



March 27



Bamberg



Irina
Smirnova-Pinchukova
MPIA, Heidelberg



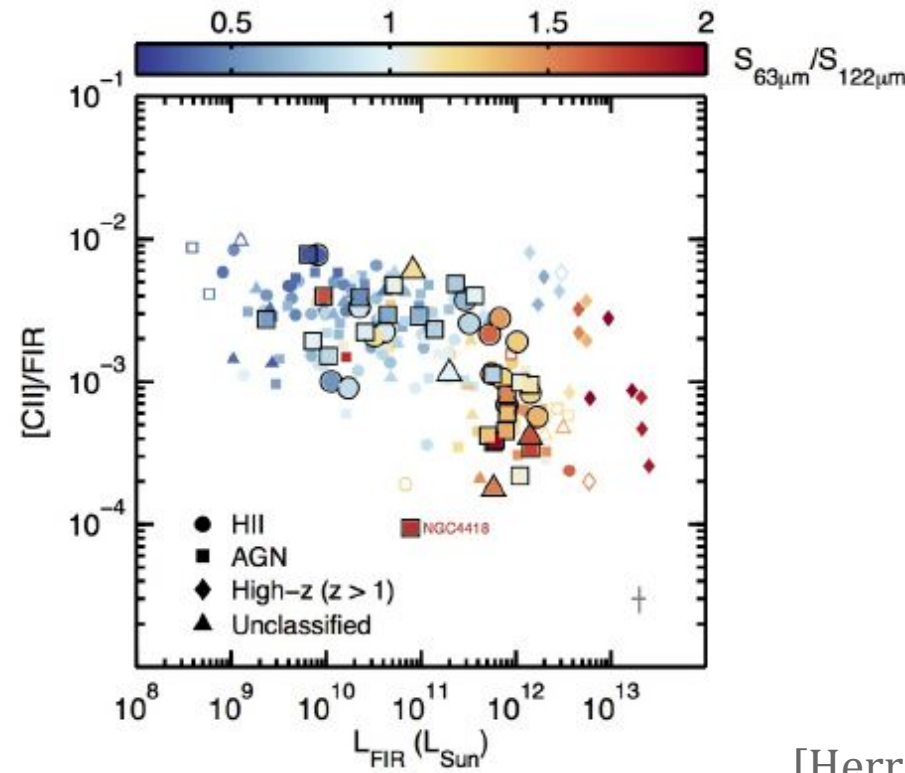
Bernd
Husemann



Gerold
Busch

Identifying the
AGN contribution
to [CII] ionization
in luminous local AGN

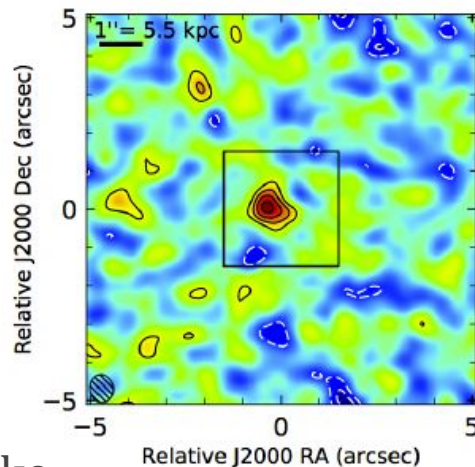
[CII] line as a SFR tracer



[Herrera-Camus+ 2018]

[CII] line as a SFR tracer

Star Formation?



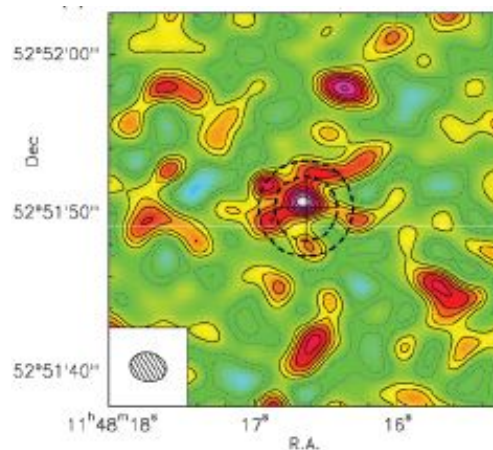
Himiko

ALMA

$z \sim 6.6$

[Carniani+ 2018]

AGN outflows?



SDSS J114816.64+525150.3

NOEMA

$z \sim 6.4$

[Maiolino+ 2012]

| INTRODUCTION | OBSERVATIONS | METHOD | RESULTS |
|--------------|--------------|--------|---------|
|--------------|--------------|--------|---------|

What is the origin of [CII] ionization?

- ❑ luminous AGN
- ❑ low-redshift galaxies
- ❑ large enough

Herschel

Space Observatory

14 May 2009 - 17 June 2013



SOFIA

Flying Observatory

26 May 2010 - now



What is the origin of [CII] ionization?

Close AGN Reference Survey



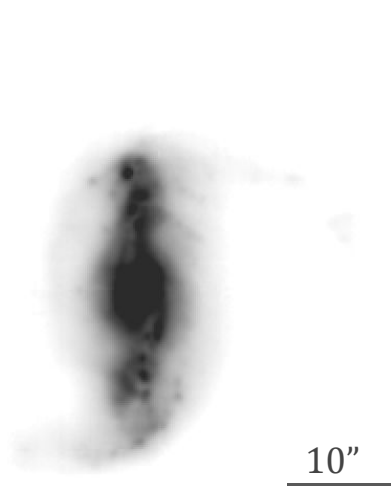
- 40 unobscured AGN
- local $0.01 < z < 0.06$
- multiwavelength observations

Subsample:

| | |
|--------------|--------------|
| HE 0108-4743 | HE 1029-1831 |
| HE 0232-0900 | HE 1108-2813 |
| HE 0433-1028 | HE 1353-1917 |
| HE 1017-0305 | HE 2211-3903 |

| introduction | observations | method | results |
|--------------|--------------|--------|---------|
|--------------|--------------|--------|---------|

