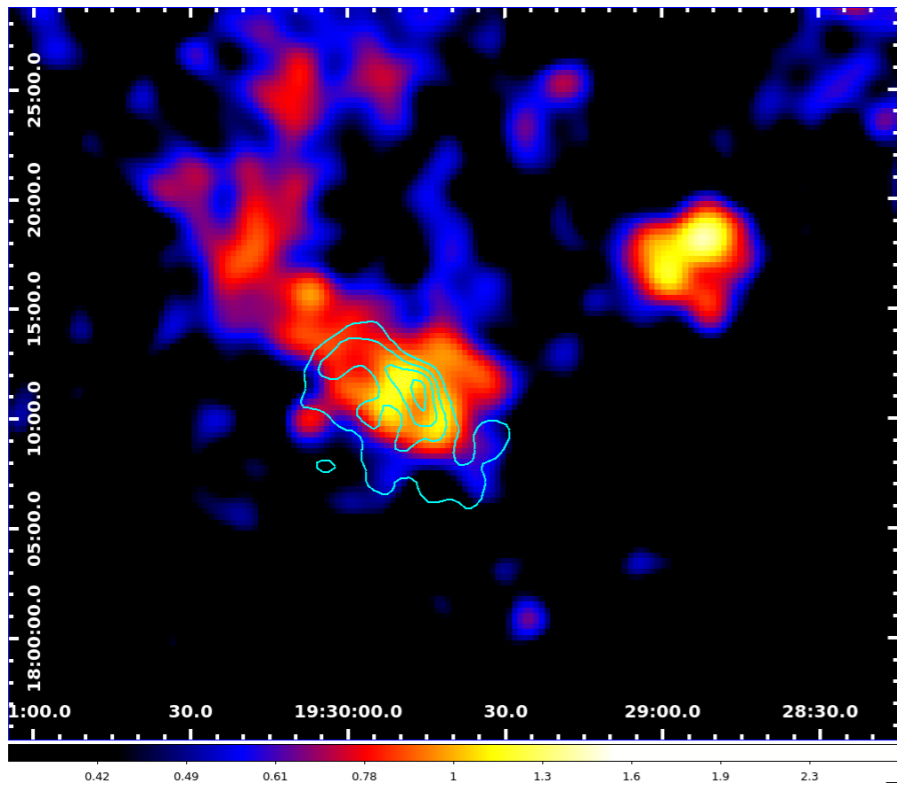




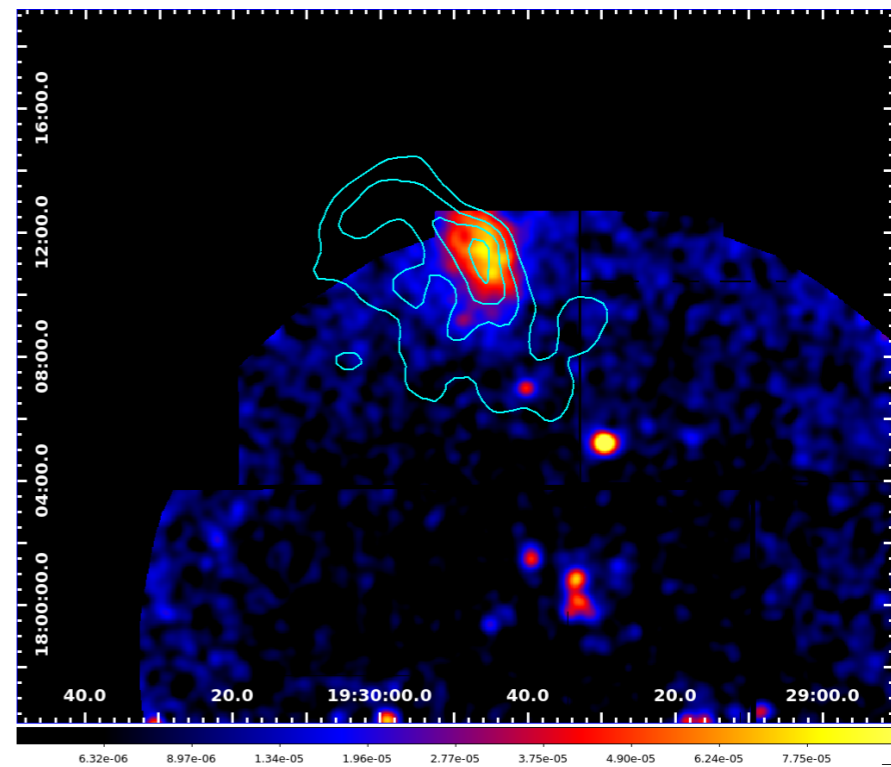
LOFAR 150 MHz



VLA 1.4 GHz



ROSAT 0.2-2.0 keV

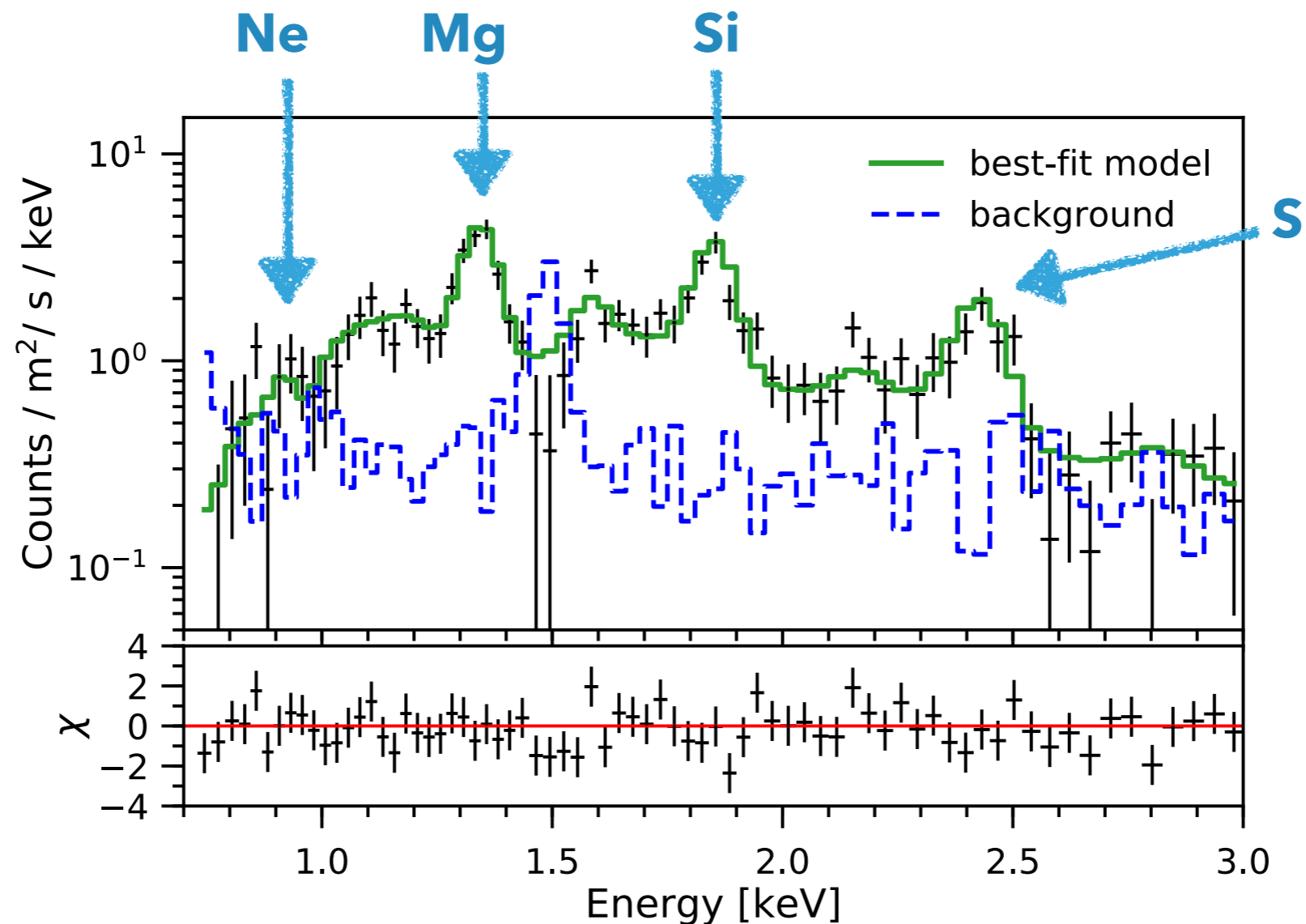


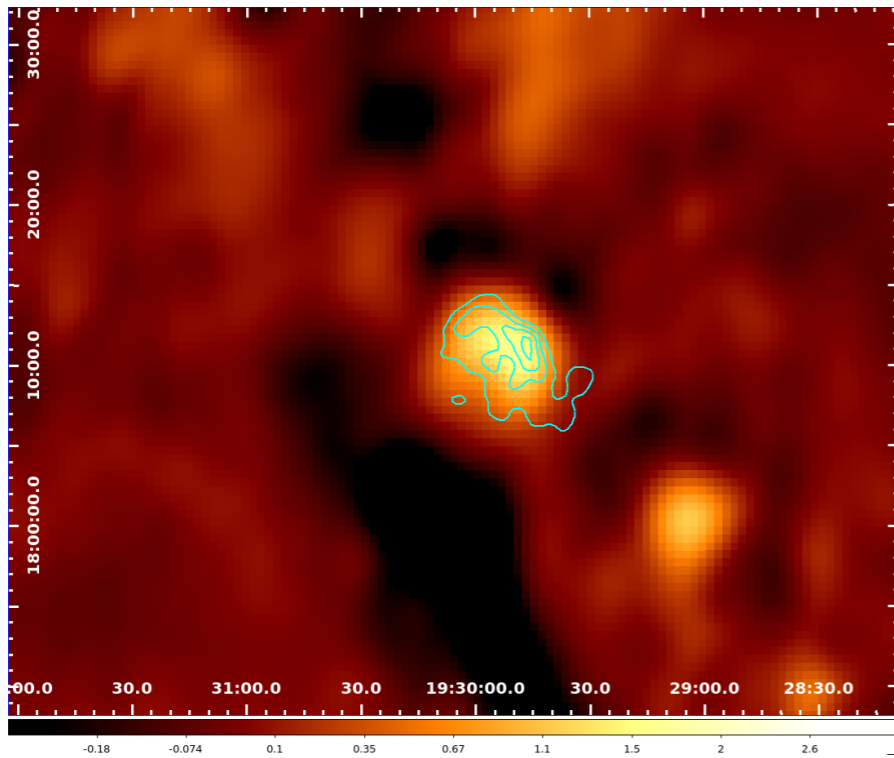
XMM-Newton EPIC-MOS2 0.3-10.0 keV

- ▶ Detected only by MOS2
- ▶ Spectrum ~2000 counts
- ▶ Several visible emission lines
- ▶ Best-fit model - NEI