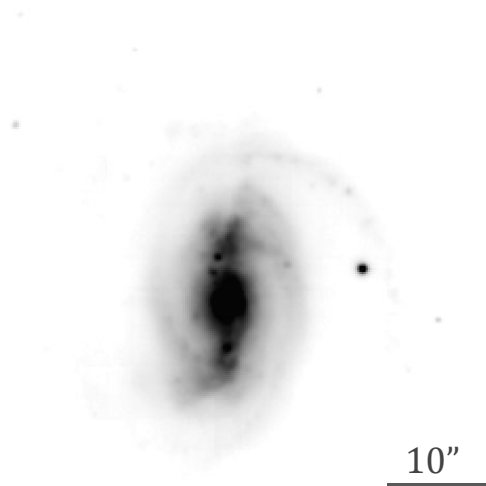
HE 0433-1028



☐ $z = 0.036$

☐ 149 Mpc

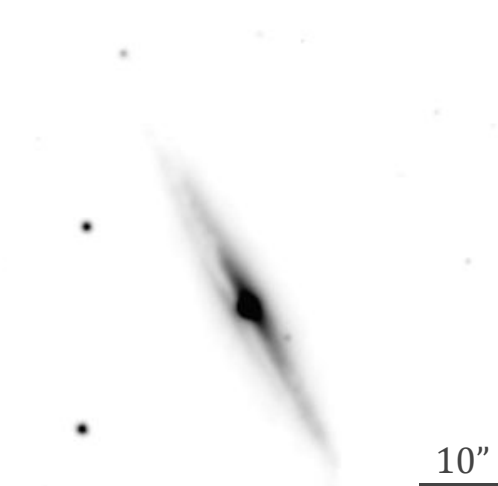
HE 1108-2813



☐ $z = 0.024$

☐ 105 Mpc

HE 1353-1917

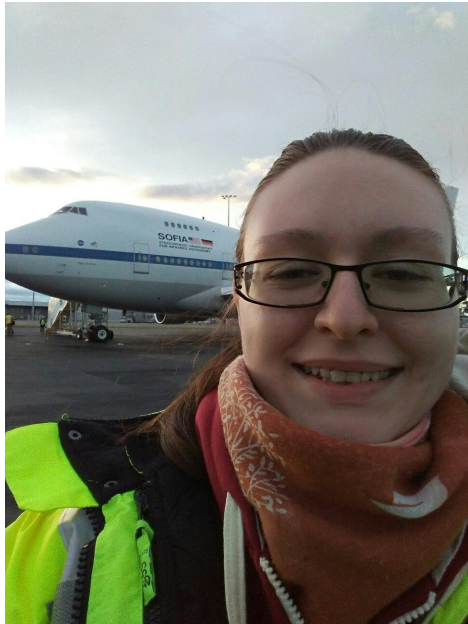


☐ $z = 0.035$

☐ 152 Mpc

| | | | |
|--------------|---------------------|--------|---------|
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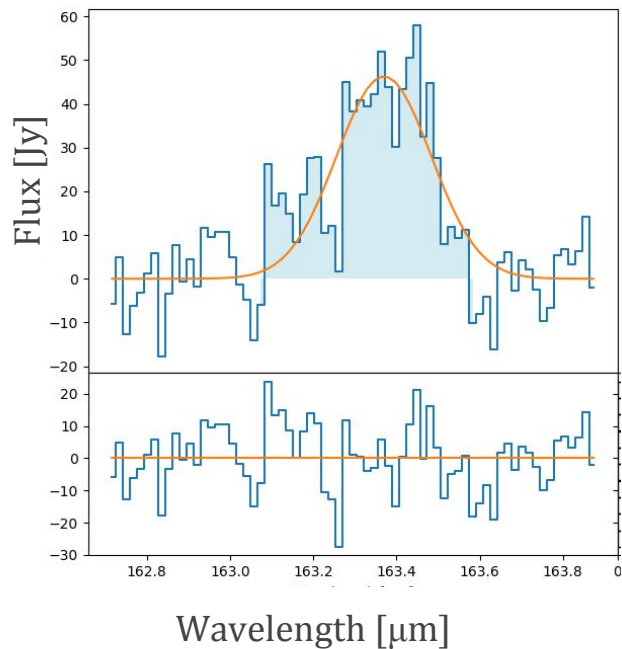
SOFIA flights, July 2017, Christchurch, New Zealand