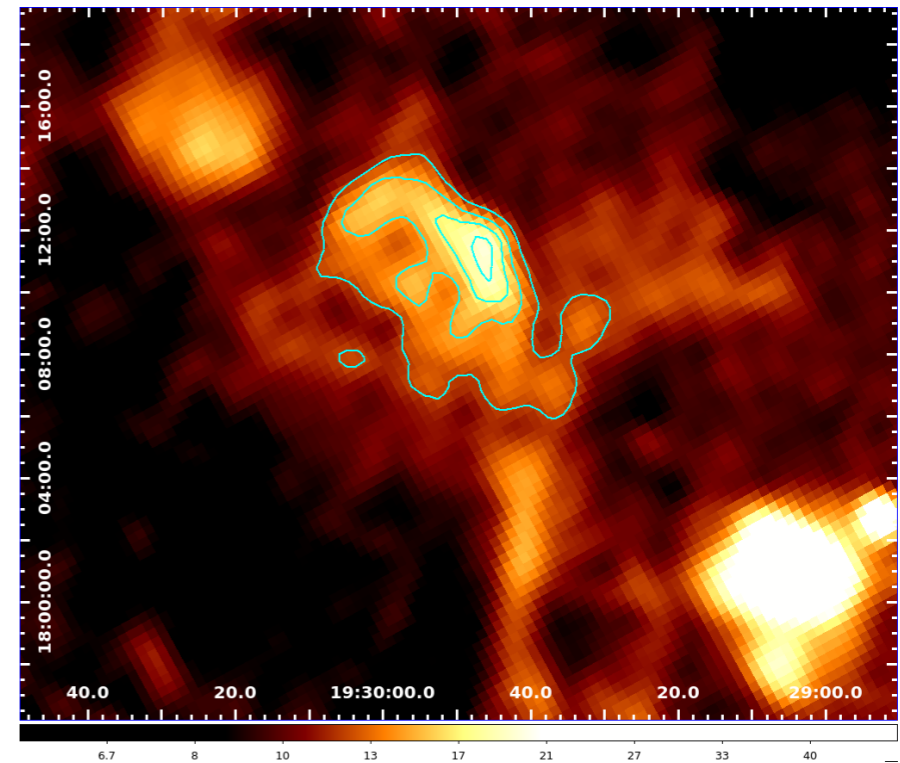
Parameter	Unit	Value	Element	Abundance
$N_H$	$10^{22} \text{ cm}^{-2}$	$2.4_{-0.2}^{+0.2}$	Ne	$0.2_{-0.2}^{+0.7}$
$n_e n_H V$	$10^{57} \text{ cm}^{-3}$	$5_{-2}^{+2}$	Mg	$0.9_{-0.2}^{+0.3}$
$T_2$	keV	$0.8_{-0.1}^{+0.2}$	Si	$0.5_{-0.1}^{+0.1}$
$\tau$	$10^{10} \text{ s cm}^{-2}$	$4_{-1}^{+2}$	S	$0.9_{-0.2}^{+0.2}$
			Fe	$1.3_{-0.5}^{+0.7}$