Alfred Krabbe
FIFI-LS PI

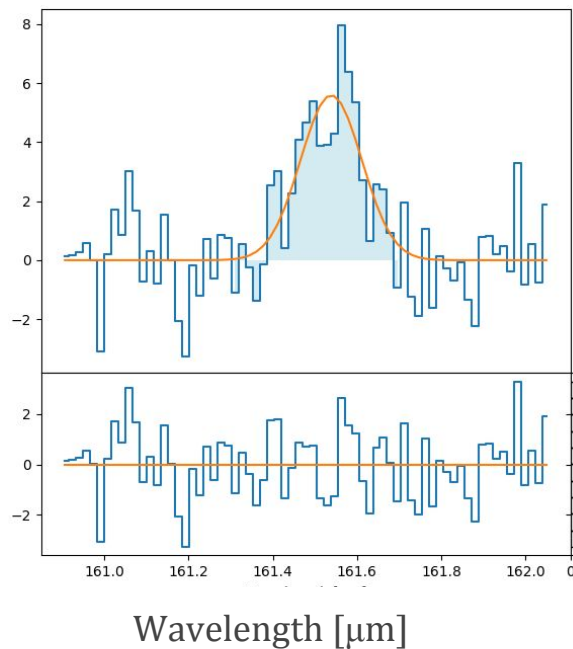
HE 0433-1028

50 min



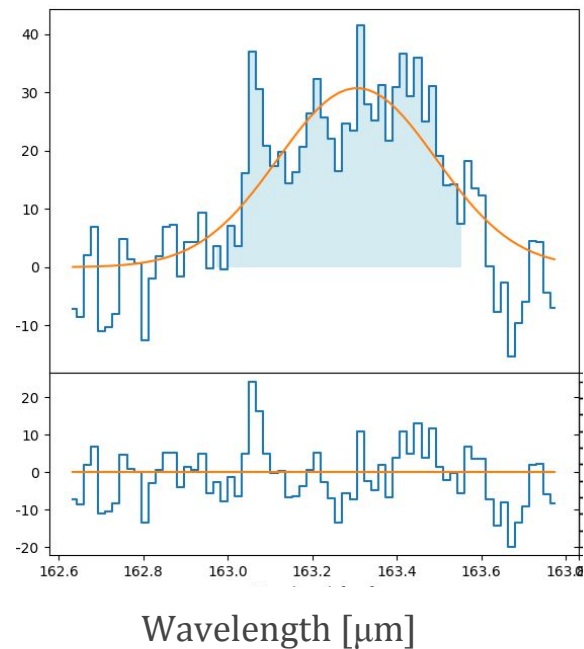
HE 1108-2813

68 min



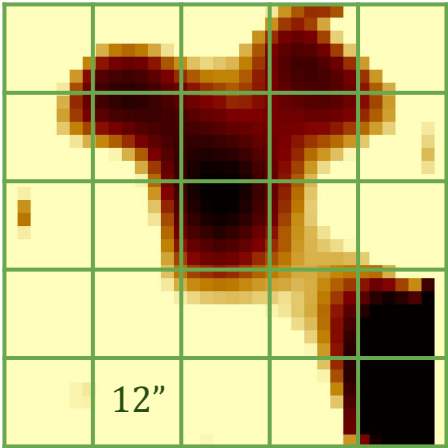
HE 1353-1917

72 min

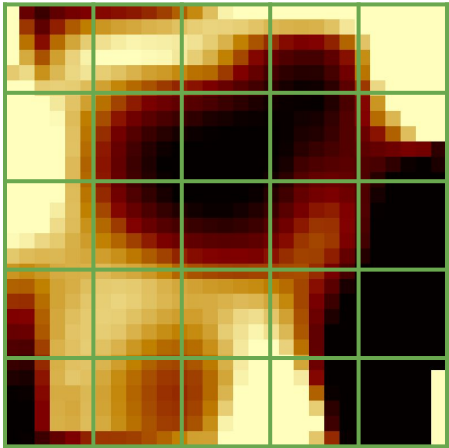


Standart pipeline output

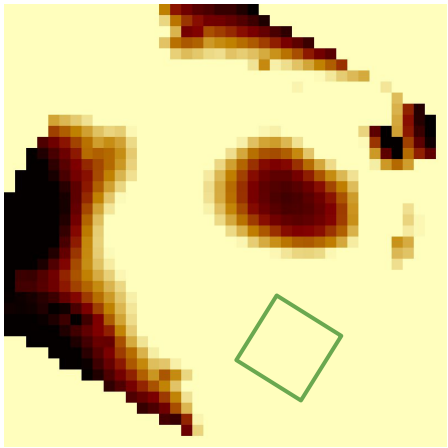
HE 0433-1028



HE 1108-2813

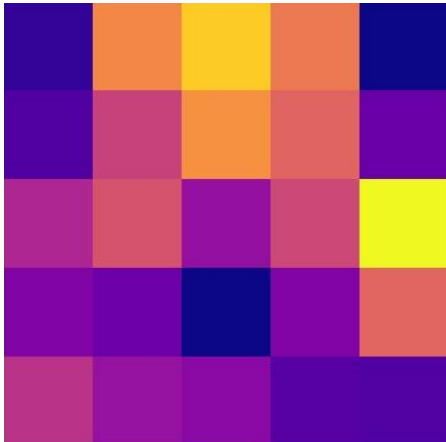


HE 1353-1917

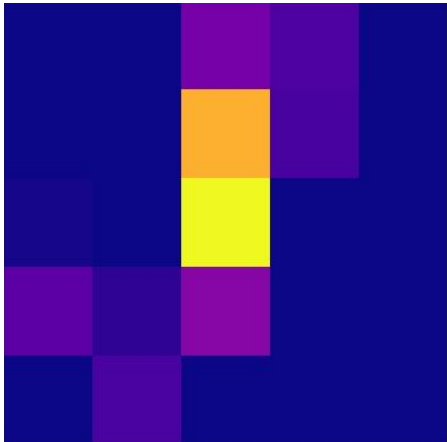


Reconstructed [CII] maps

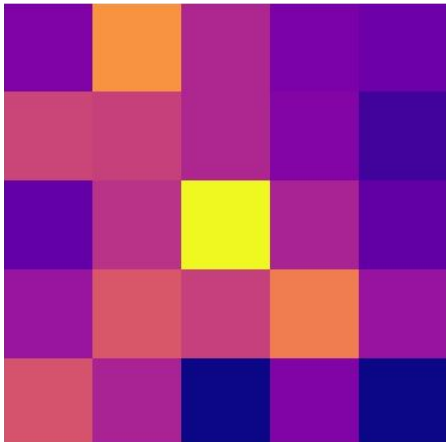
HE 0433-1028



HE 1108-2813



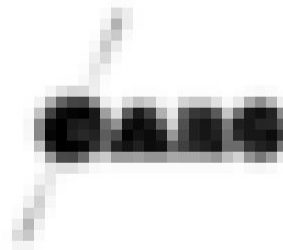
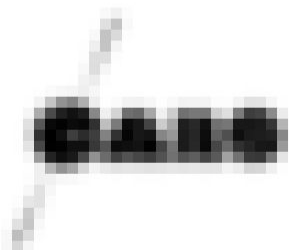
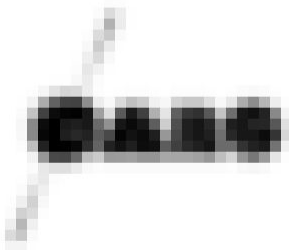
HE 1353-1917



Pixel size = 12"

| | | | |
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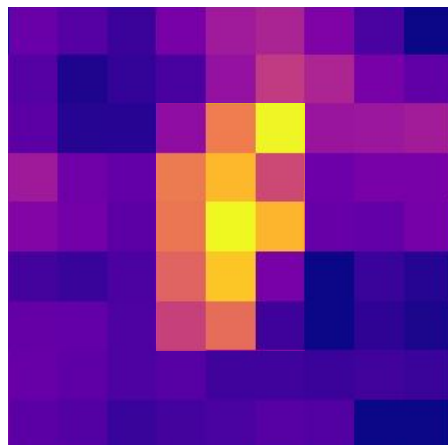
Drizzle algorithm