Cstat/d.o.f 83.48/64

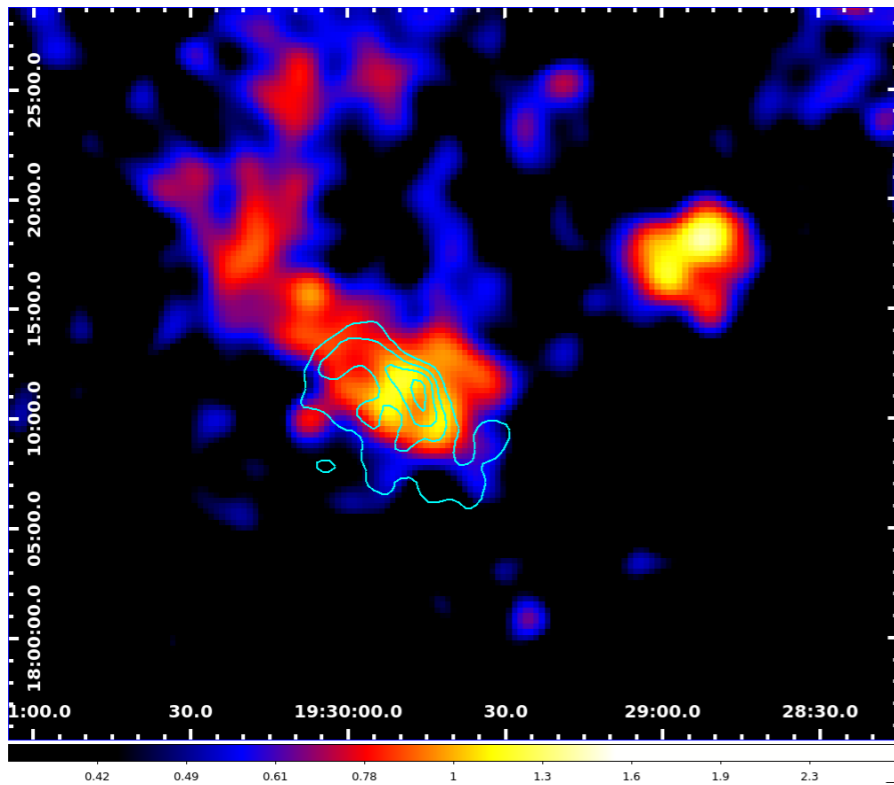




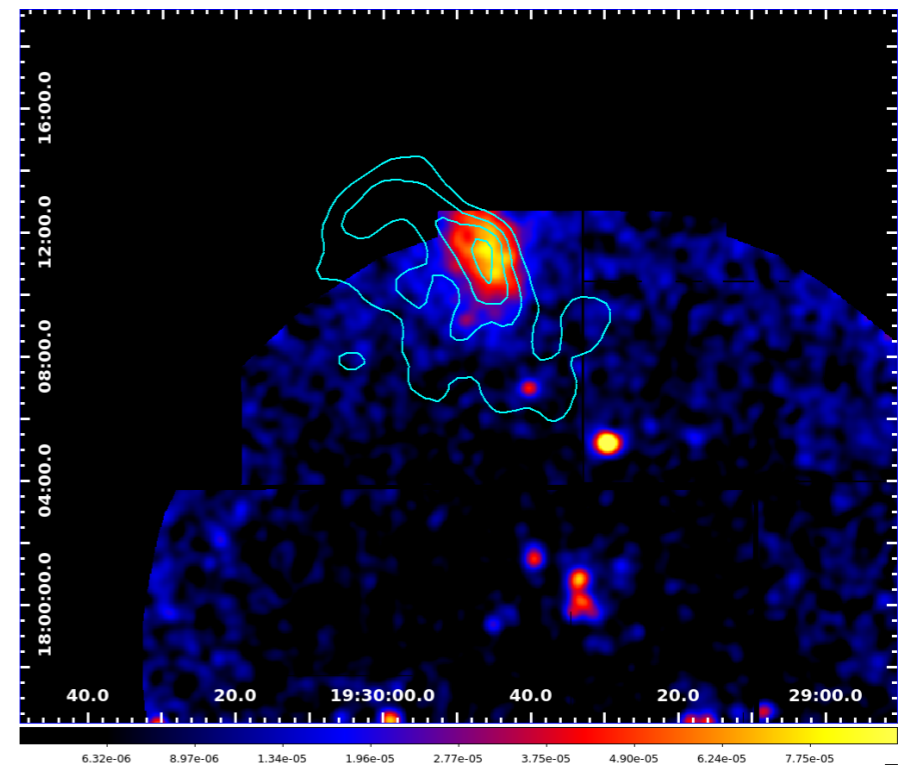
LOFAR 150 MHz



VLA 1.4 GHz



ROSAT 0.2-2.0 keV

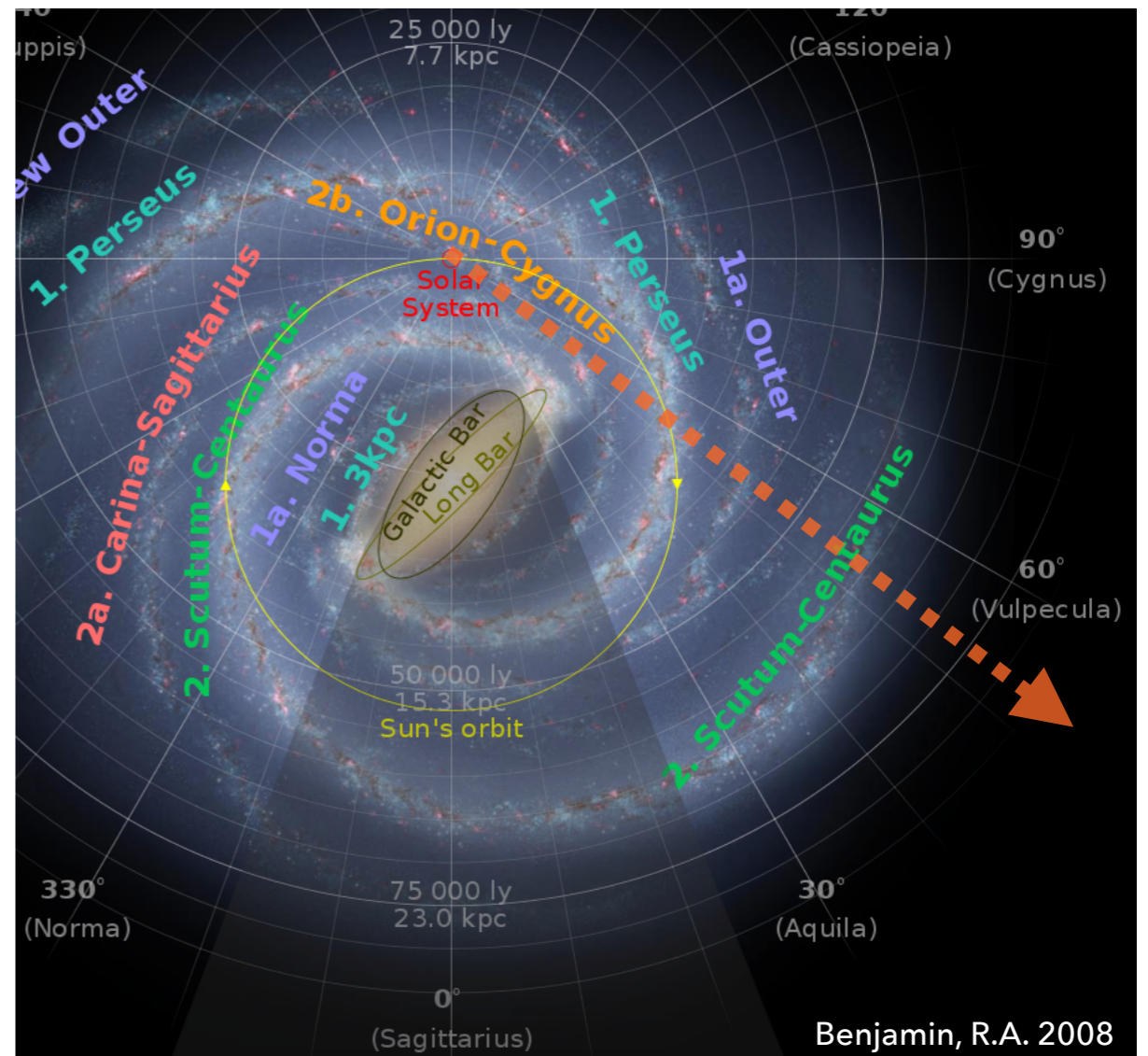
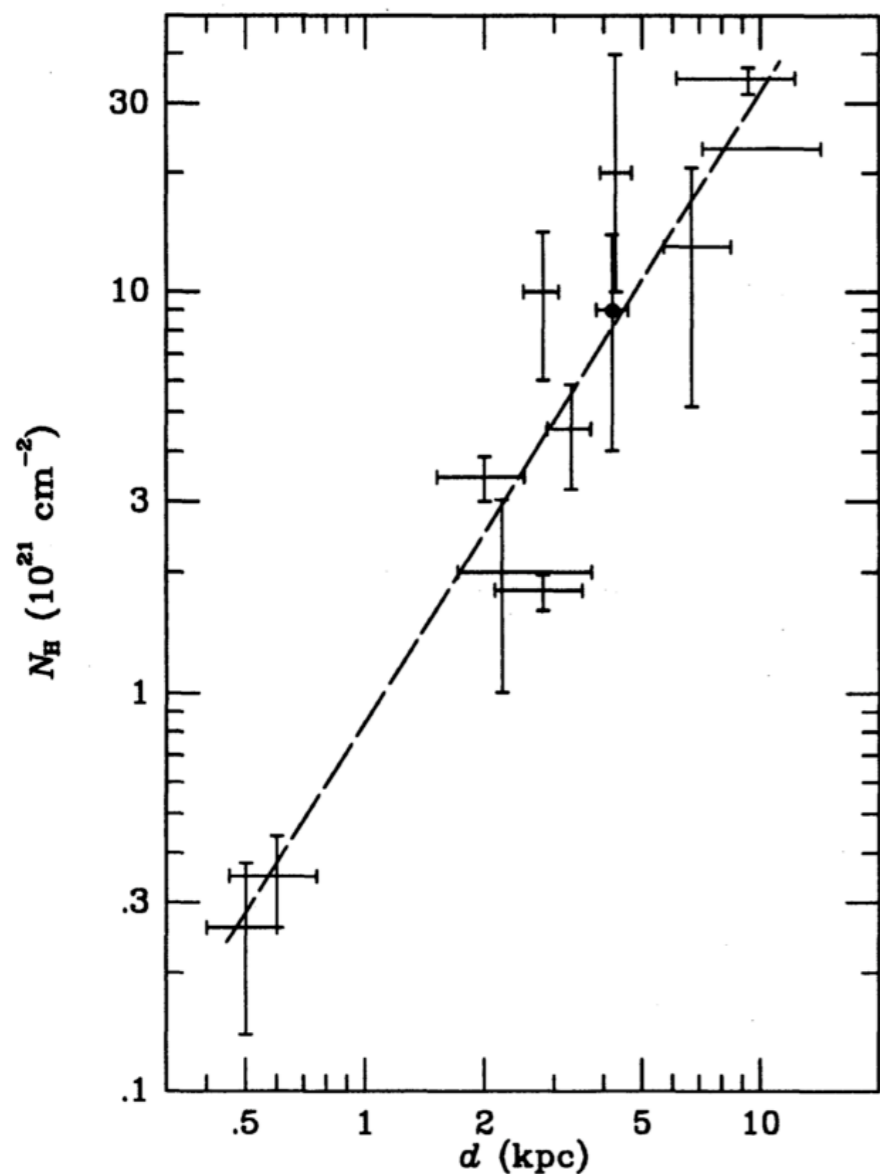