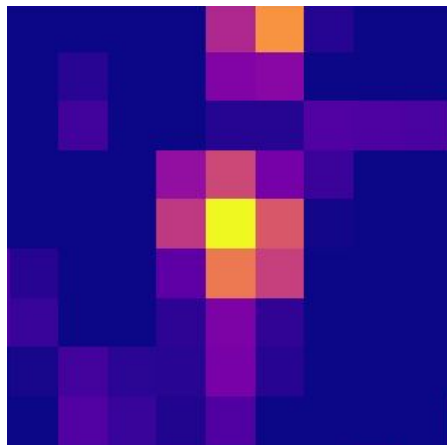
| | | | |
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Drizzled [CII] maps

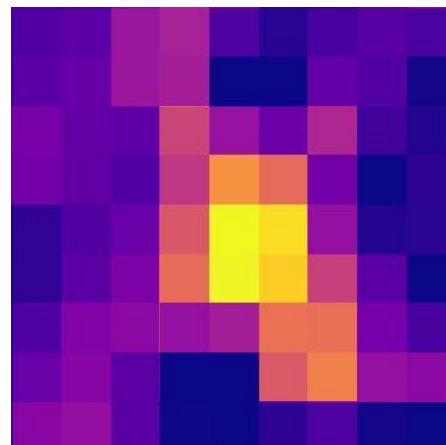
HE 0433-1028



HE 1108-2813



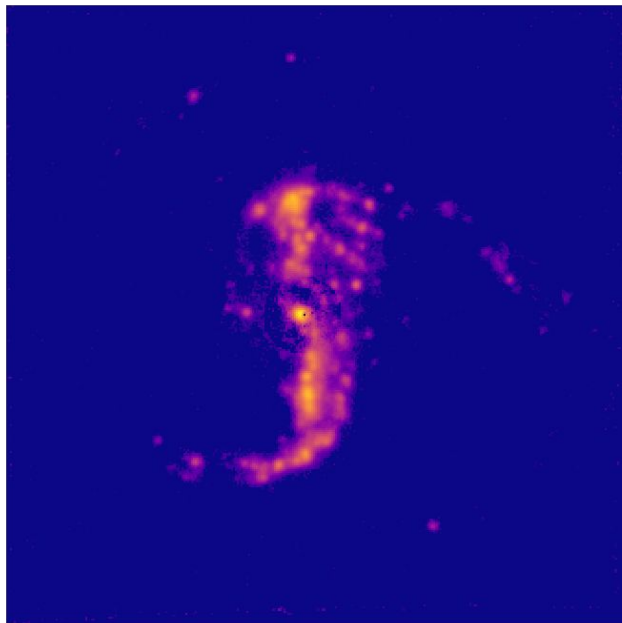
HE 1353-1917



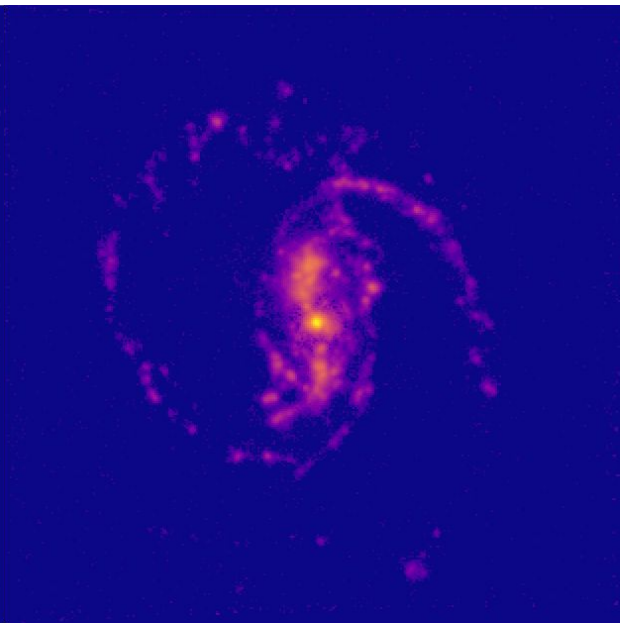
Pixel size = 6"