XMM-Newton EPIC-MOS2 0.3-10.0 keV

# DISTANCE OF G53.41+0.03

- ▶  $N_H$  - Distance relation (Strom 1994):  $D \sim 8$  kpc
- ▶ Galactic LoS  $\ell=53.4$  - Carina-Sagittarius Arm

$$N_H = d_{\text{kpc}}^{1.58} 8.4 \times 10^{20} \text{ cm}^{-2}$$



1) Ionisation age  $\tau = n_e t$ ;  $\Rightarrow t \approx 1600$  years

$$\tau = 4_{-1}^{+2} \cdot (10^{10} \text{ cm}^{-3} \text{ s}) \quad n_e = \sqrt{\frac{n_{\text{orm}}}{1.2V_x}} \approx 0.8 d_{7.5}^{-3/2} \text{ cm}^{-3}$$

$$t = \frac{\tau}{n_e} \approx 1600 d_{7.5}^{-3/2} \text{ years}$$

2) Sedov-Taylor self-similar evolution model

a) Shock Temperature, Radius  $\Rightarrow t \approx 5300$  years

$$v_s = \sqrt{\frac{kT}{1.2}} \cdot 1000 \text{ [km/s]} \approx 800 \text{ km/s}$$

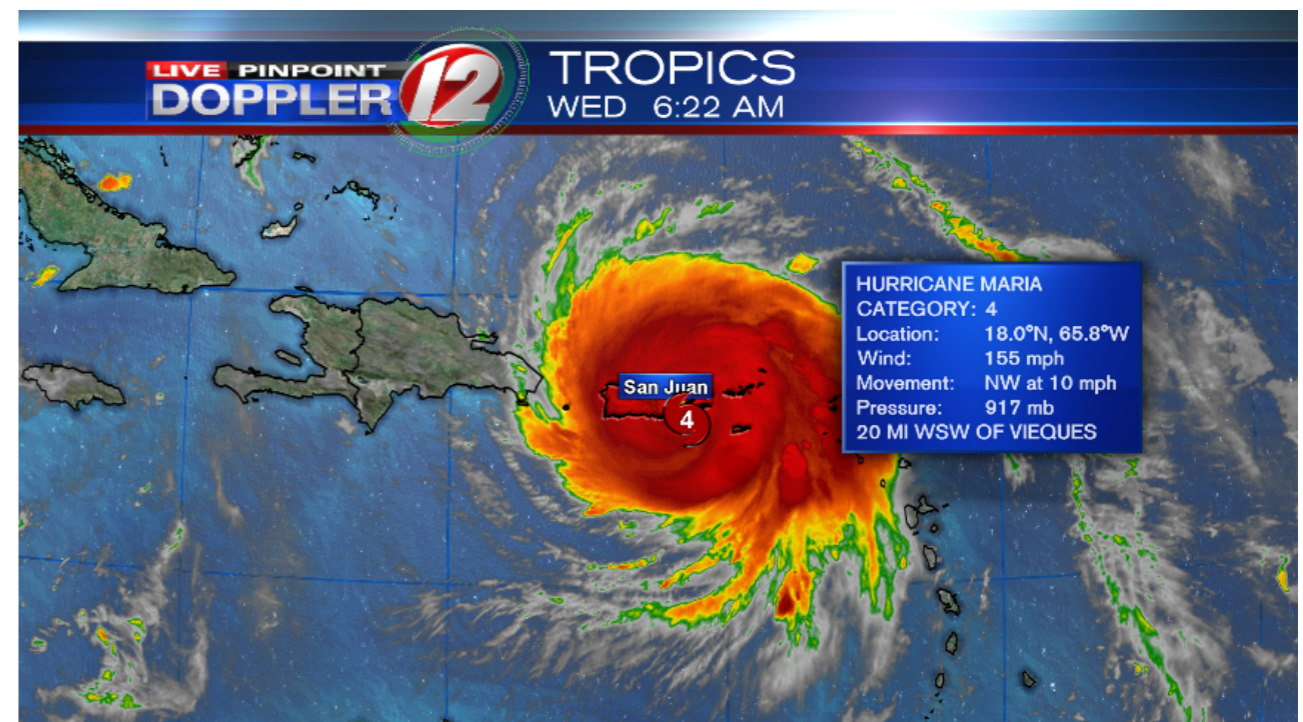
$$t = \frac{5}{2} \frac{v_s}{R_{SNR}} \approx 5300 d_{7.5}^{-1} \text{ years}$$

b) SN Energy, Radius  $\Rightarrow t \approx 7800$  years

$$R^5 = 2.026 \frac{Et^2}{n_H}$$

$$t = 0.7026 \left( \frac{n_H R^5}{E} \right)^{1/2} \approx 7800 d_{7.5}^{5/2} \text{ years}$$