Comparison with H α MUSE/VLT

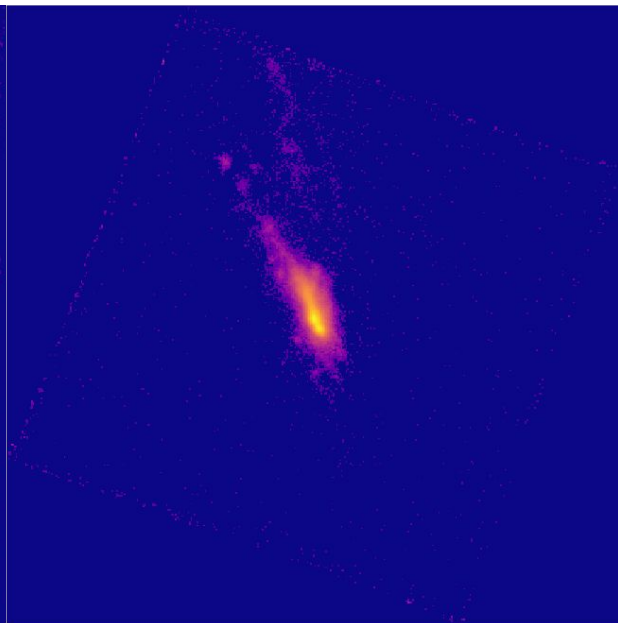
HE 0433-1028



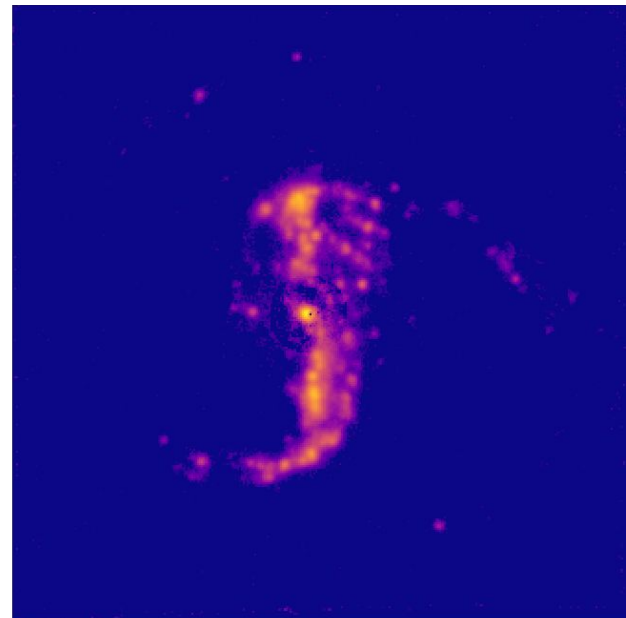
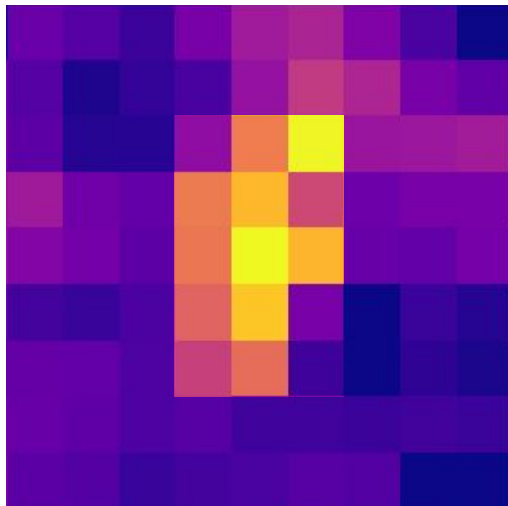
HE 1108-2813



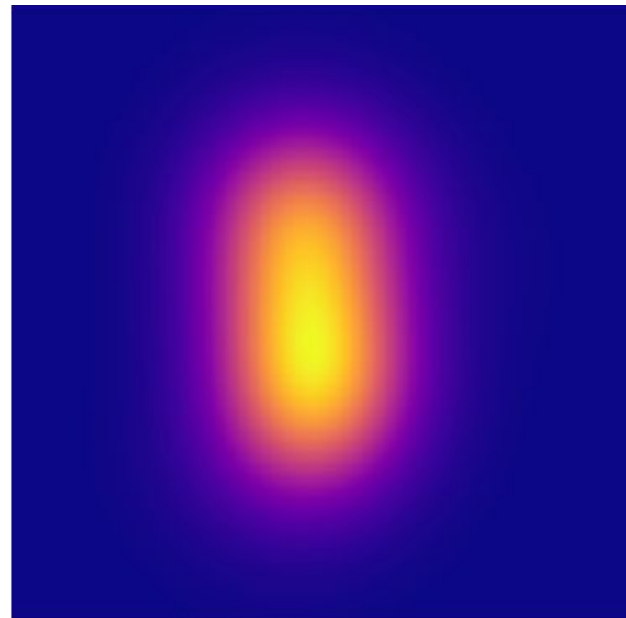
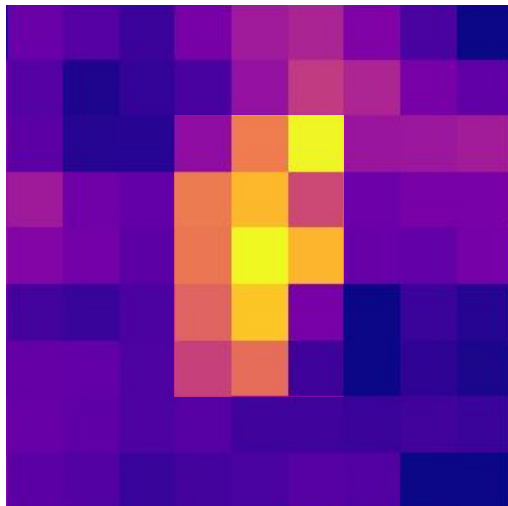
HE 1353-1917



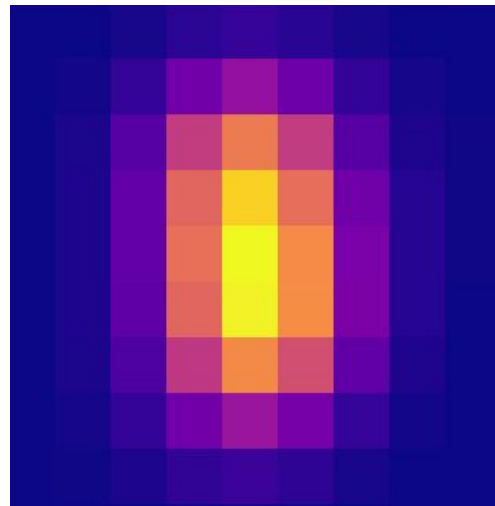
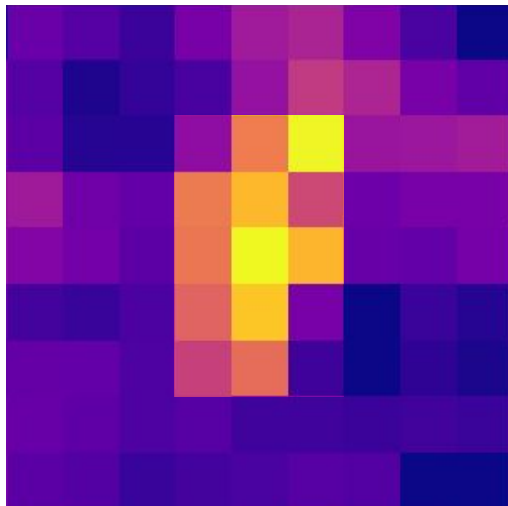
Comparison with H α MUSE/VLT