- ▶ Arecibo analysis delayed by weather
- ▶ Non-detection with upper limits
  1. On-centred pulsar:  $S_{\max} = 0.011$  mJy
  2. Off-centred pulsar:  $S_{\max} = 0.045$  mJy
- ▶ X-ray compact source still a possibility



## 1. Case study - G53.41+0.03

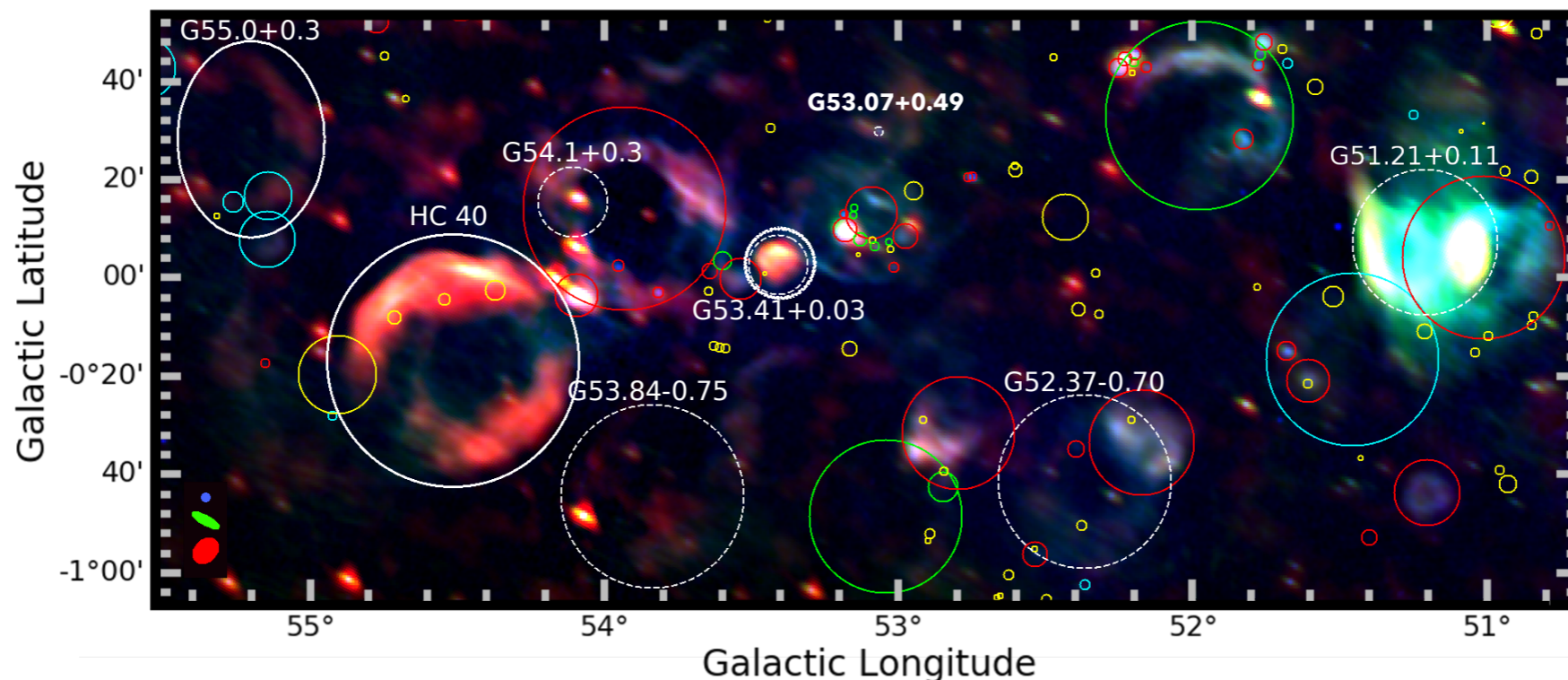
- ▶ Radio
- ▶ X-rays (Chandra, XMM-Newton)

## 2. Surveys detections

- ▶ Radio - LOFAR
- ▶ X-rays - eRosita



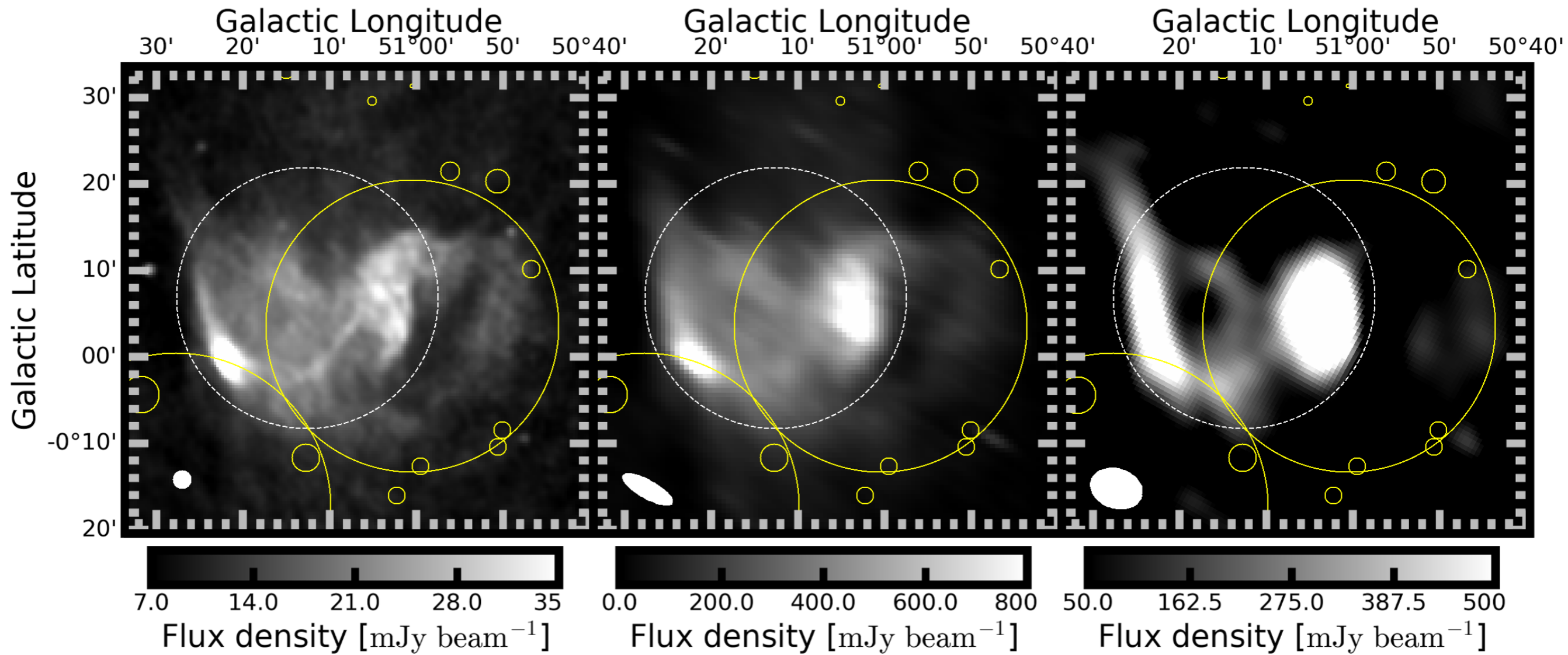
- ▶ We confirm that G53.41+0.03 is a supernova remnant
  - ▶ Deeper centred-on observation needed
- ▶ Other SNR candidates in FoV
  - ▶ Three candidates unlikely
  - ▶ G51.21+0.11 requires further investigation