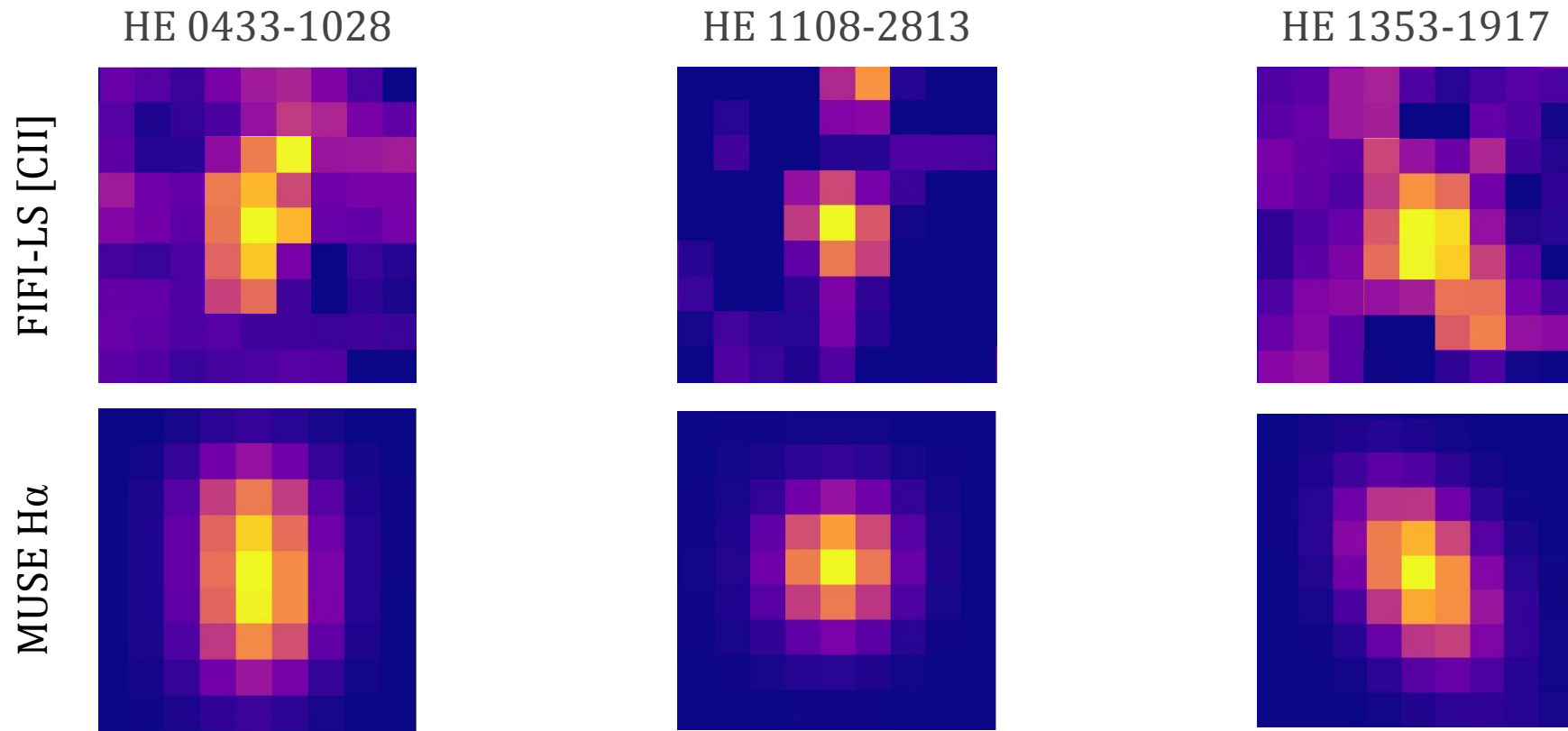
Comparison with H α MUSE/VLT



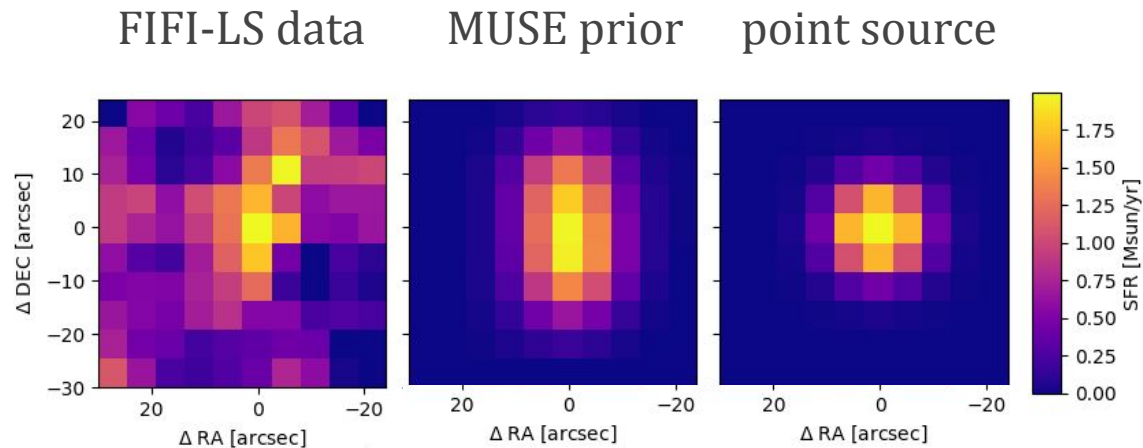
Comparison with H α MUSE/VLT



Comparison with H α MUSE/VLT

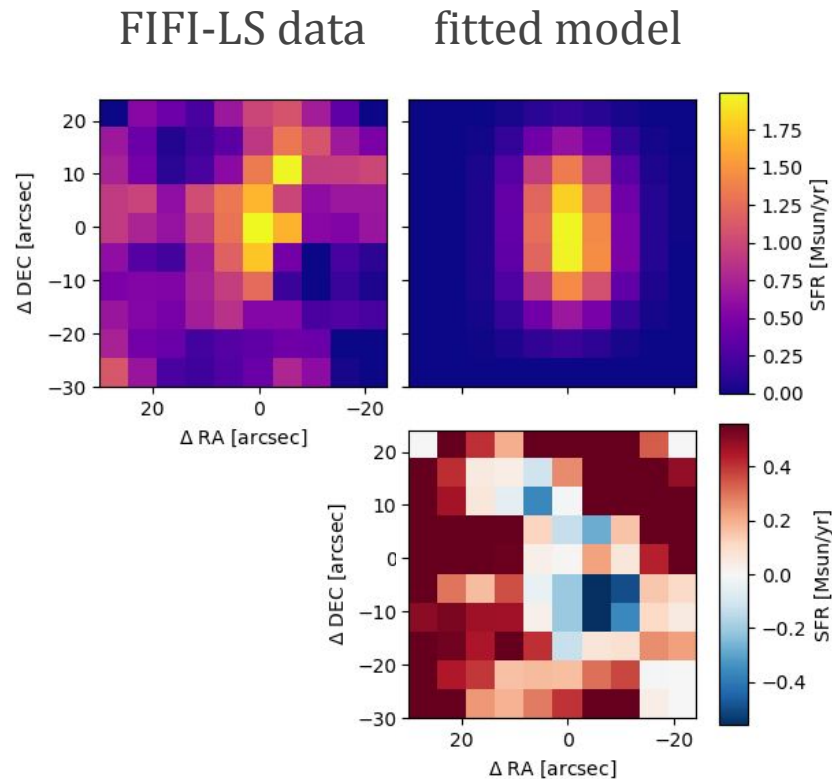
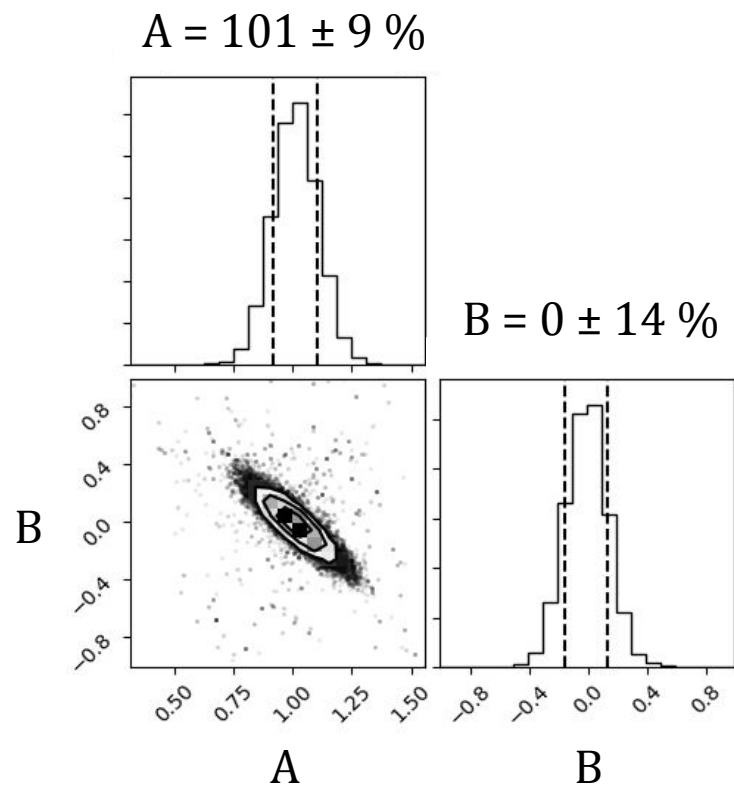