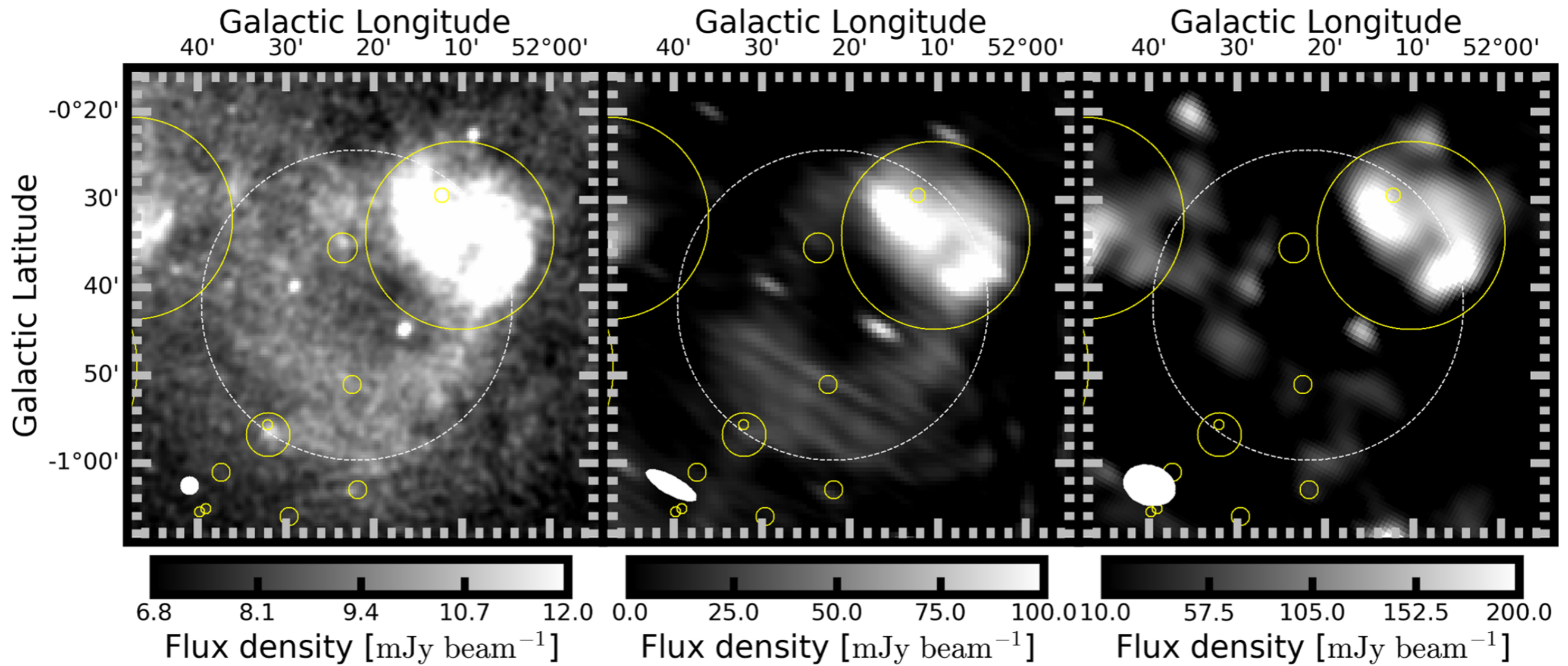


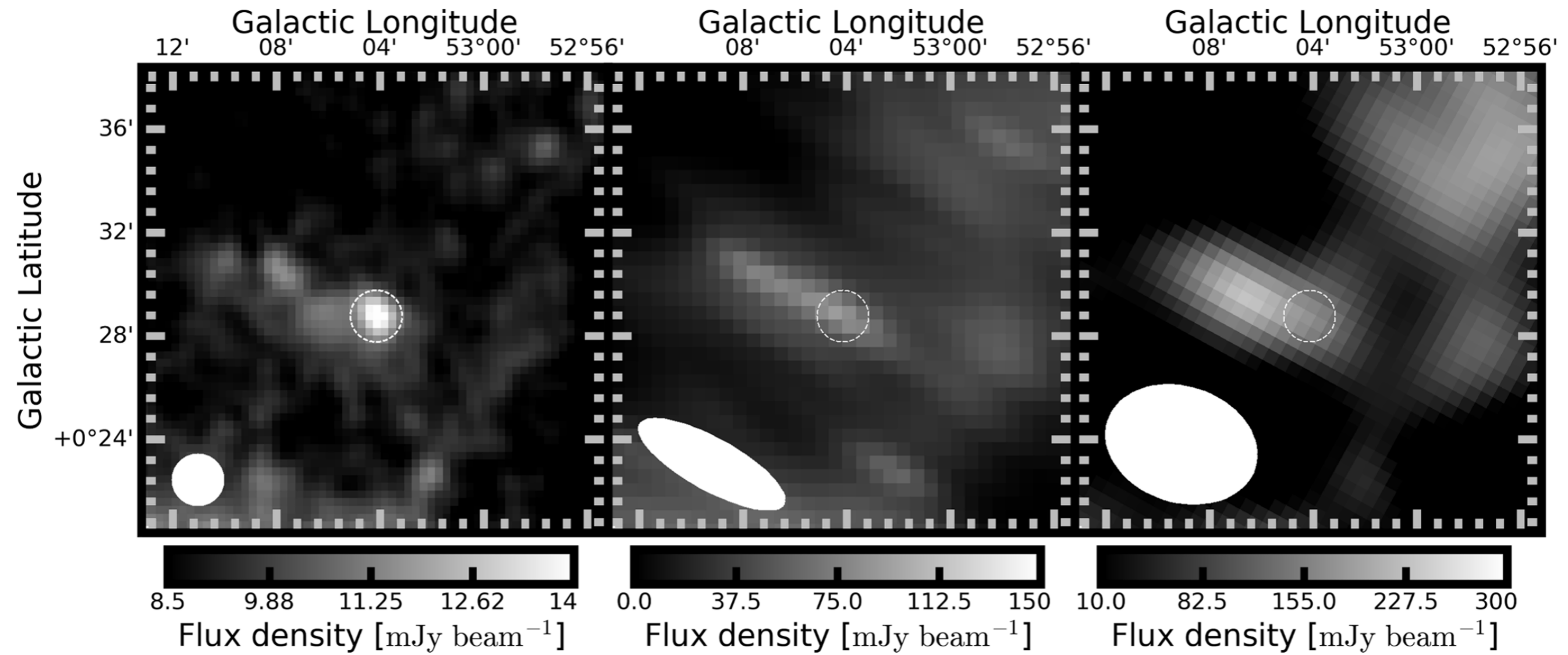


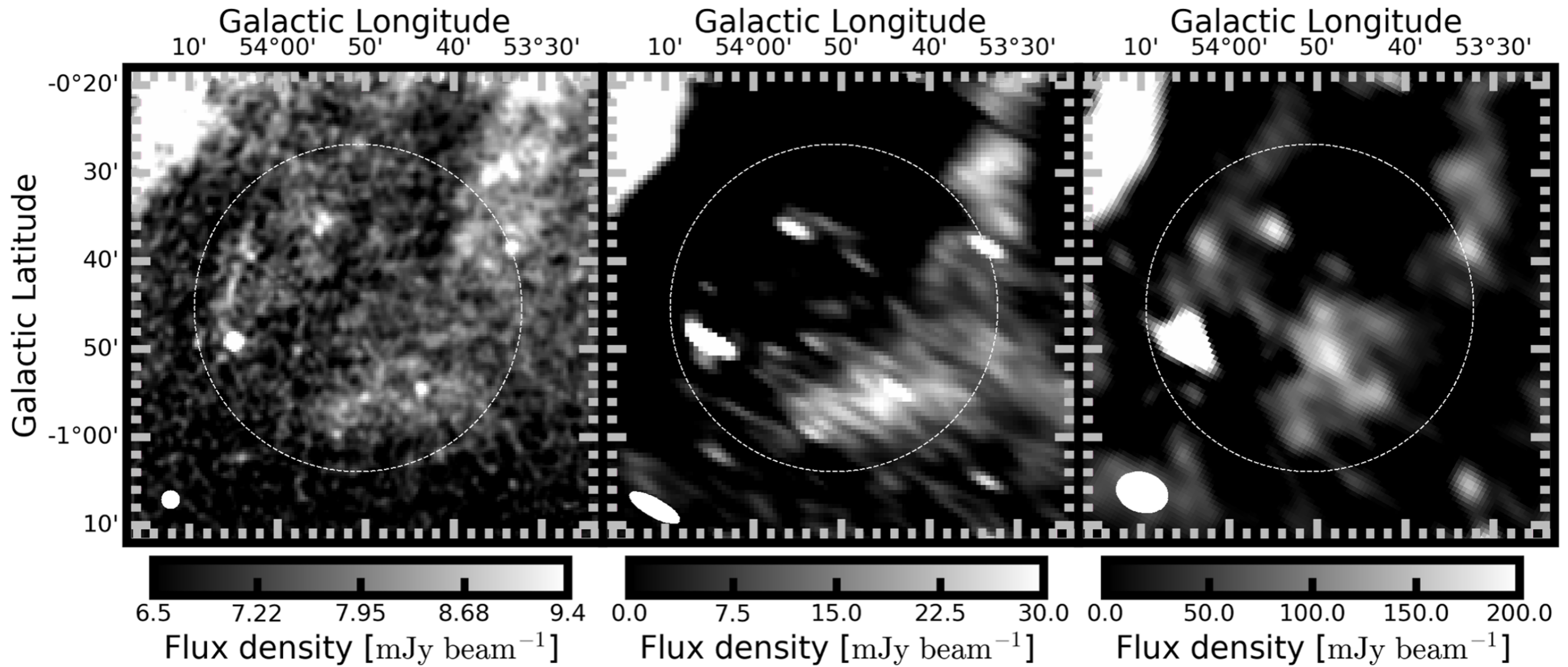
# BACKUP SLIDES

# G51.21+0.11









# G54.1+0.3