Preliminary results HE 0433-1028

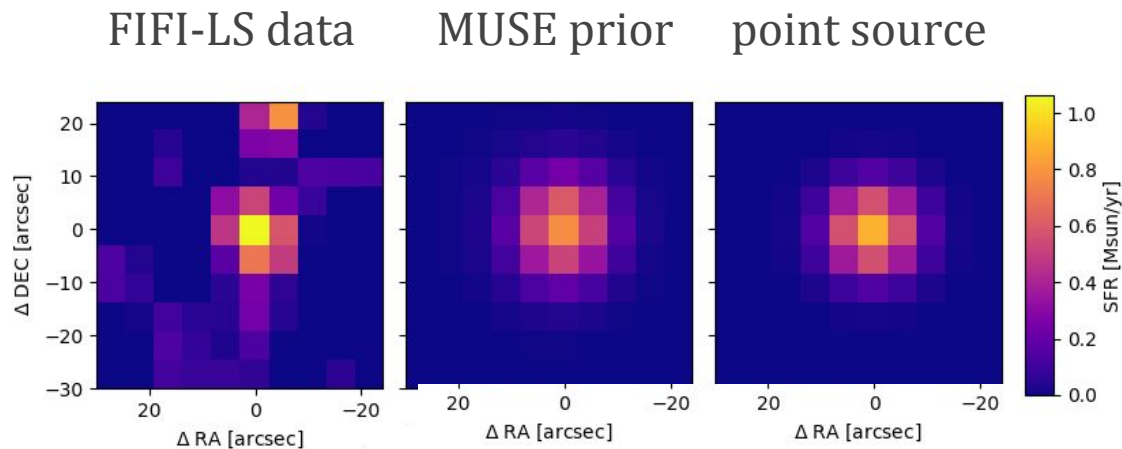


$$\text{Model} = A \times \text{MUSE prior} + B \times \text{point source}$$

Preliminary results HE 0433-1028

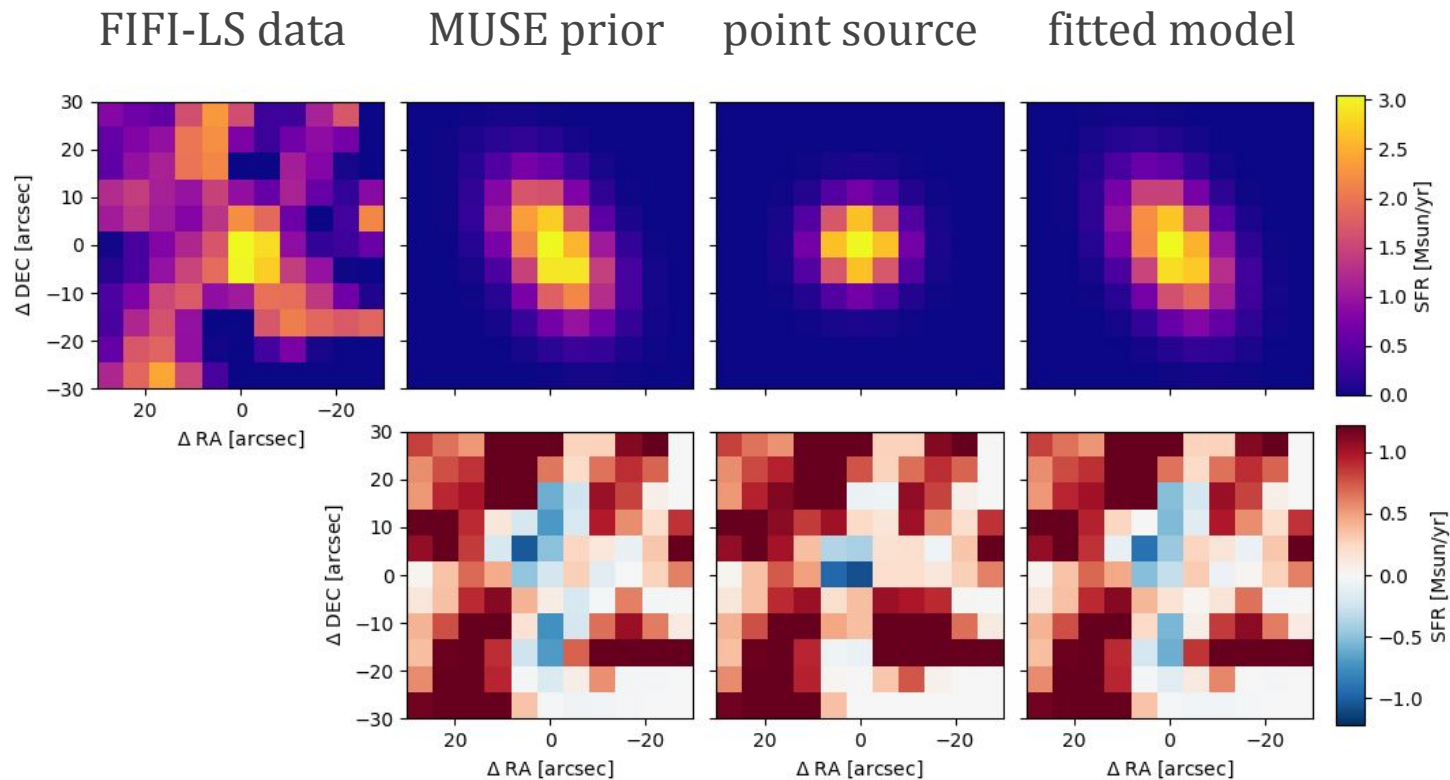


Preliminary results HE 1108-2813

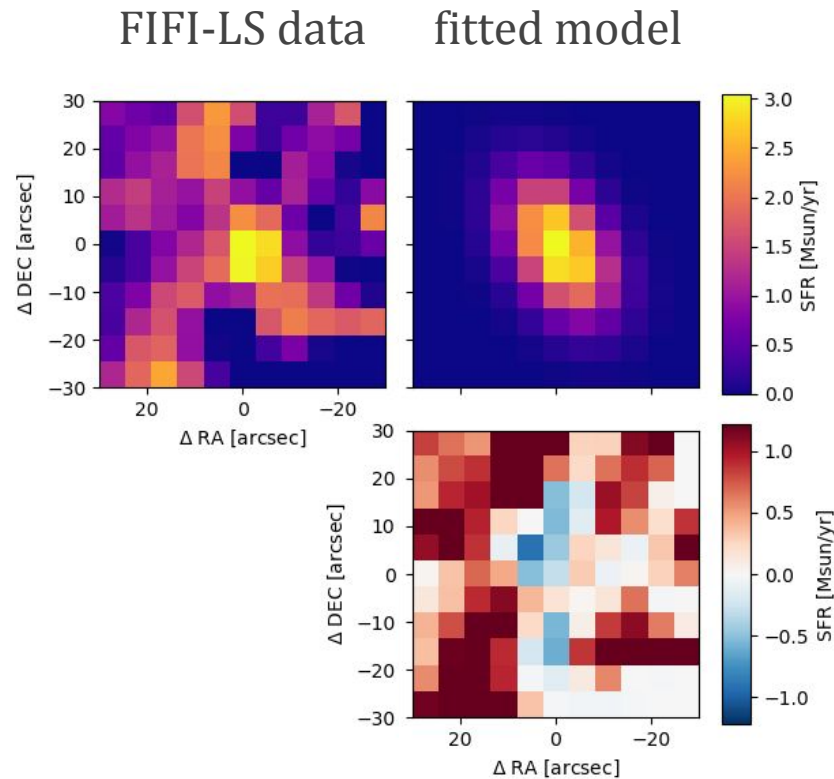
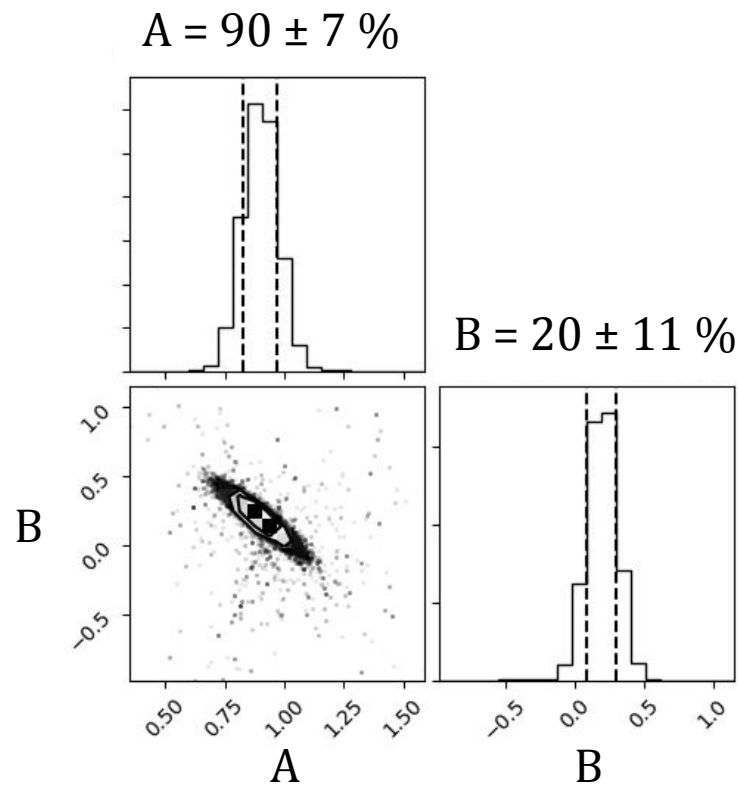


$$\text{Model} = A \times \text{MUSE prior} + B \times \text{point source}$$

Preliminary results HE 1353-1917



Preliminary results HE 1353-1917



Take away messages

- ❑ SOFIA resolves galaxies at $z \sim 0.03$
- ❑ It is possible to distinguish AGN contribution to [CII] ionization
- ❑ L [CII] - SFR relation as a function of AGN luminosity

FUTURE GOALS

- ❑ [Busch et al.] in prep
- ❑ Increase the sample
- ❑ Herschel SHINING [Herrera-Camus+ 2018]

THANK YOU!